

# LYRA

## Self-interacting dark matter: What dwarf galaxies can teach us about our Universe



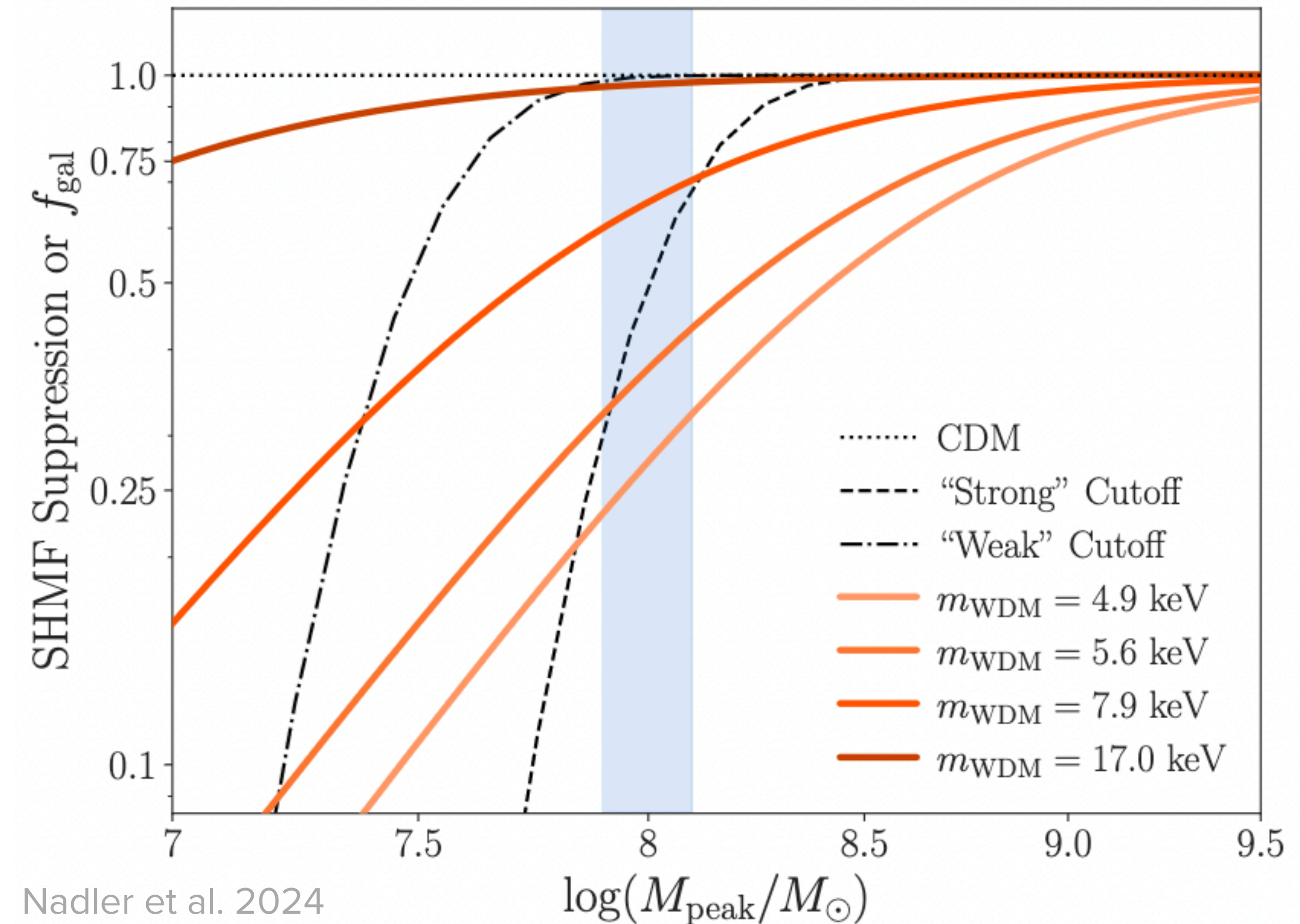
**THALES A GUTCKE**  
**IFA — UNIVERSITY OF HAWAII**

With Volker Springel, Rüdiger Pakmor, Christoph Pfrommer  
Azadeh Fattahi, Shaun Brown, Joaquin Sureda, Finn Giddings

# RELEVANCE OF DWARF GALAXIES FOR DARK MATTER RESEARCH

- One of the tightest **constraints on dark matter** models
- Cusp - core question at the centers of dwarf galaxies
- Constrains **feedback** models
- **Smallest dwarfs** might provide the tightest constraints

