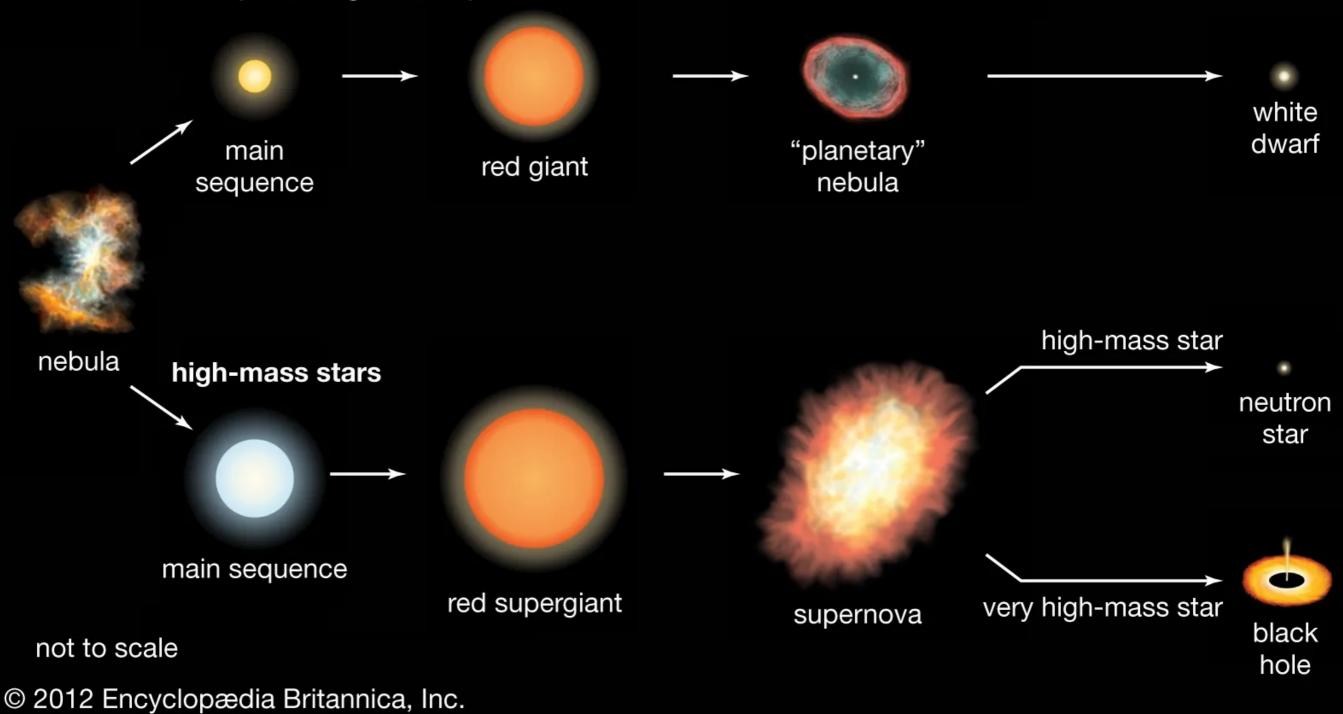
# Characterizing Exoplanetary Rocks Using Polluted White Dwarfs

#### Isabella Trierweiler UCLA

#### **Stellar evolution**

low- and medium-mass stars (including the Sun)



 $\sim$ I/3 of white dwarfs are polluted by heavy elements

## WHITE DWARF SYSTEMS ARE ACTIVE



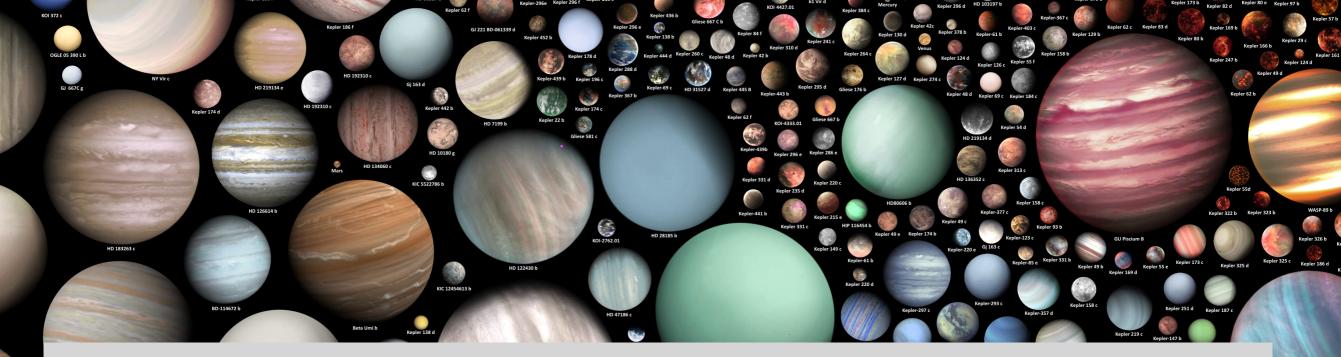
Gas, dust, major and minor planets have been detected around white dwarfs in a variety of orbits and states of decay

