



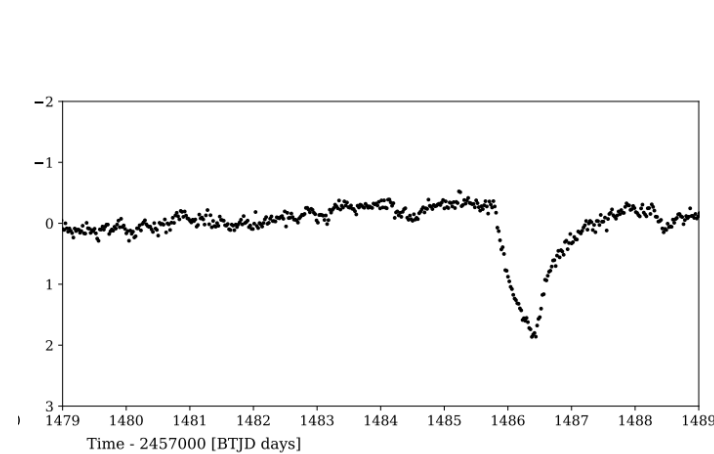
# A Deep Learning Framework for Identifying Transit Anomalies



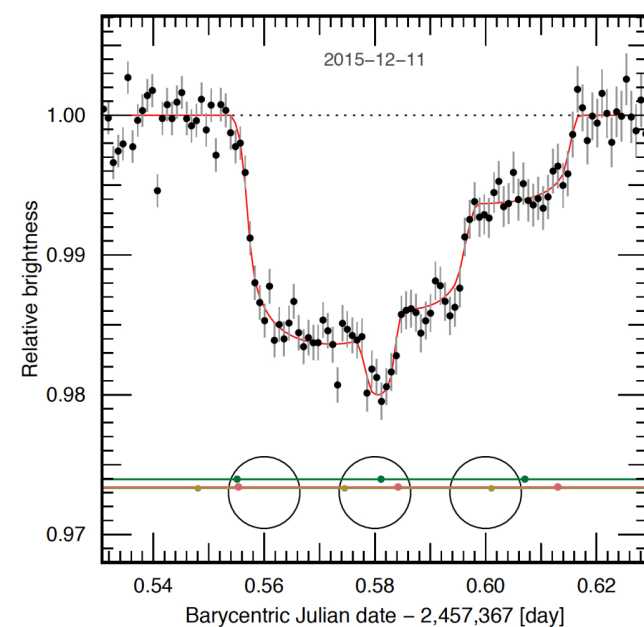
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Transit detection missions are also sensitive to **any objects that cause periodic dimming of their host stars**, from exoplanets and other well-understood astrophysical phenomena (e.g., disintegrating worlds, exo-comets, complex multi-planet systems, etc.) to possible alien megastructures. Such objects **manifest as “anomalous” transit signals** that deviate from expected behavior due to a gravitationally bound spherical planet. Here we outline a **deep learning framework** for identifying non-spherical transiting objects in missions like Kepler and TESS.

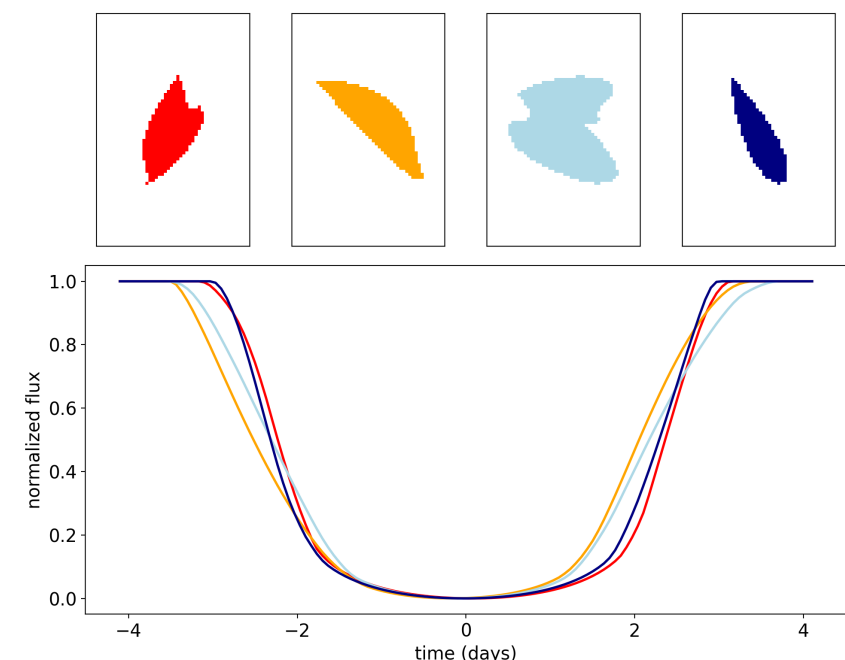
## Sources of Anomalous Transit Signals



*Exocomet tail from  $\beta$  Pictoris produces asymmetrical transit that is visibly different from circular transit models. (Zieba et al. 2020)*