## WHICH DWARFS CONSTRAIN DM BEST?



# EXPLOSION OF DWARF GALAXY RESEARCH BEYOND THE LOCAL GROUP

## Until now:

<100 dwarfs known fainter than  $M_V > -10$

Almost all associated with galaxy/cluster

## Now:

DESI - detection

DECam - detection

HST/JWST - follow up

EUCLID - detection

## Soon:

4MOST - follow up

Vera Rubin Observatory / LSST - detection

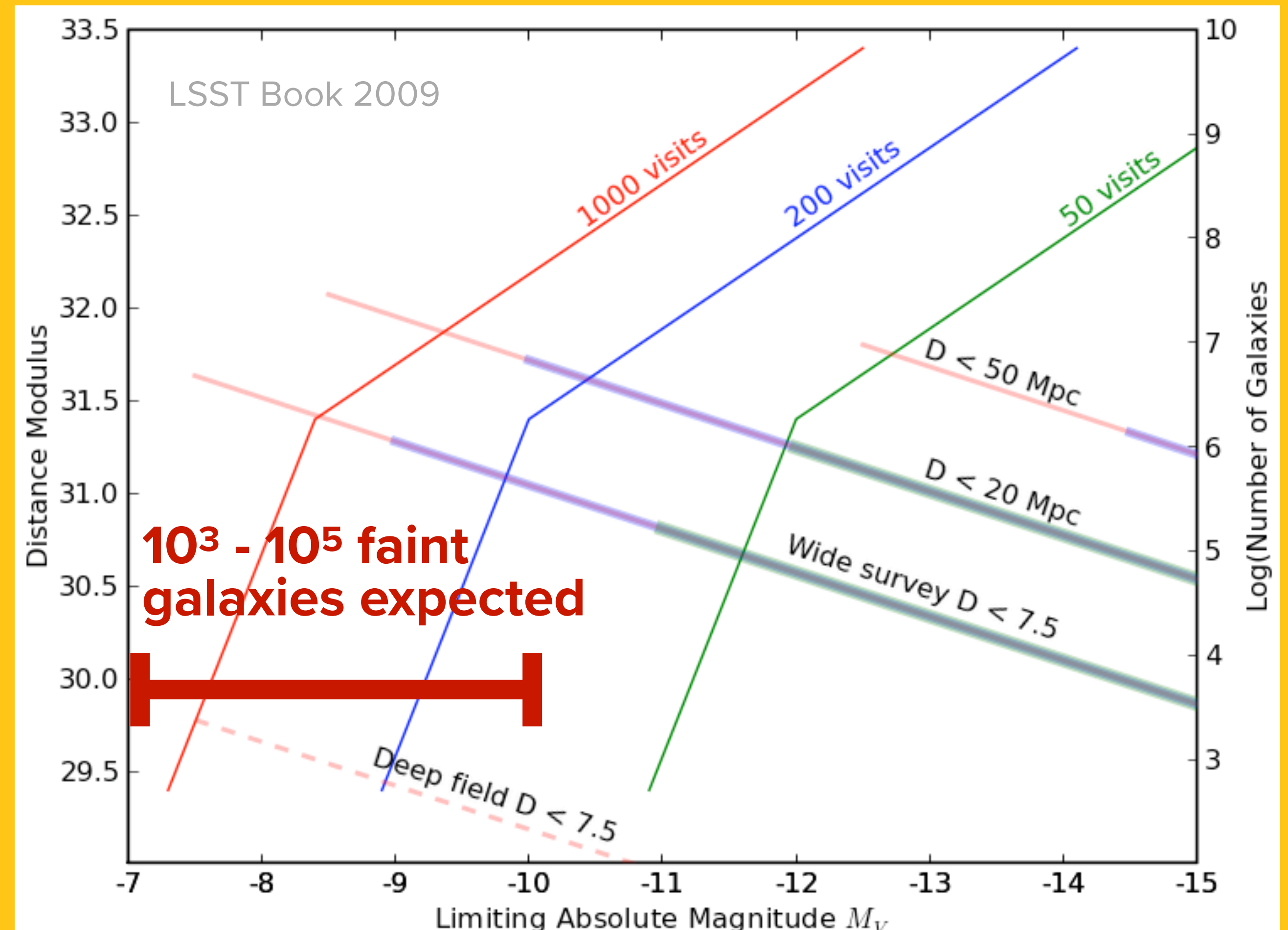
Roman telescope

## Future:

Wide-field spectroscopic telescope