#### Planetary metals observed in white dwarf photospheres

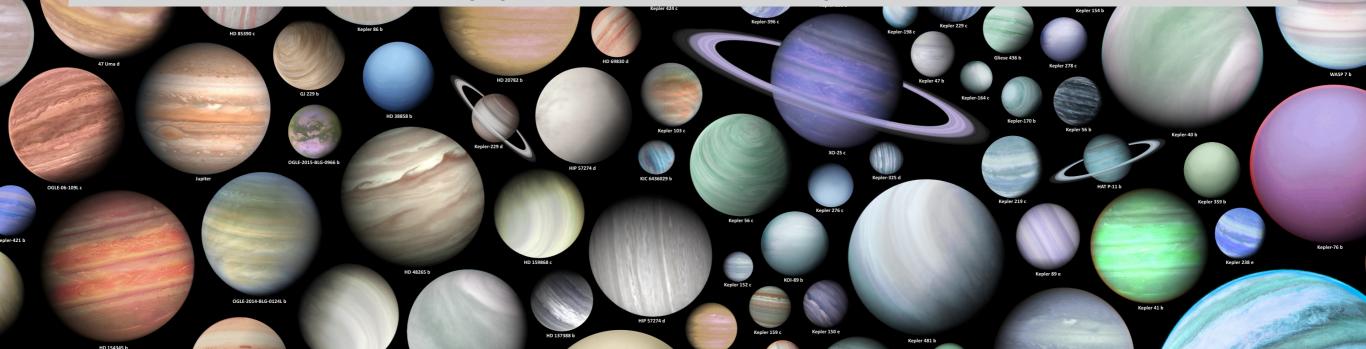
				Lith	oph	iles		Chalcophiles									
Li	Be			Side	ropł	niles		Volatiles						С	Ν	0	
Na	Mg													Si	Ρ	S	
К	Ca	Sc		Ti	V	Cr	Mn	Fe	Co	Ni	Cu						
	Sr																

Veras 2021

- Pollution is mostly rocky and sometimes icy
- >1000 polluted white dwarfs, ~50 highly polluted
- Pollution candidates include asteroids, comets, KBOs, exomoons

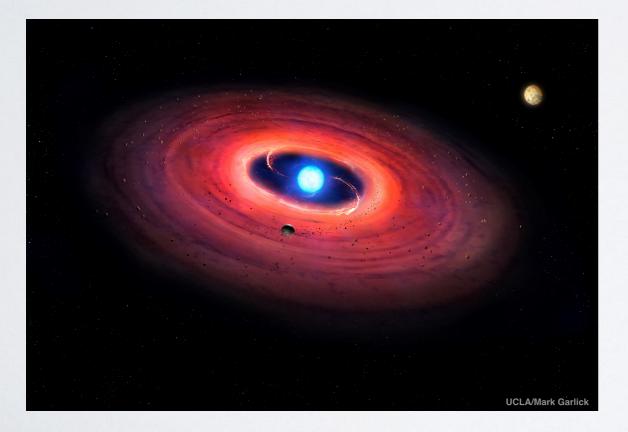


# How do exoplanet compositions vary across the galaxy? Are the rocky planets in our Solar System typical or unique?



# HOW DO EXOPLANET COMPOSITIONS VARY ACROSS THE GALAXY?

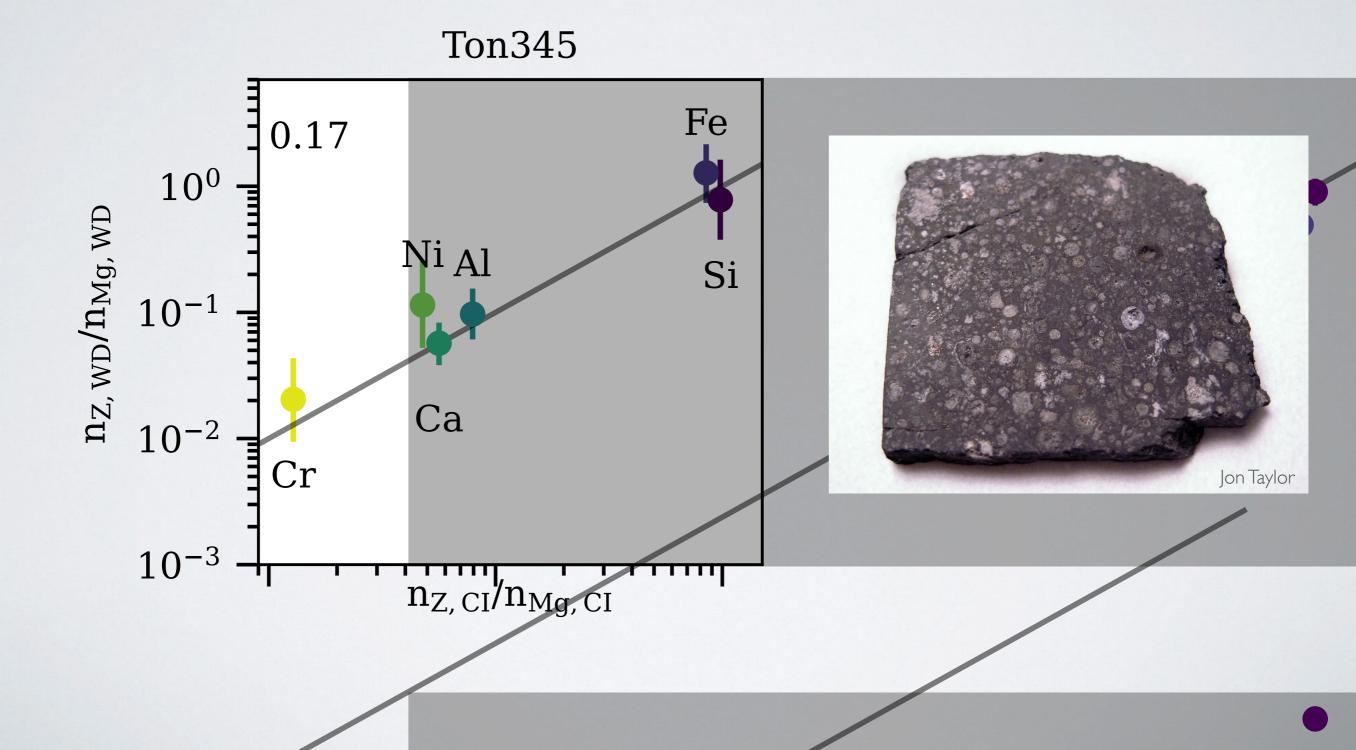
- White dwarf pollution provides samples of exoplanetary rocks in the solar neighborhood
- We test local exoplanet compositions by comparing polluted white dwarfs to solar system rocks and benchmark against local stellar abundances