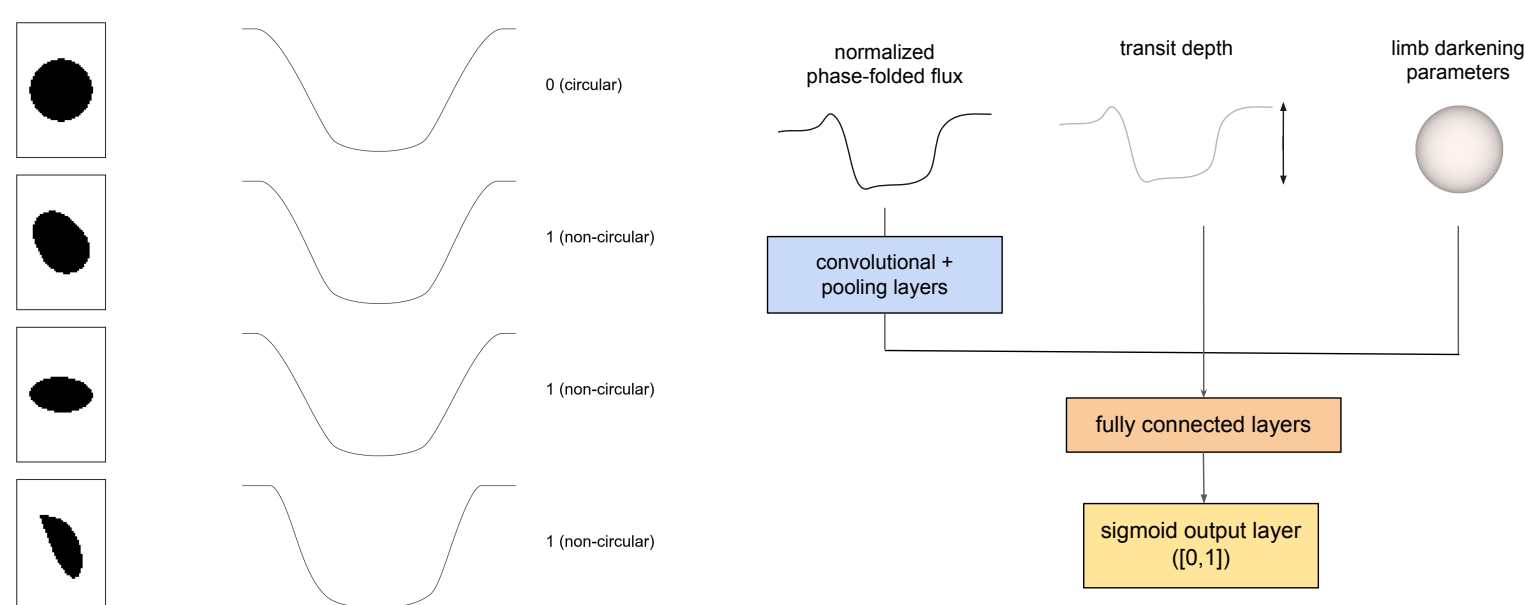
*Co-occurring transits of three circular planets in TRAPPIST-1 produce anomalous transit shape (Gillon et al. 2017)*



*Transits of different occulting shapes simulated with EightBitTransit (Sandford & Kipping 2019). These objects show subtle deviations that are most prominent at ingress and egress (Angelo et al., in preparation)*

## A Deep Learning Framework for Identifying Non-Circular Transits

The analytic relationship between a transiting object’s shape and its phase-folded light curve has degenerate solutions and is difficult to optimize. We can sidestep these issues by training a neural network to classify objects as circular and non-circular based on their light curves:



**Left:** Simulated light curve examples of different shapes are paired with associated “circular” or “non-circular” classification as training data. **Right:** Diagram of our neural network structure. Our neural network learns the relationship between an object’s shape and its light curve and host star properties.

## Data Challenges

There are a number of considerations we need to account for both in assembling representative training data and processing TESS data to analyze:

- TTV’s
- ephemeris uncertainties
- stellar activity and instrumental effects

If you have experience treating these in phase-folded Kepler or TESS data, we’d love to hear from you!

### References

- S. Zieba, K. Zwintz, M.A. Kenworthy, et al. 2020, “Transiting exocomets detected in broadband light by TESS in the  $\beta$  Pictoris system”
- M. Gillon, A. Triaud, B. Demory, et al. 2017, “Seven temperate terrestrial planets around the nearby ultracool dwarf star TRAPPIST-1”
- E. Sandford & D. Kipping 2019, “Shadow Imaging of Transiting Objects”