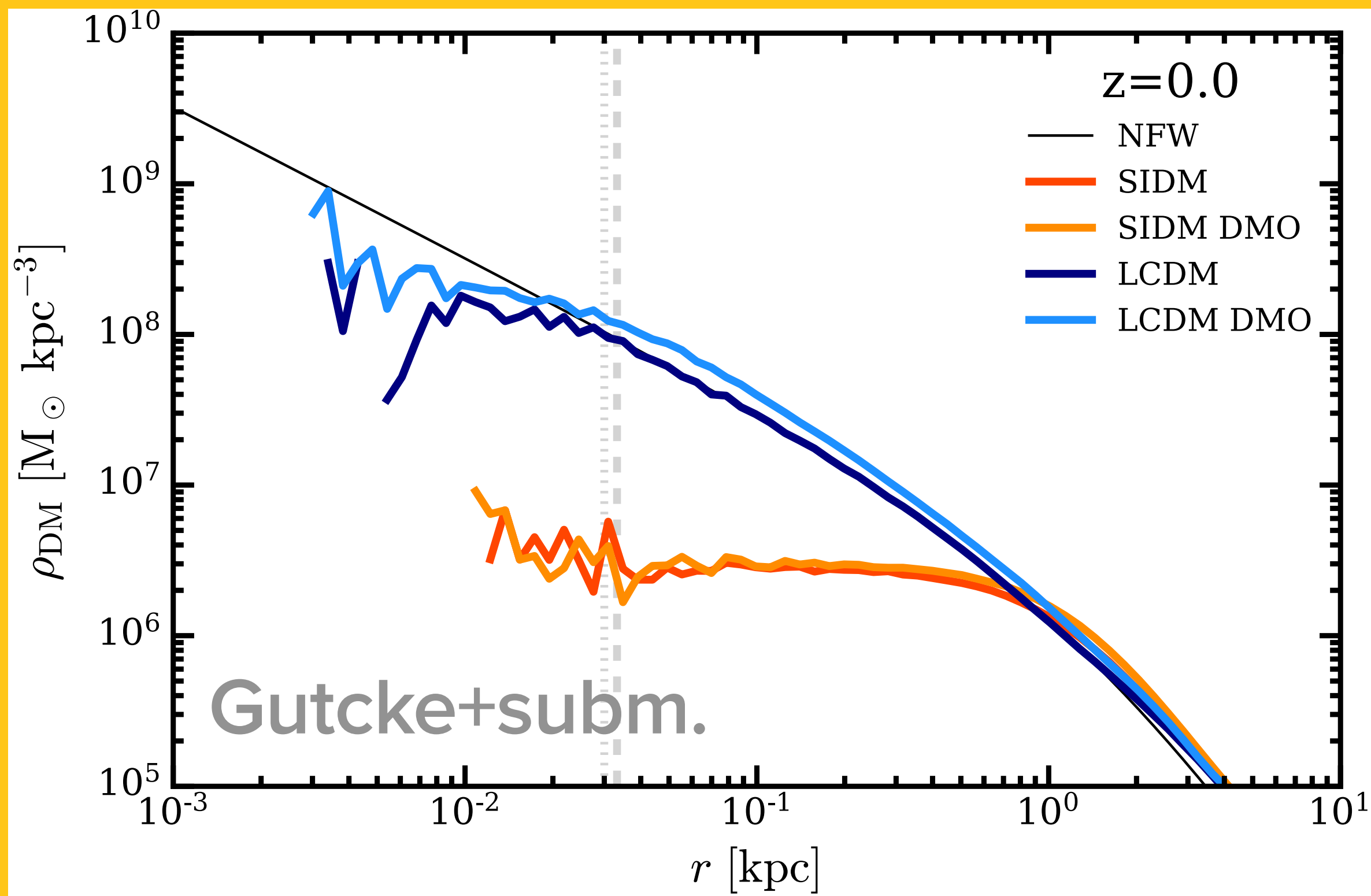
Arrakhis satellite

## DWARF GALAXIES WITH RUBIN

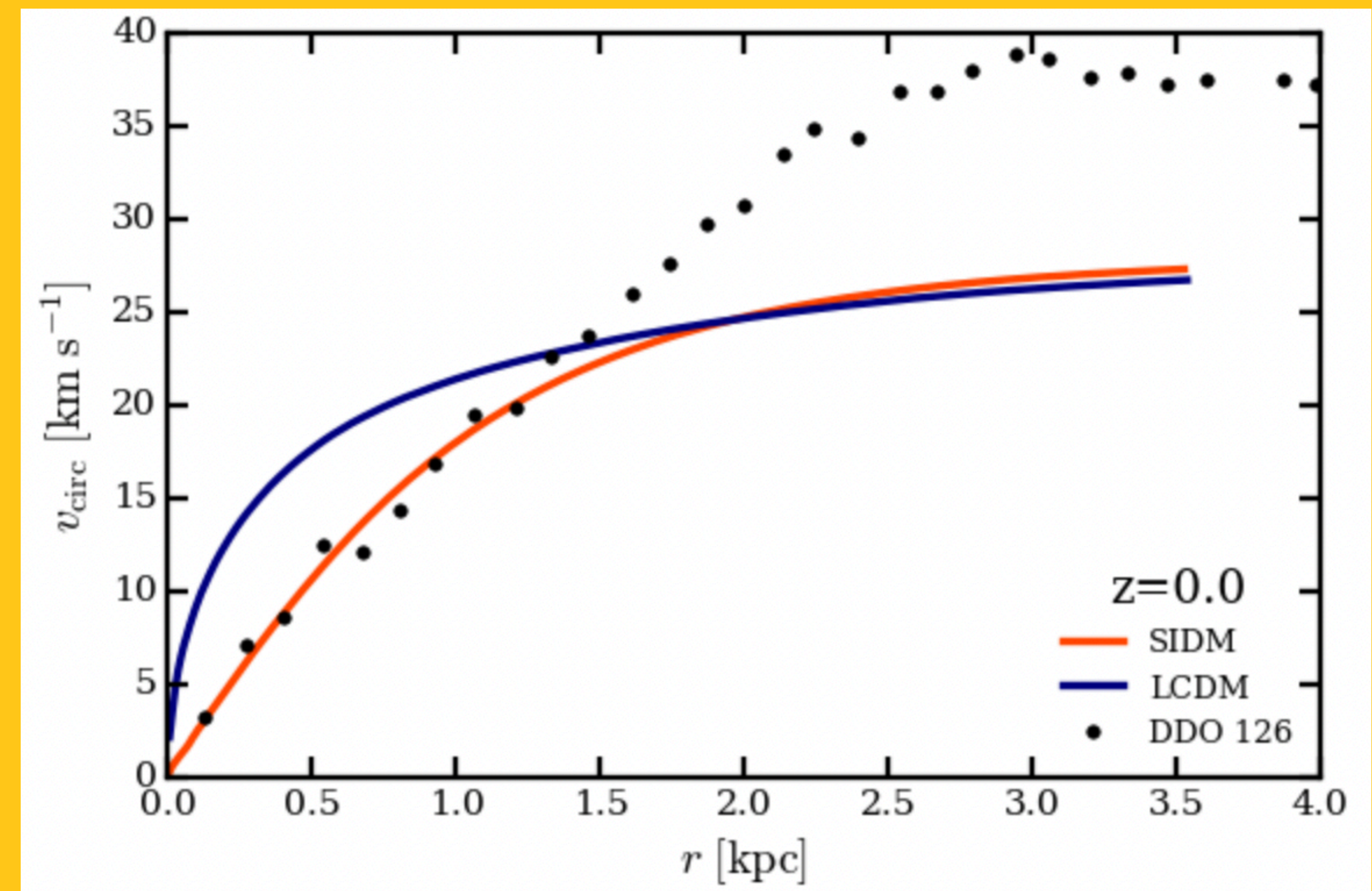


# DM DENSITY PROFILE AND ROTATION CURVE

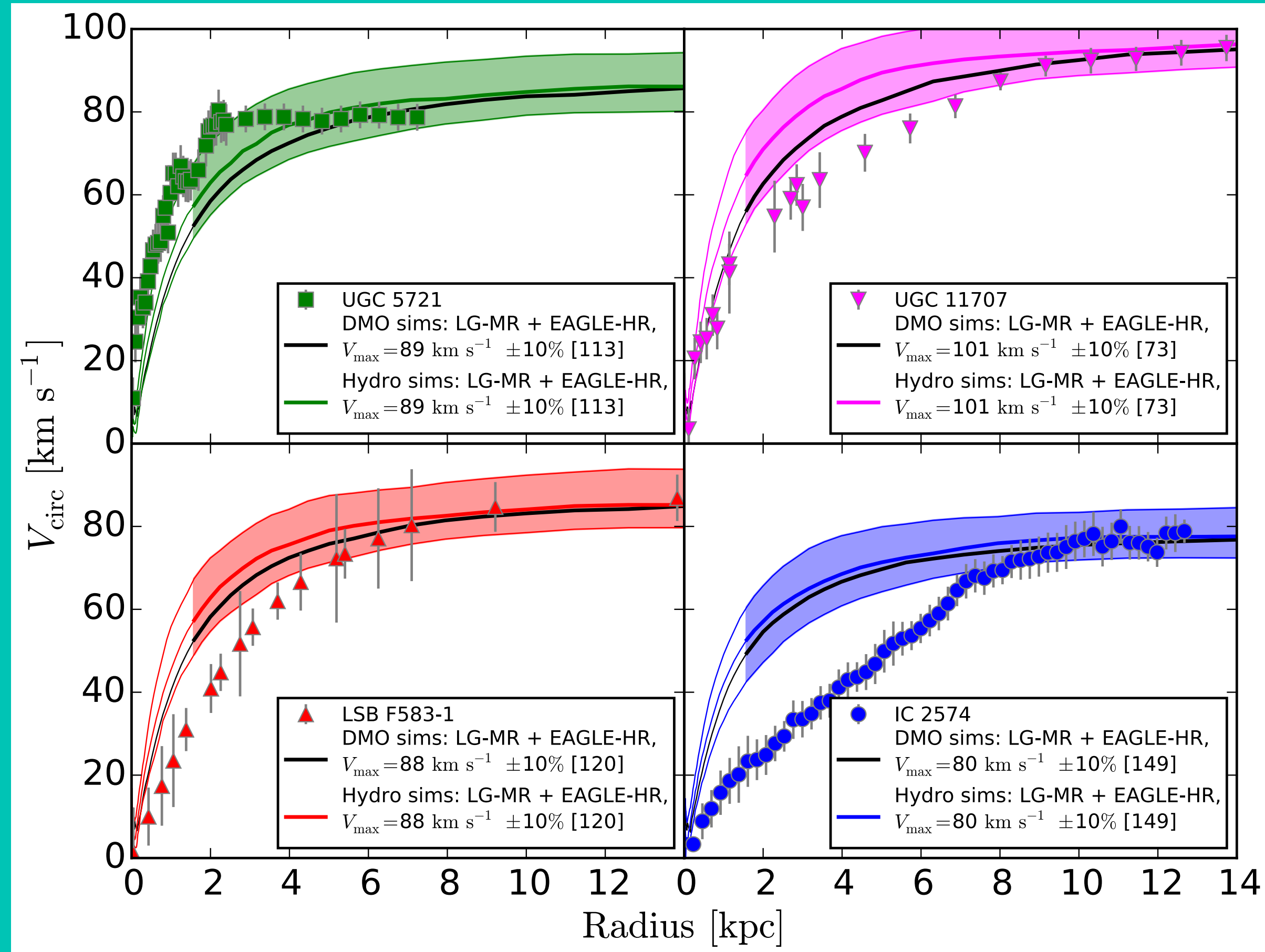
## DM profiles at z=0



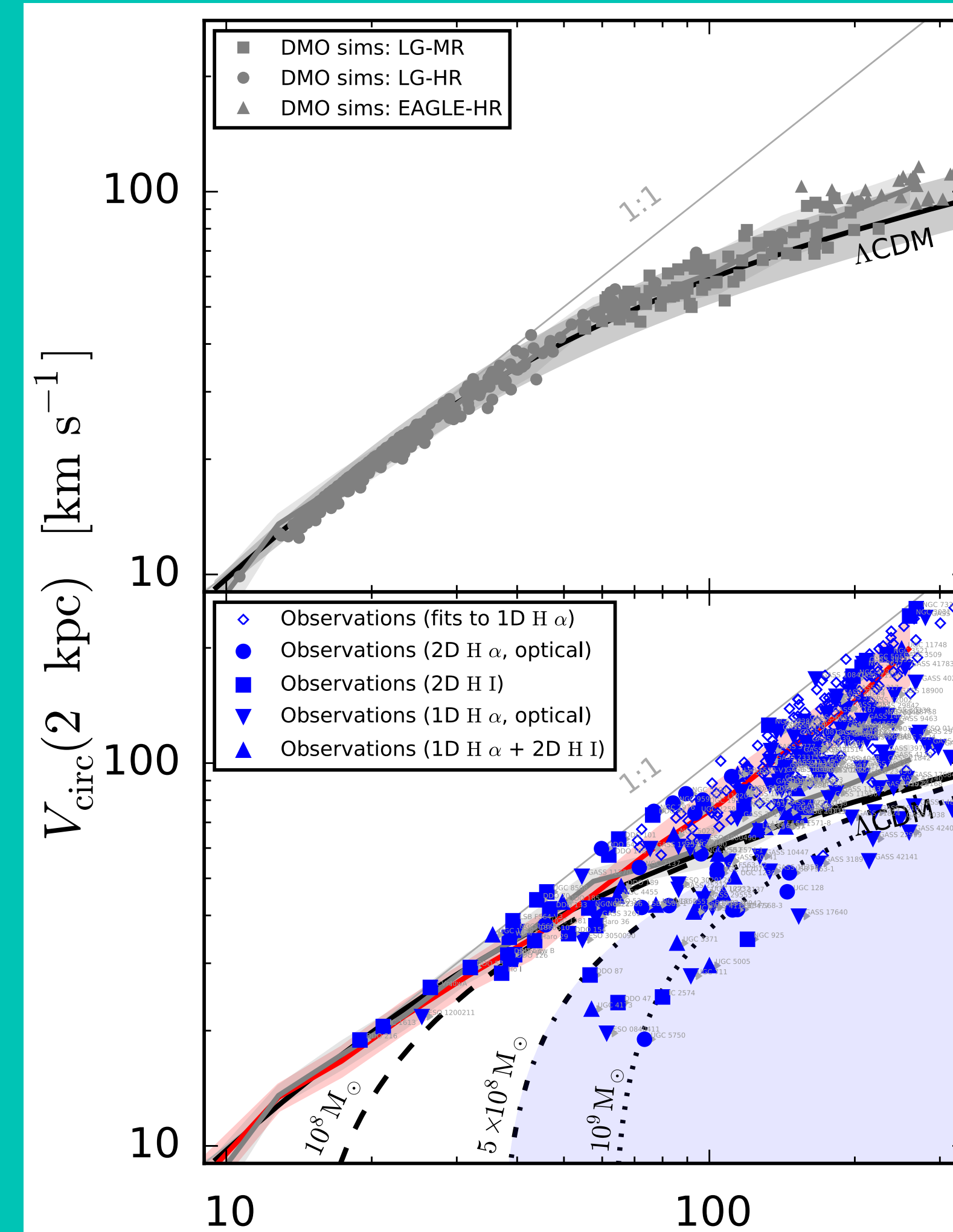
## Rotation curves



# DIVERSITY OF DWARF GALAXIES



Oman et al. 2022

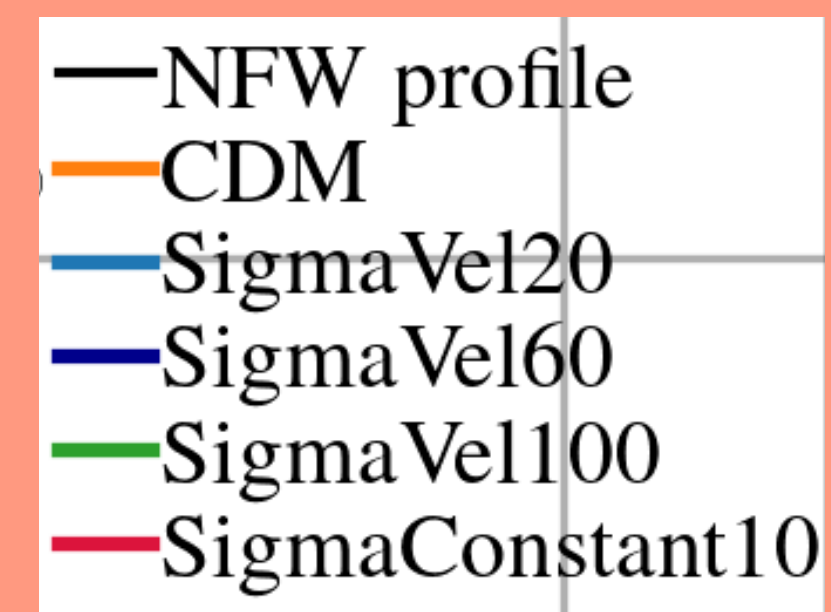
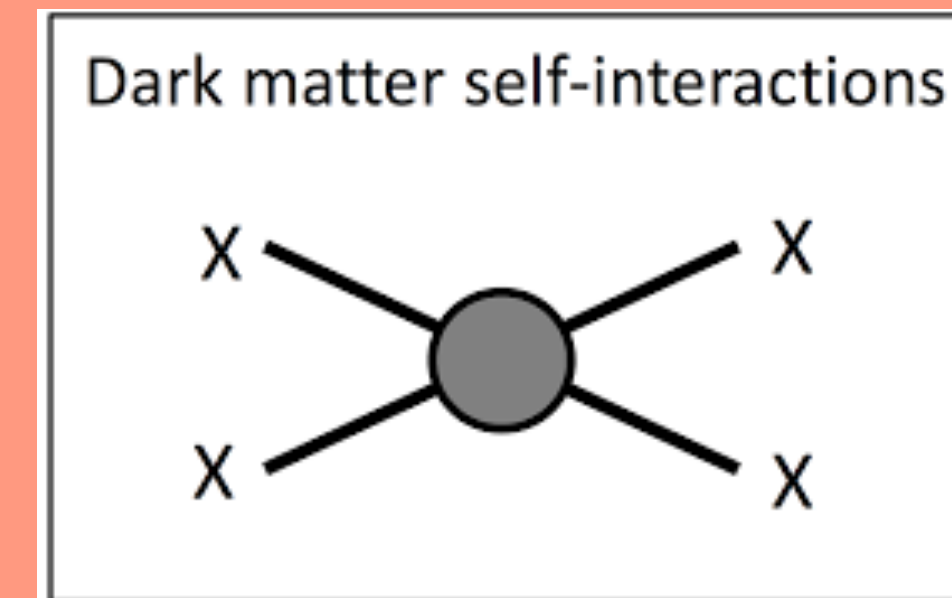


