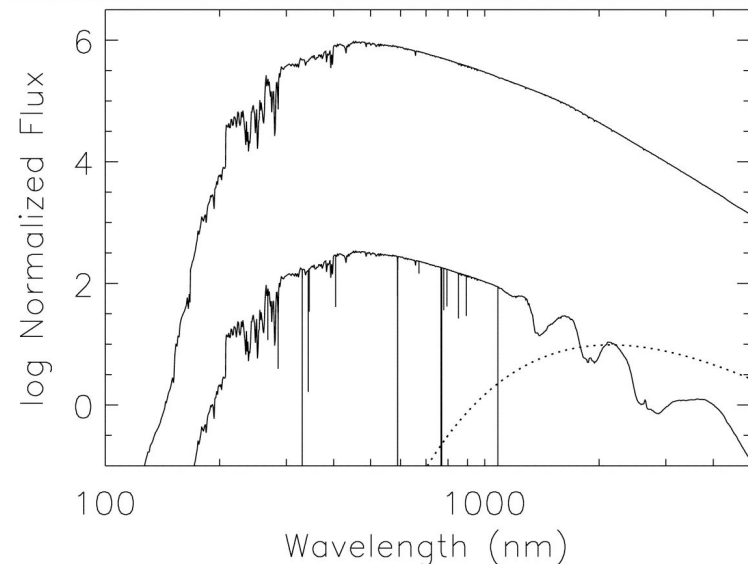


# Transmission spectroscopy

- During transit, light from the star passes through the planet's atmosphere
- Some wavelengths absorbed
- Smaller flux at these wavelengths – deeper transit
- Absorption lines superimposed on stellar spectrum
- Sodium a strong absorber



# Atmospheric Model

## **SWARM**

- Code written in IDL
- Takes parameters from MCMC file
- Calculates position of planet

## **Sodium**

- Calculates orbital radial velocity of planet
- Calculates decrease in flux due to transit

## **WASP**

- Uses synthetic spectra to generate model transit spectrum

## **Atmospheric**

## **Rossiter**

- Calculates radial velocity of model transit spectrum, including RM effect

## **Model**

- Will incorporate model atmosphere data
  - Observations compared to model
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