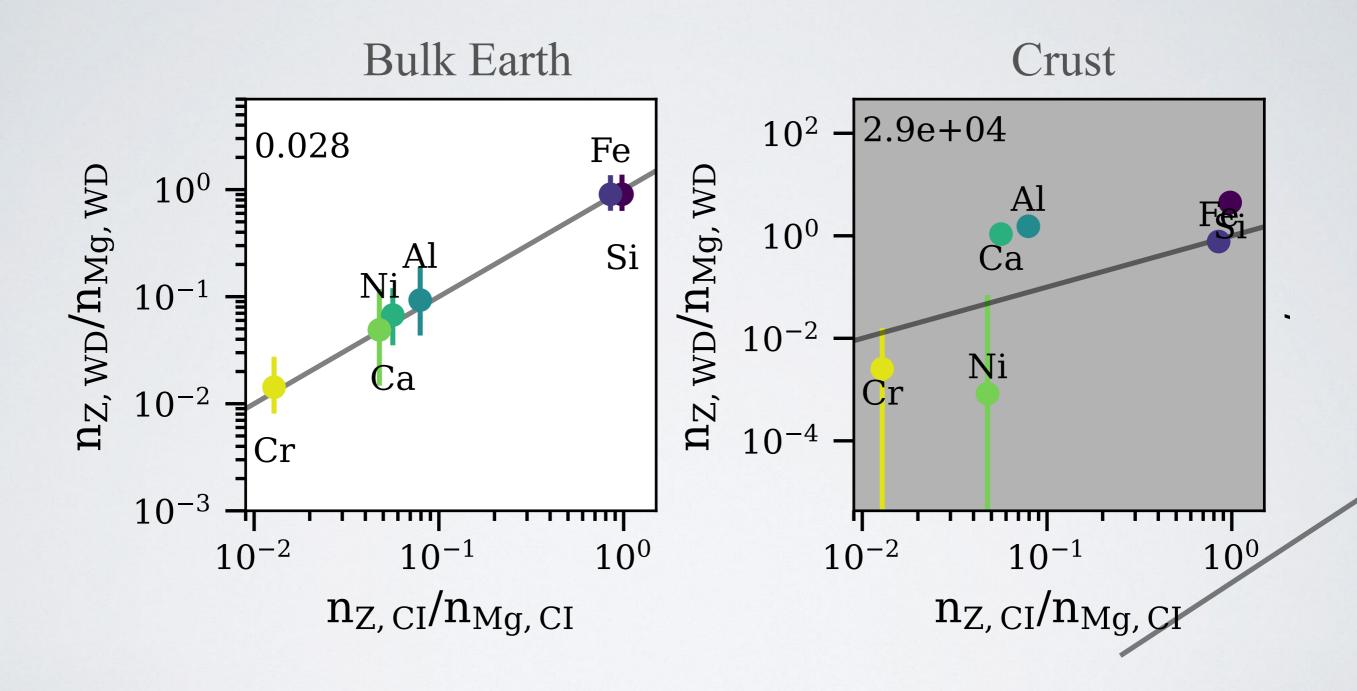


# **COMPARING EXO-ROCKS TO SOLAR SYSTEM CHONDRITES**

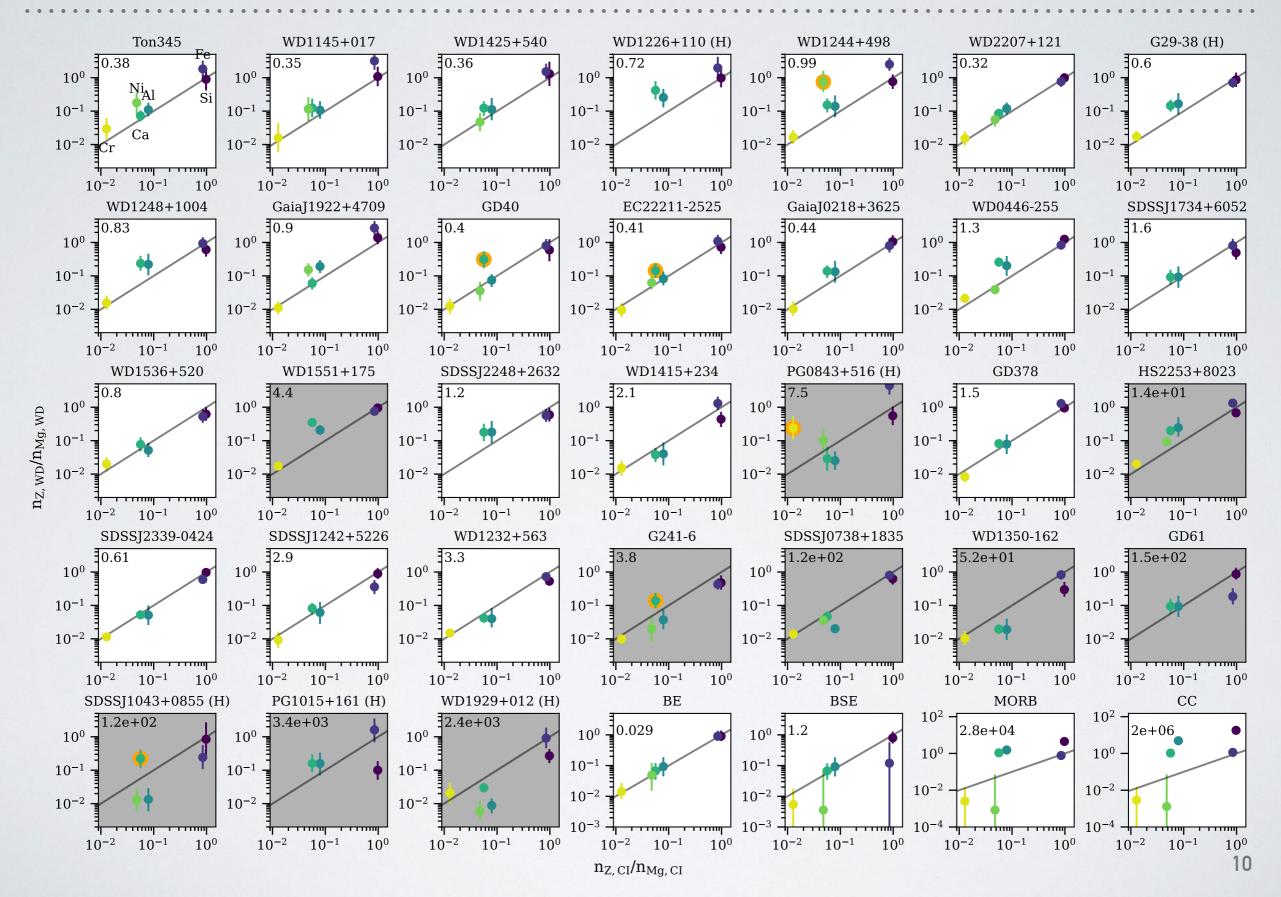
Chondrites are best representation of Solar System compositions