All flat

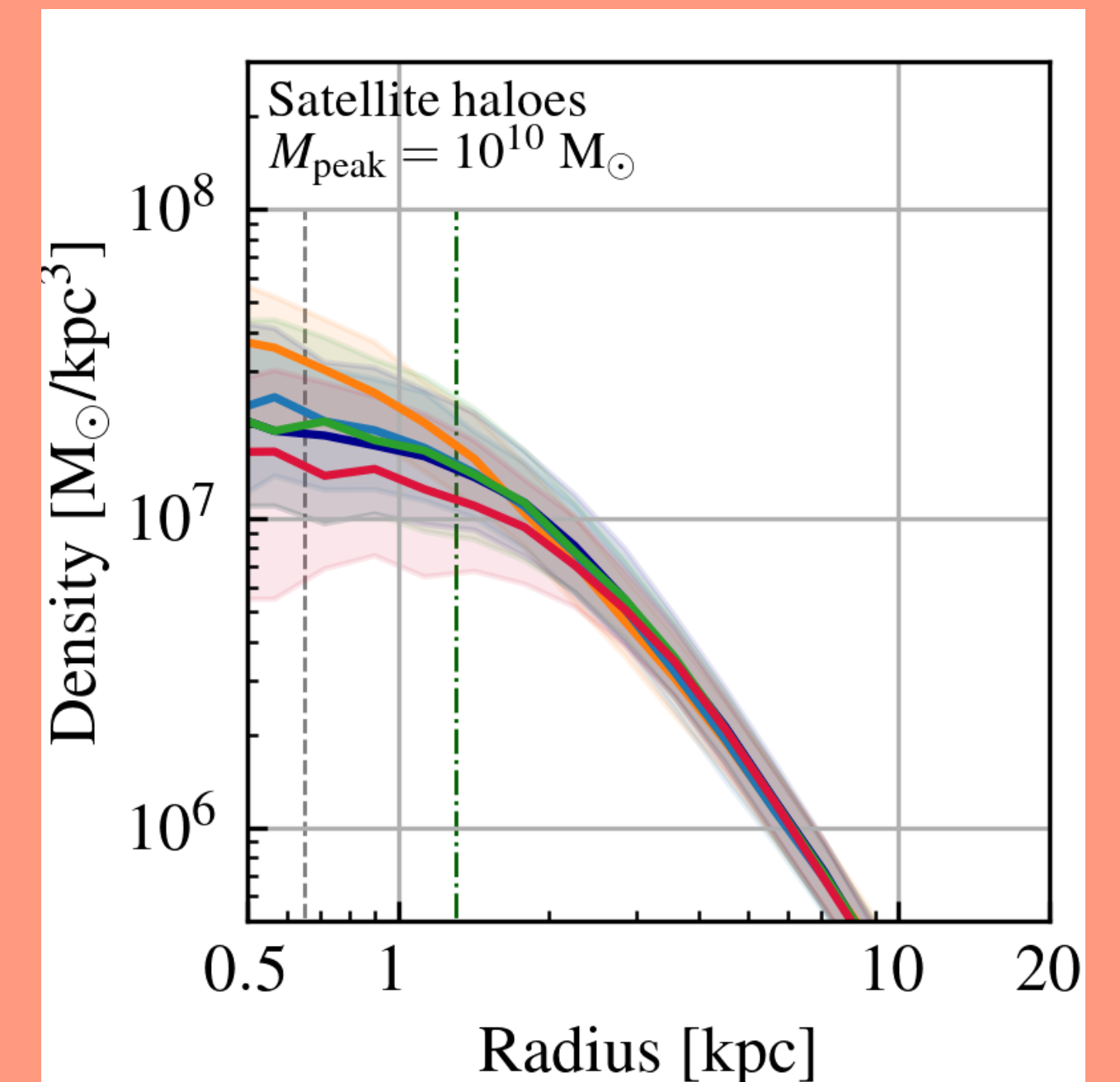
Some flat, some peaked

# SELF-INTERACTING DARK MATTER (SIDM)

- Proposed interaction between dark matter particles
  - Requires invoking a **new gauge boson**
  - Allows scattering or annihilation of dark matter
- Most common parameter is the **scattering cross section**
  - Now often includes a velocity dependency
- Mostly testing as dark matter-only
  - Need to study **with baryonic physics!**



Correa et al. 2022

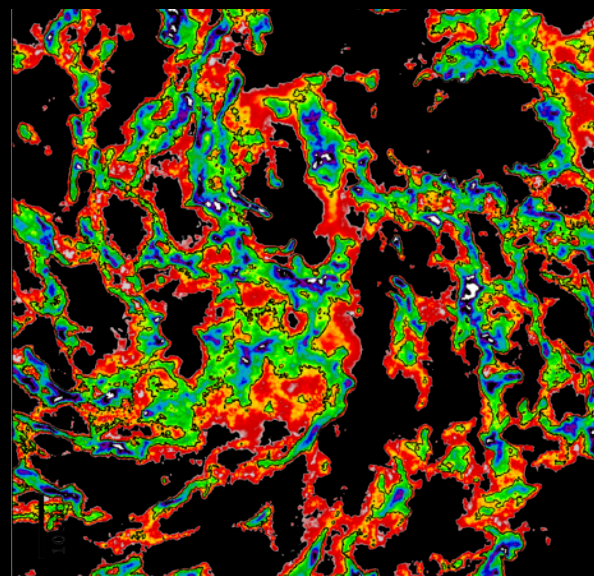


# THE LYRA MODEL

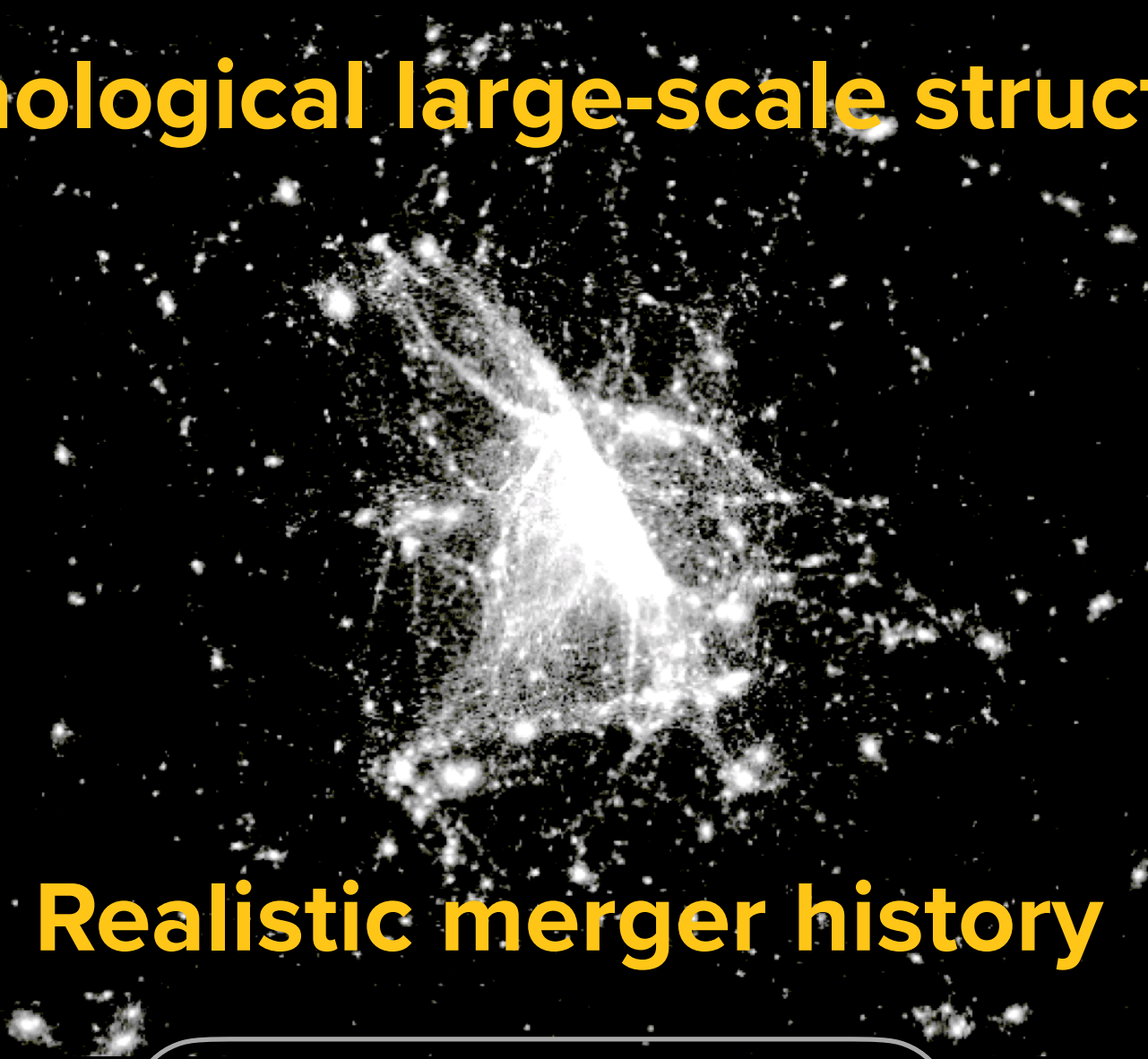
Model (Gutcke+2021)  
built within AREPO code

## Cosmological large-scale structure

### Multi-phase ISM

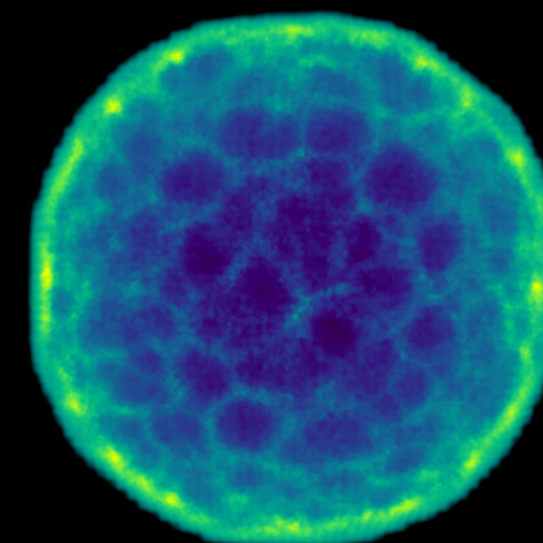


Molecular clouds



Realistic merger history

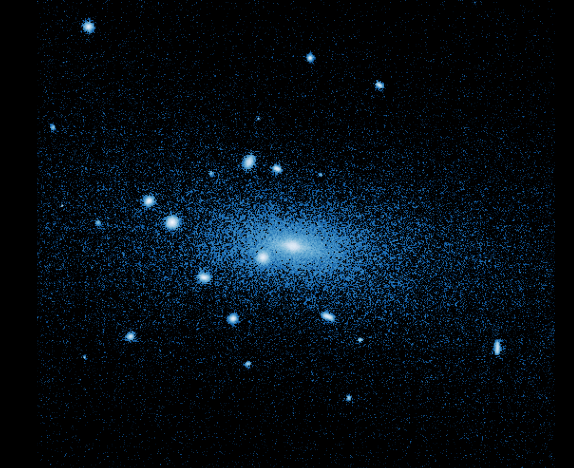
### Resolved



supernovae

## Emergent features

### Star clusters



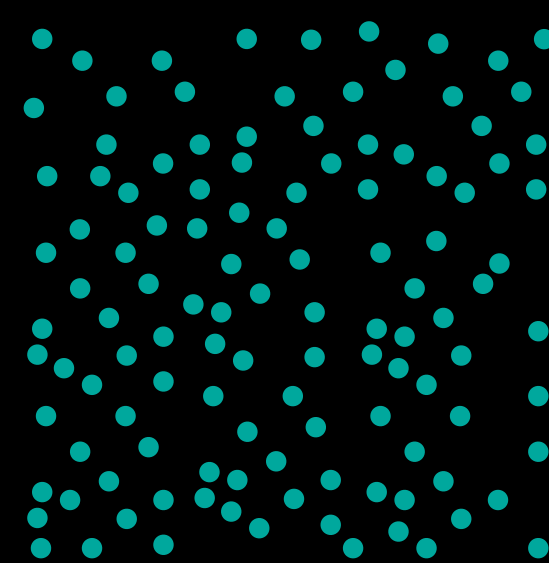
Clustered SNe

### Individual stars



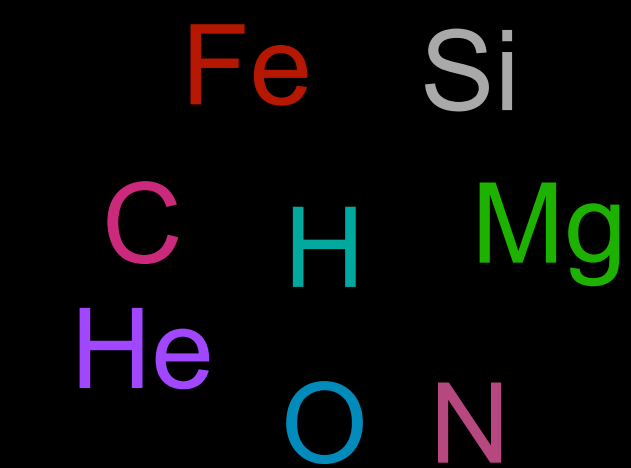
follow IMF