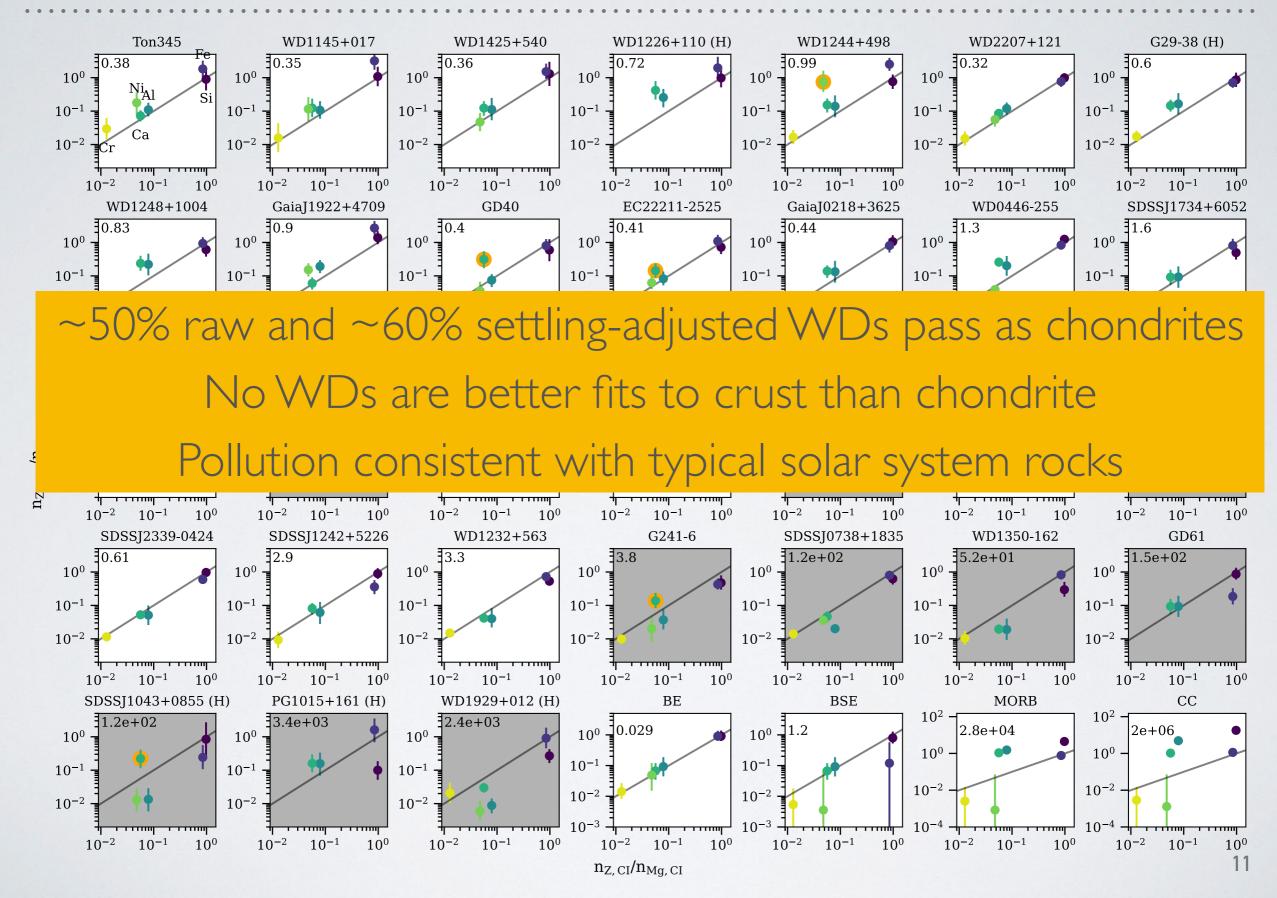
# STATISTICALLY DISTINGUISH BETWEEN CHONDRITE AND CRUST



### WHITE DWARFS ABUNDANCES

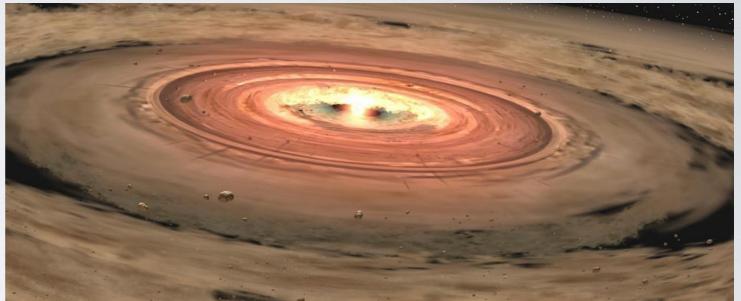


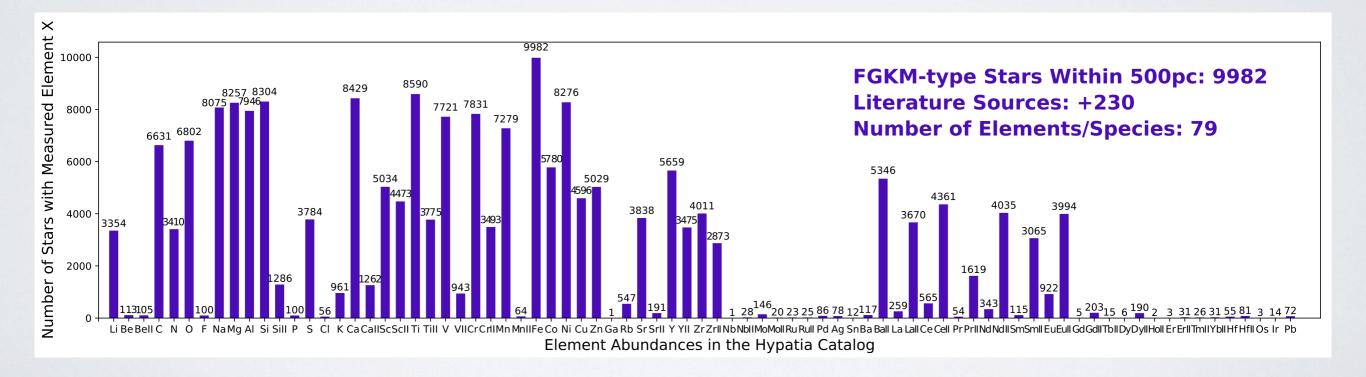