### Metal seeding



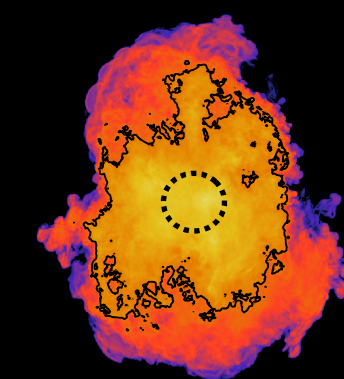
in PopIII halos

### Discrete



chemical yields

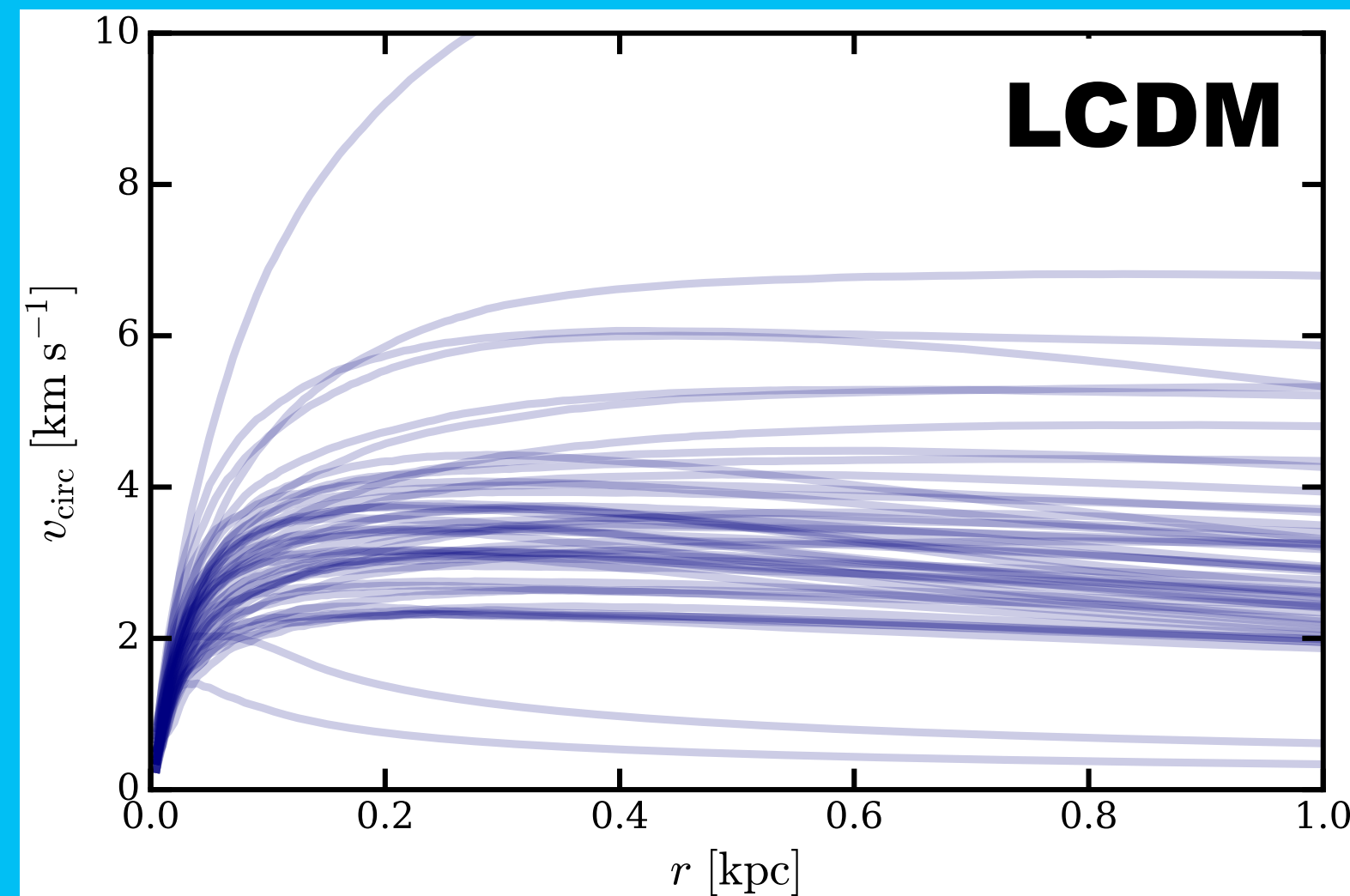
### SN-driven outflows



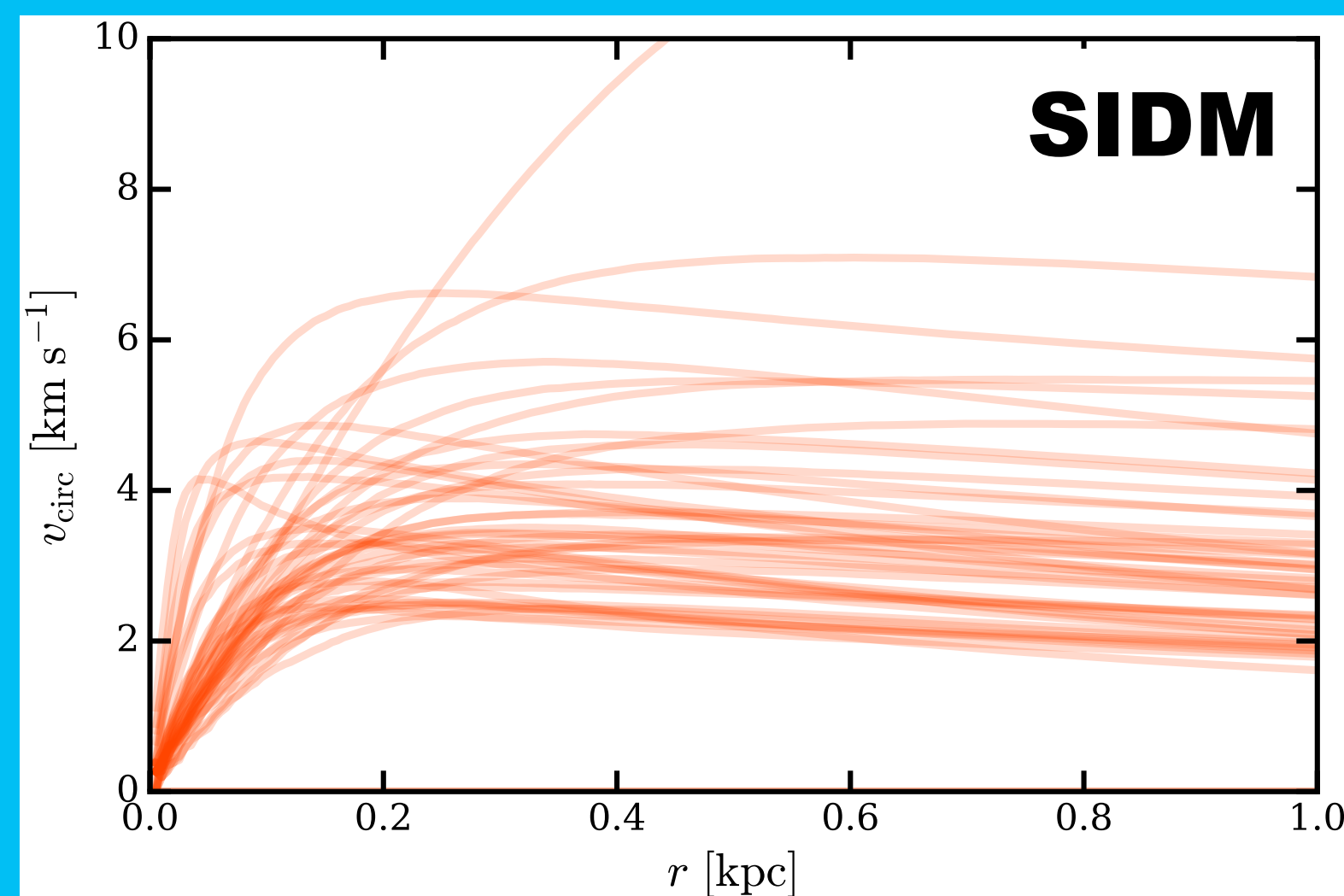
Realistic CGM

# DIVERSITY OF ROTATION CURVES WITH SIDM

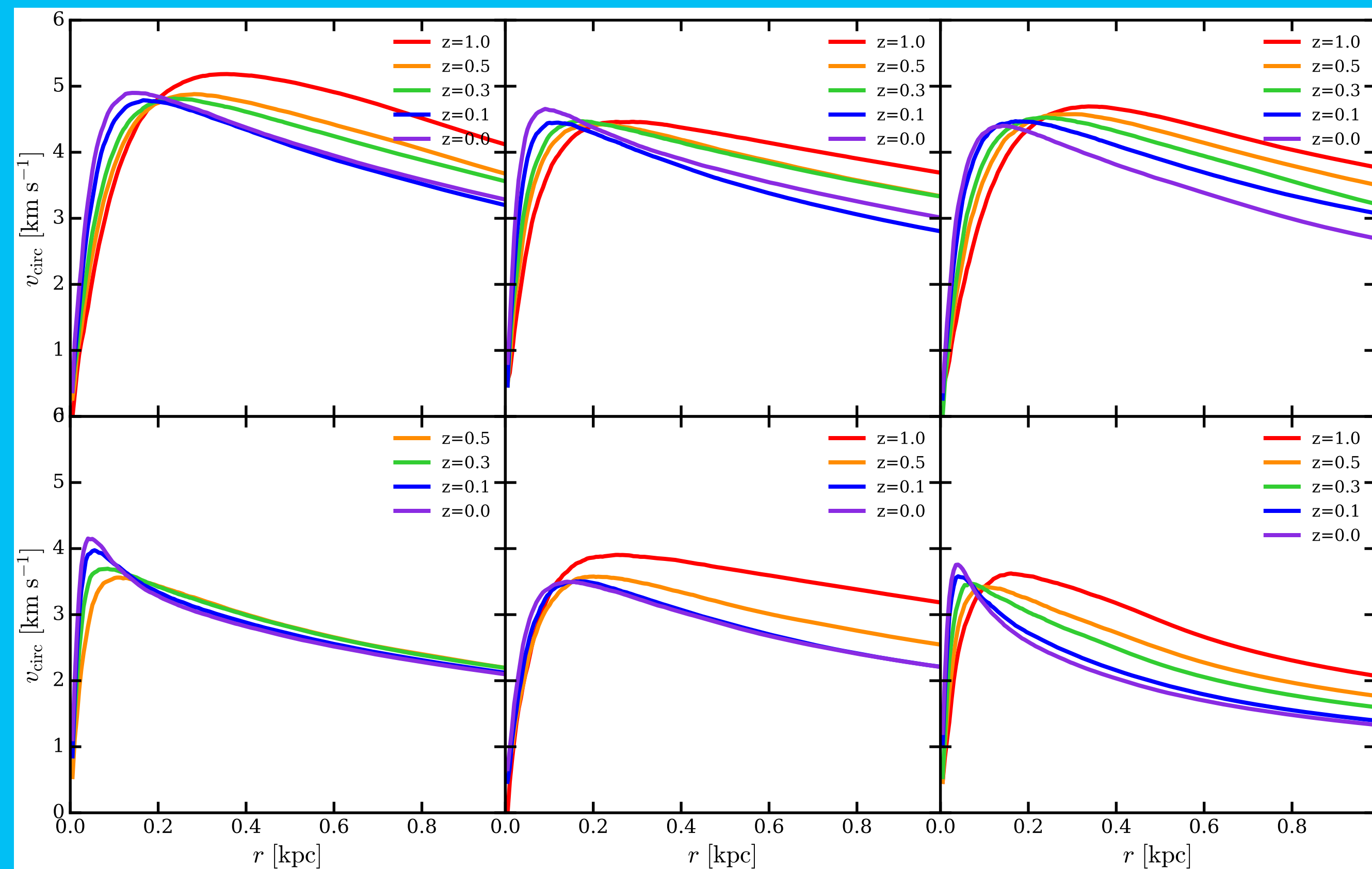
All flat



Some flat, some peaked



## Core-collapsing subhalos

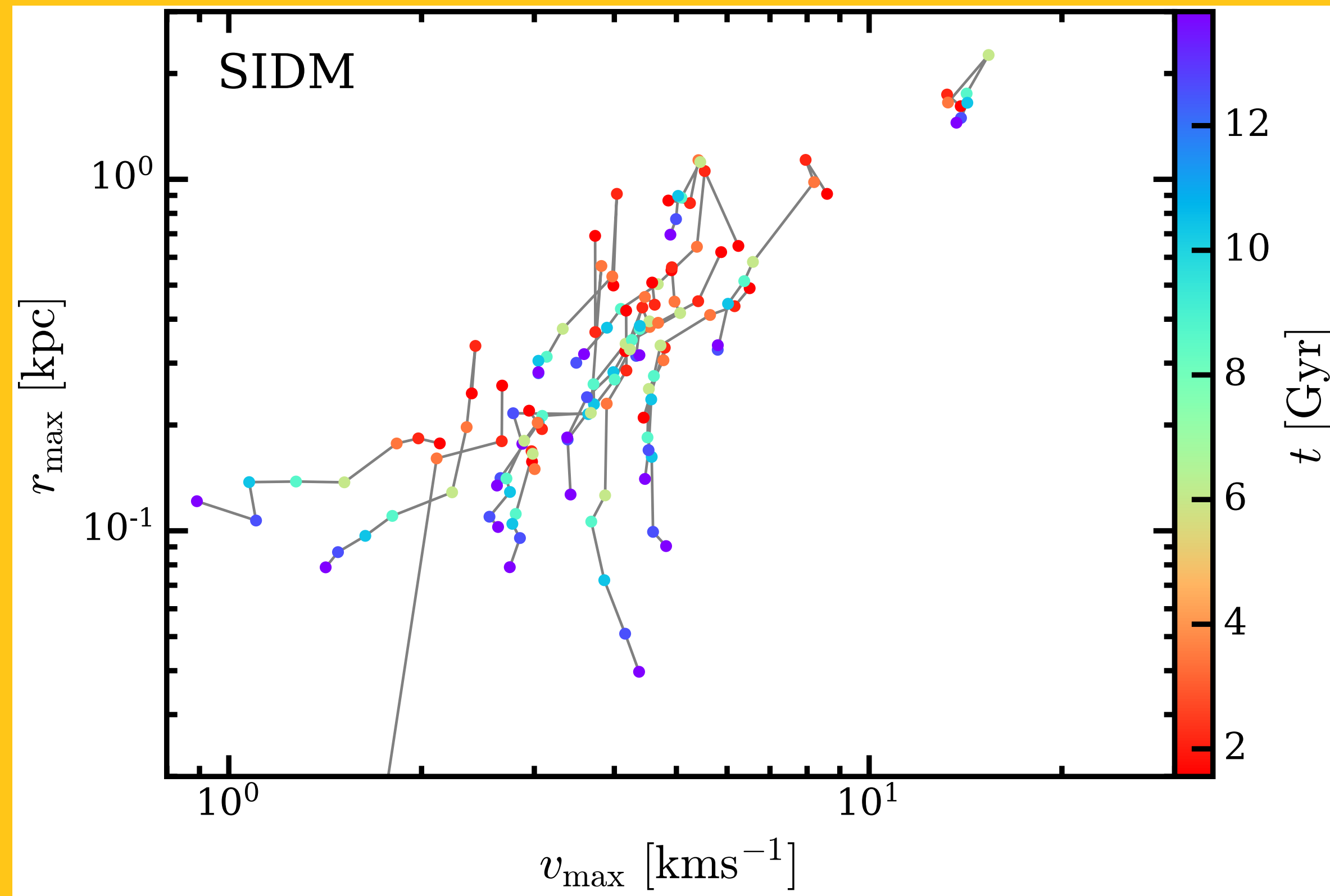


Gutcke+subm.

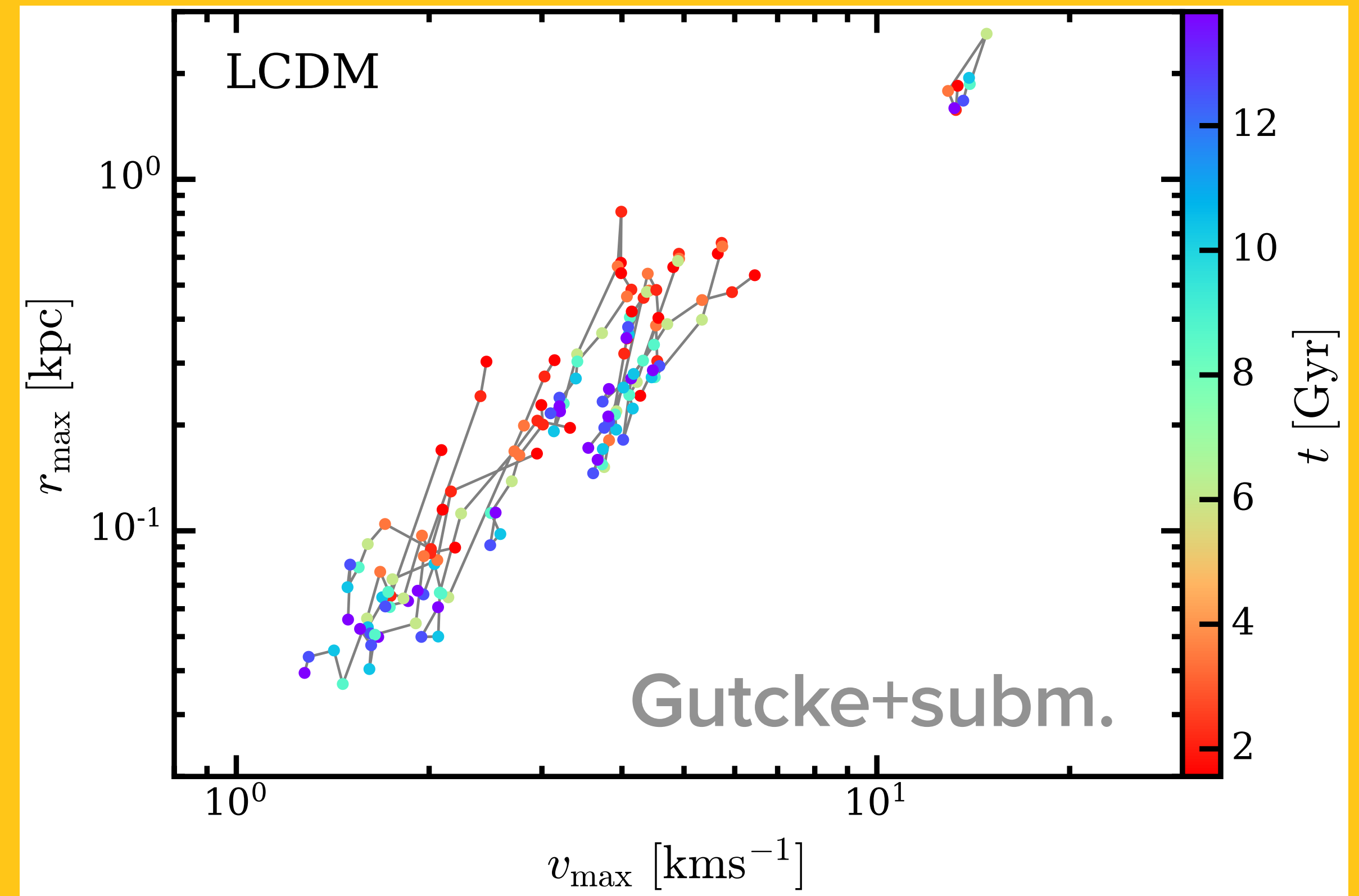


# CORE-COLLAPSED SATELLITES

SIDM

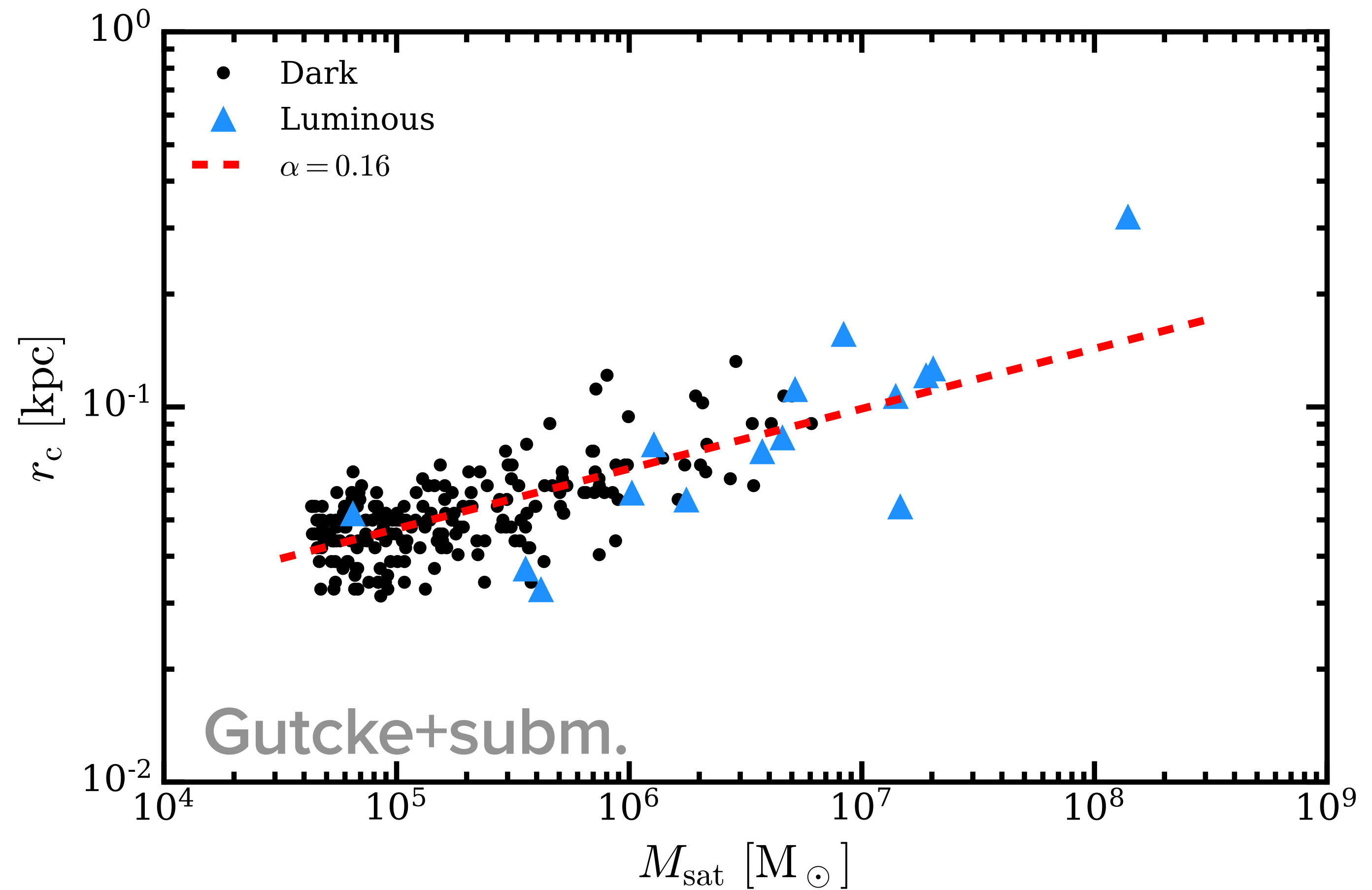


LCDM



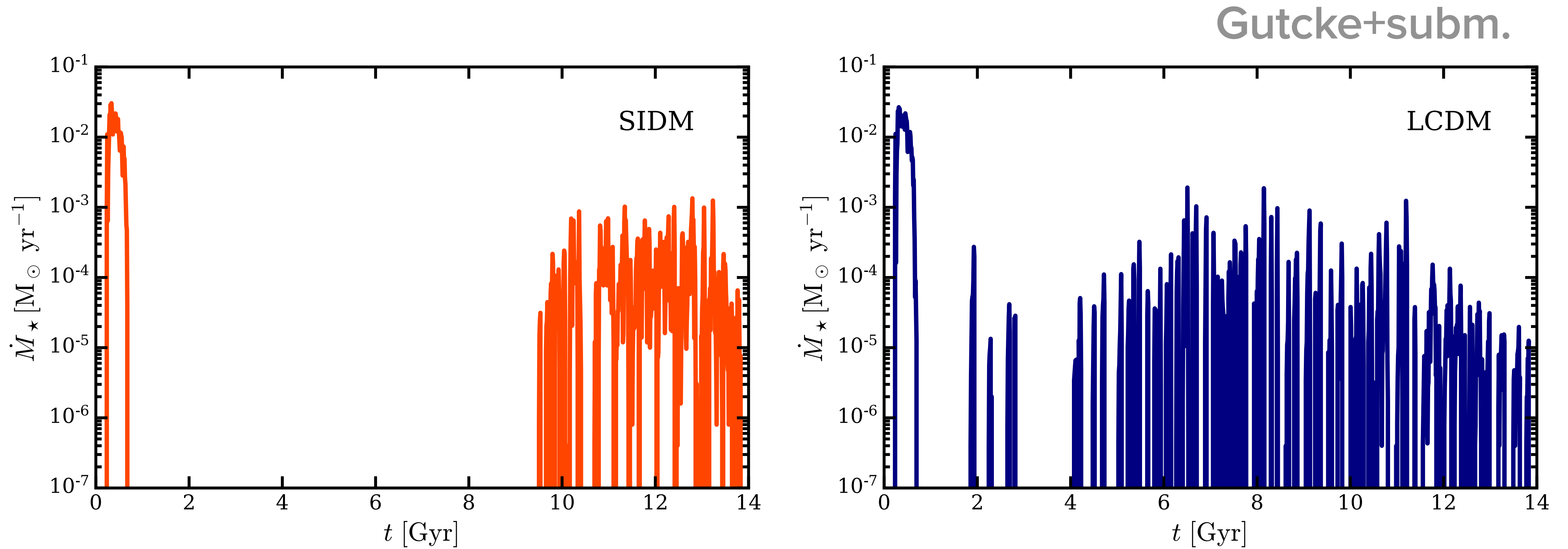
# CORE SIZES IN SATELLITES

Power law  
relation between  
core size and  
dynamical mass



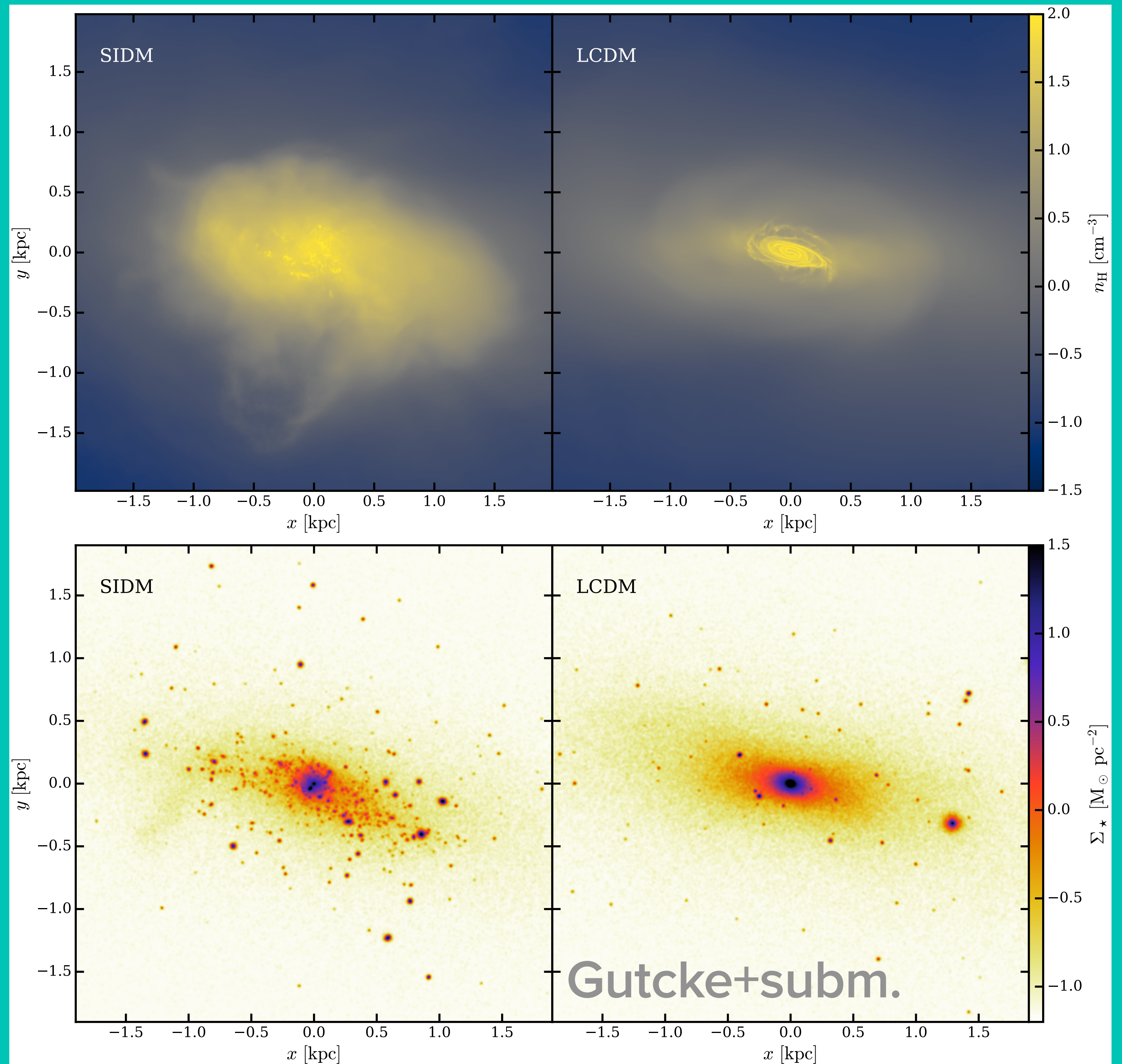
# LOOKING AT THE CENTRAL GALAXY

# STAR FORMATION HISTORY IS DELAYED



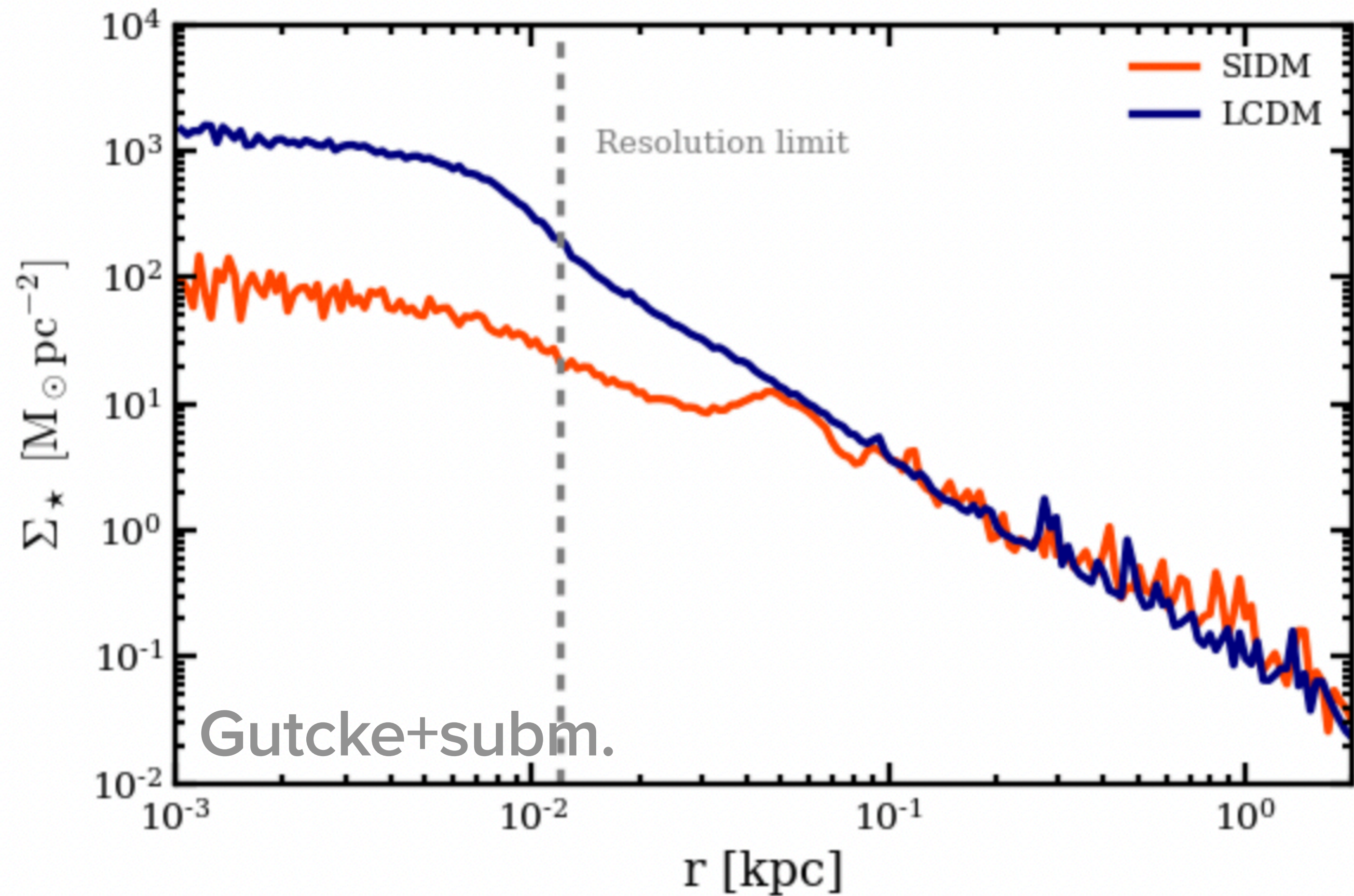
# EFFECT OF SIDM ON GAS AND STARS

- Flatter core region allows gas to be extended
- Star formation occurs in various star forming regions
- Star clustering likely over-estimated due to lack of 2-body evaporation