#### WHITE DWARFS ABUNDANCES



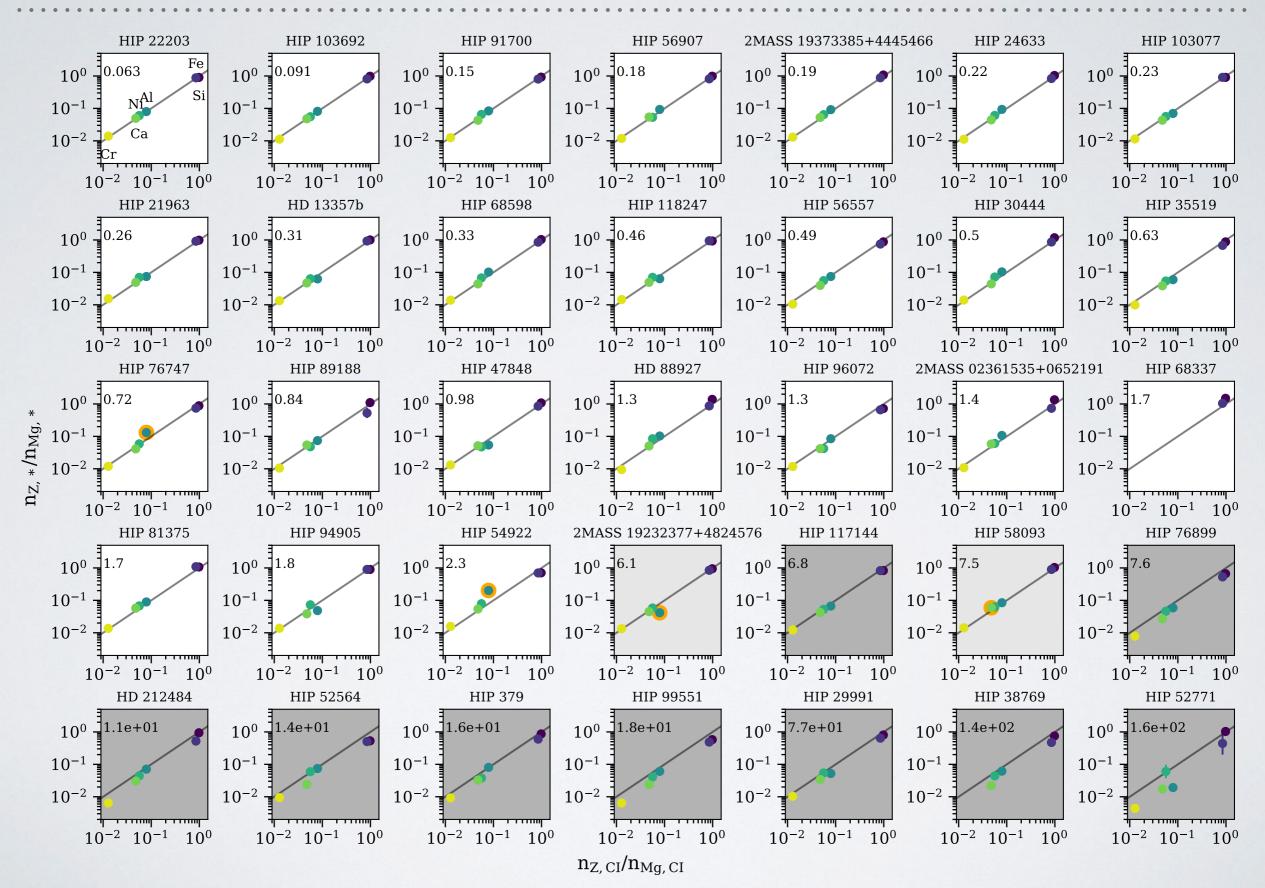
# **HYPATIA CATALOGUE STARS**

Stellar abundances represent protoplanetary disk environments and correlate with planet compositions (Thiabaud+ 2015, Bonsor+ 2021)



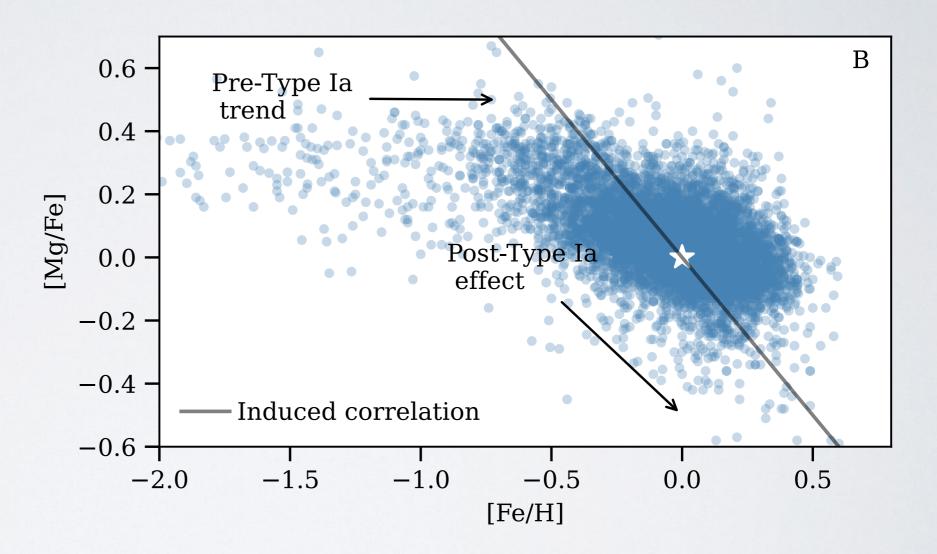


# **HYPATIA CATALOGUE STARS**



# **CONSIDERING EFFECTS OF GALACTIC CHEMICAL EVOLUTION**

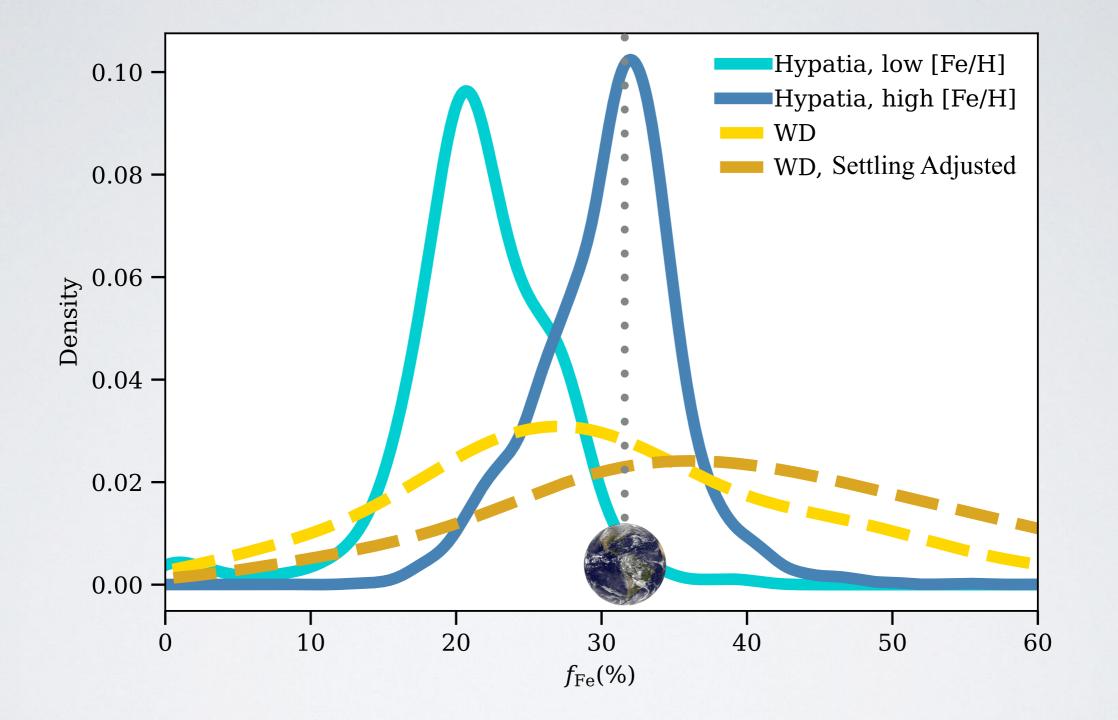
- Metallicity [Fe/H] increases with time
- Change in lithophile/ siderophile ratios due to late effects from Type I a supernovae







#### **EFFECTS OF GALACTIC CHEMICAL EVOLUTION – SMALLER CORES?**



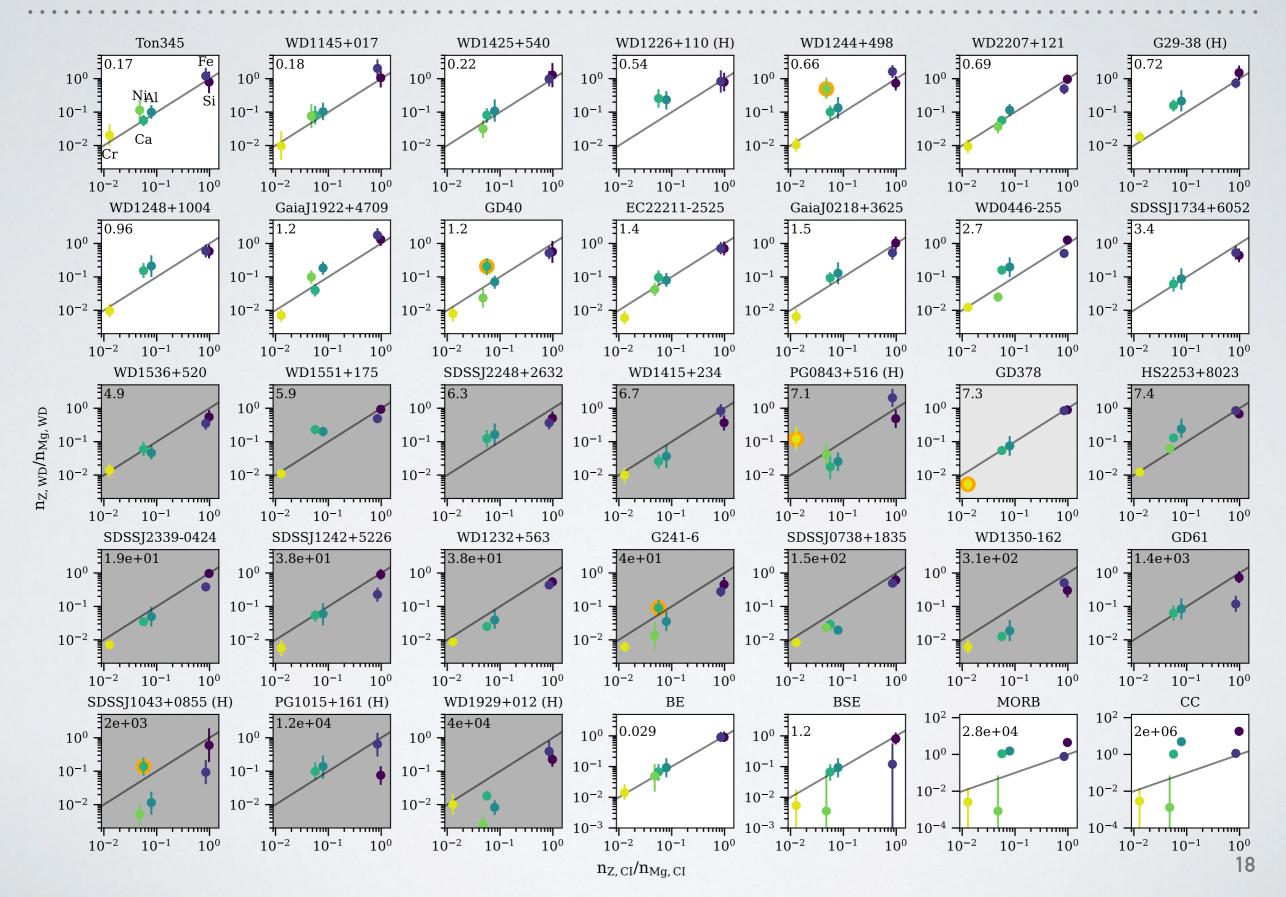
Expect smaller cores at early times or in very low metallicity regions of the galaxy