# STELLAR SURFACE DENSITY

Observable  
reduction of  
stellar density  
in central  
region



# SUMMARY

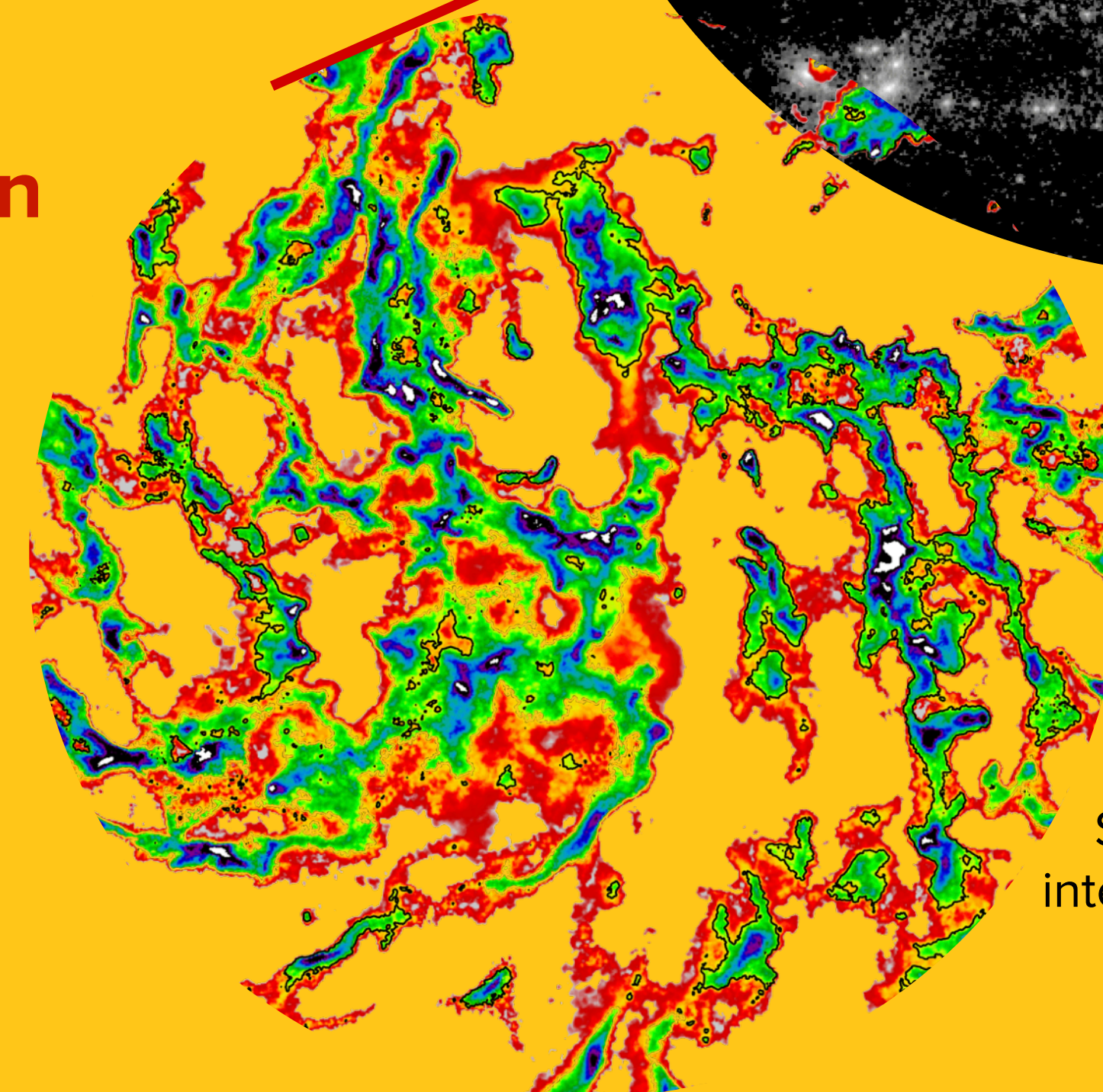
**Lyra** brings together small scale baryonic physics and the cosmological context providing a novel testing ground for SIDM

- Highly timely in light of dwarf galaxy data explosion
- Provides much needed spatial details to constrain DM while including baryonic physics

Model developments are ongoing

Let me know if you would like to collaborate!

Cosmological large-scale structure



Small-scale interstellar medium