# SUMMARY

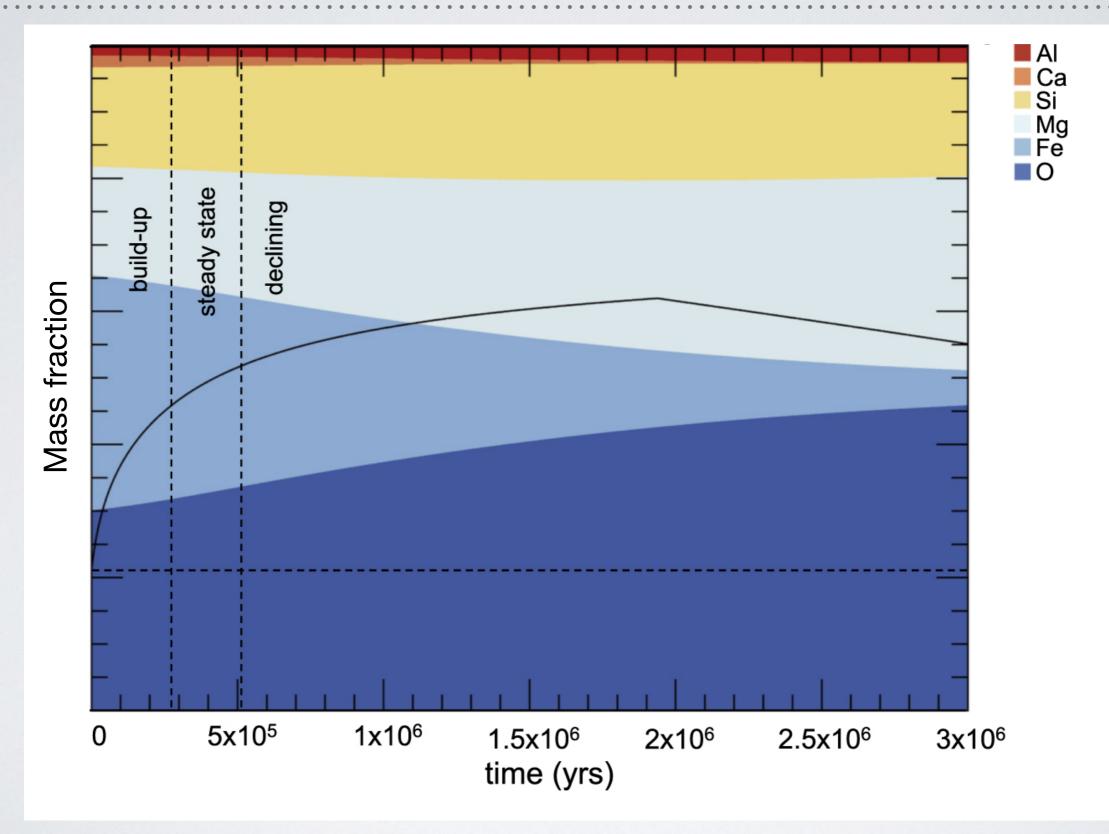
Solar system rocks are typical in the solar neighborhood

- >50% of exo-rocks sampled by polluted WDs are consistent with chondritic material
- ~75% of local stars sample by Hypatia Catalog are consistent with chondrites
- No evidence for accretion of crust by WDs
- On galactic scales, chemical evolution may lead to systematically smaller iron cores in planets at early times, but galactic effects are likely not impacting the current sample of polluted WD compositions

#### WHITE DWARFS - RAW DATA

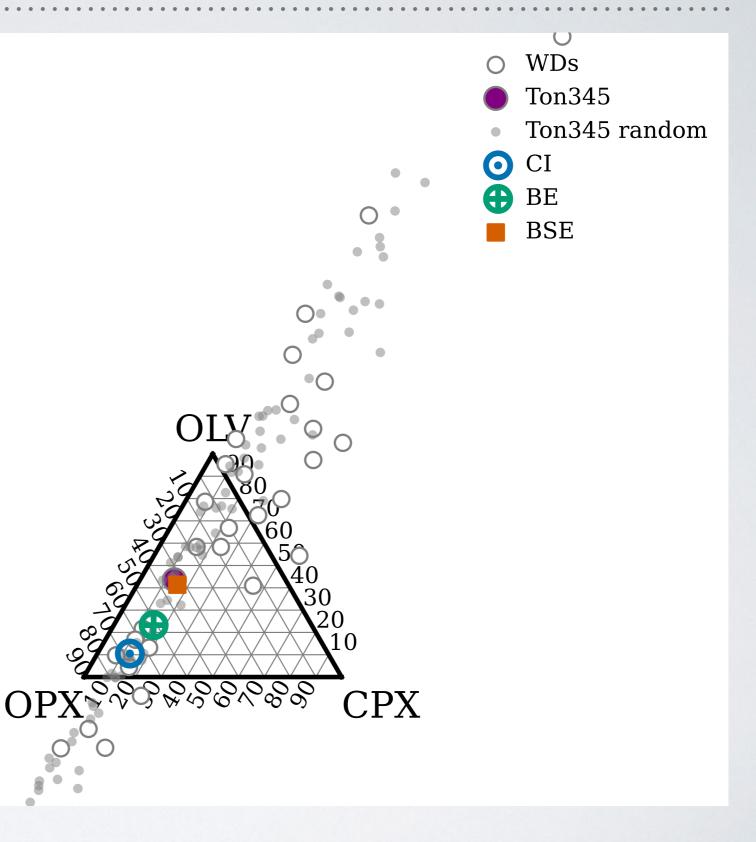


# **DIFFERENTIAL SETTLING INFLUENCES MEASURED COMPOSITION**

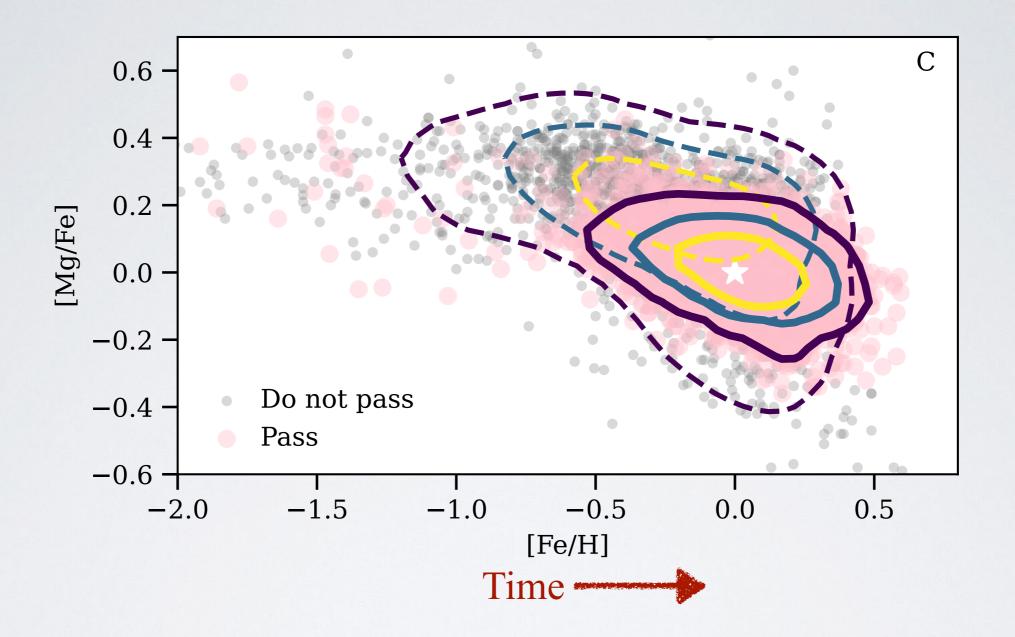


# **CLASSIFYING WHITE DWARF POLLUTION BY MINERALOGY**

Uncertainties in Mg, Si, Ca abundances too large to constrain mineralogy

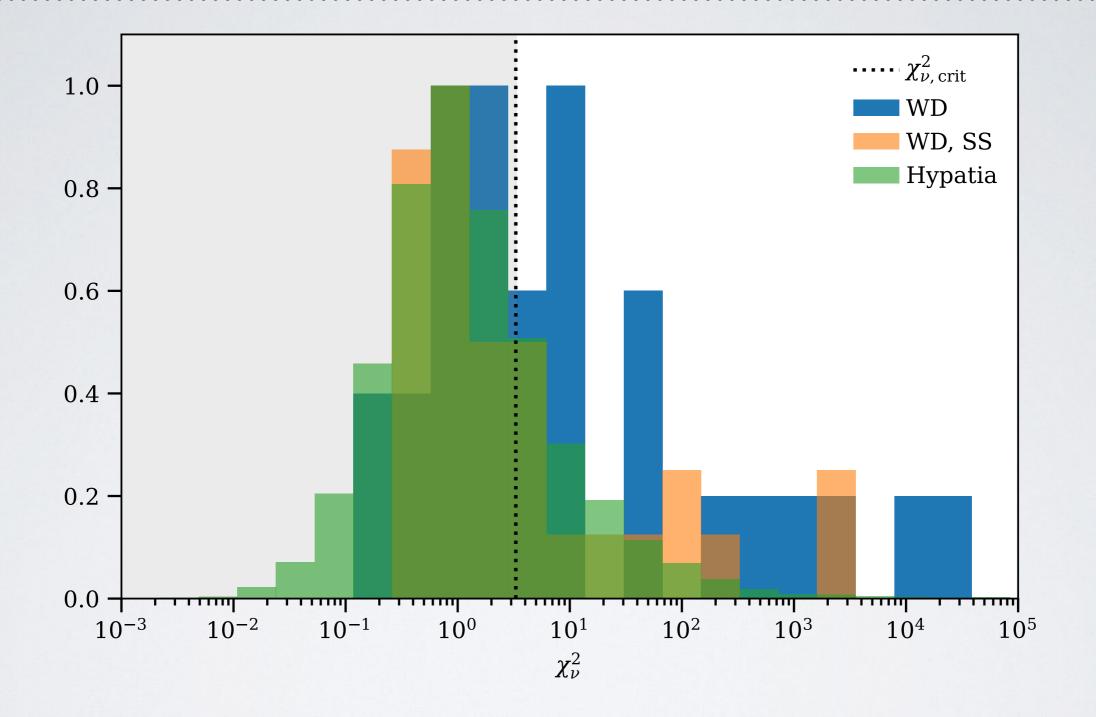


# **CONSIDERING EFFECTS OF GALACTIC CHEMICAL EVOLUTION**



- Lithophile/siderophile ratio (Mg/Fe) changes with time
- Lower metallicity stars (forming at earlier times) tend to be less consistent with chondrites

### **HYPATIA CATALOGUE STARS ARE CHONDRITIC**



~75% of Hypatia Catalogue stars pass as chondritic