ASTROMETRIC MICROLENSING WITH GAIA

Łukasz Wyrzykowski

(pron: Woo-cash Vi-zhi-kov-ski) Warsaw University Astronomical Observatory, Poland



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COLLABORATORS



Krzysztof Rybicki (PhD student)



here

Kasia Kruszyńska (PhD student)



Mariusz Gromadzki (postdoc)



Zuzanna Kostrzewa-Rutkowska (postdoc at SRON, NL)

> Alex Hamanowicz Master student -> PhD @ ESO)

Gaia Alerts team in Cambridge (UK)









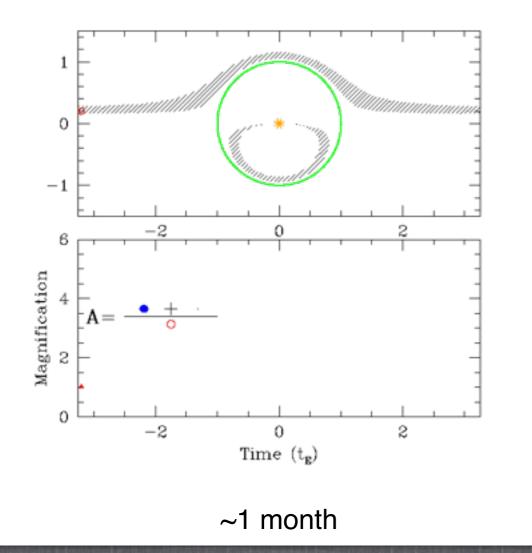
MICROLENSING

Gravitational lensing by compact lenses (stellar or remnants)

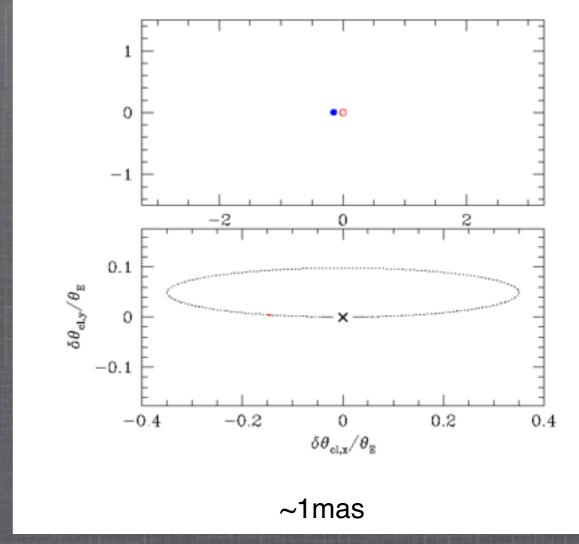
Mass range from Earth-like to ~100 MSun

• Sources: background stars (chance: 10⁻⁶ in the Bulge, 10⁻⁸ in the LMC)

photometry (sum of images)



astrometry (centroid motion)



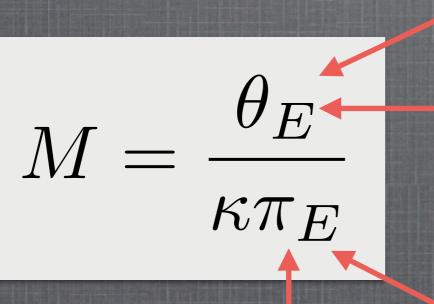
animations by S.Gaudi

BLACK HOLES MICROLENSING

 About 0.8% of microlensing events should be due to Black Holes! (Gould 2000)

2000 events found every year -> ~16 black holes every year!!

• so, where are they?



high amplification events/finite source (rare)

astrometry (VLT/AO, HST, <u>Gaia</u>)

Earth parallax

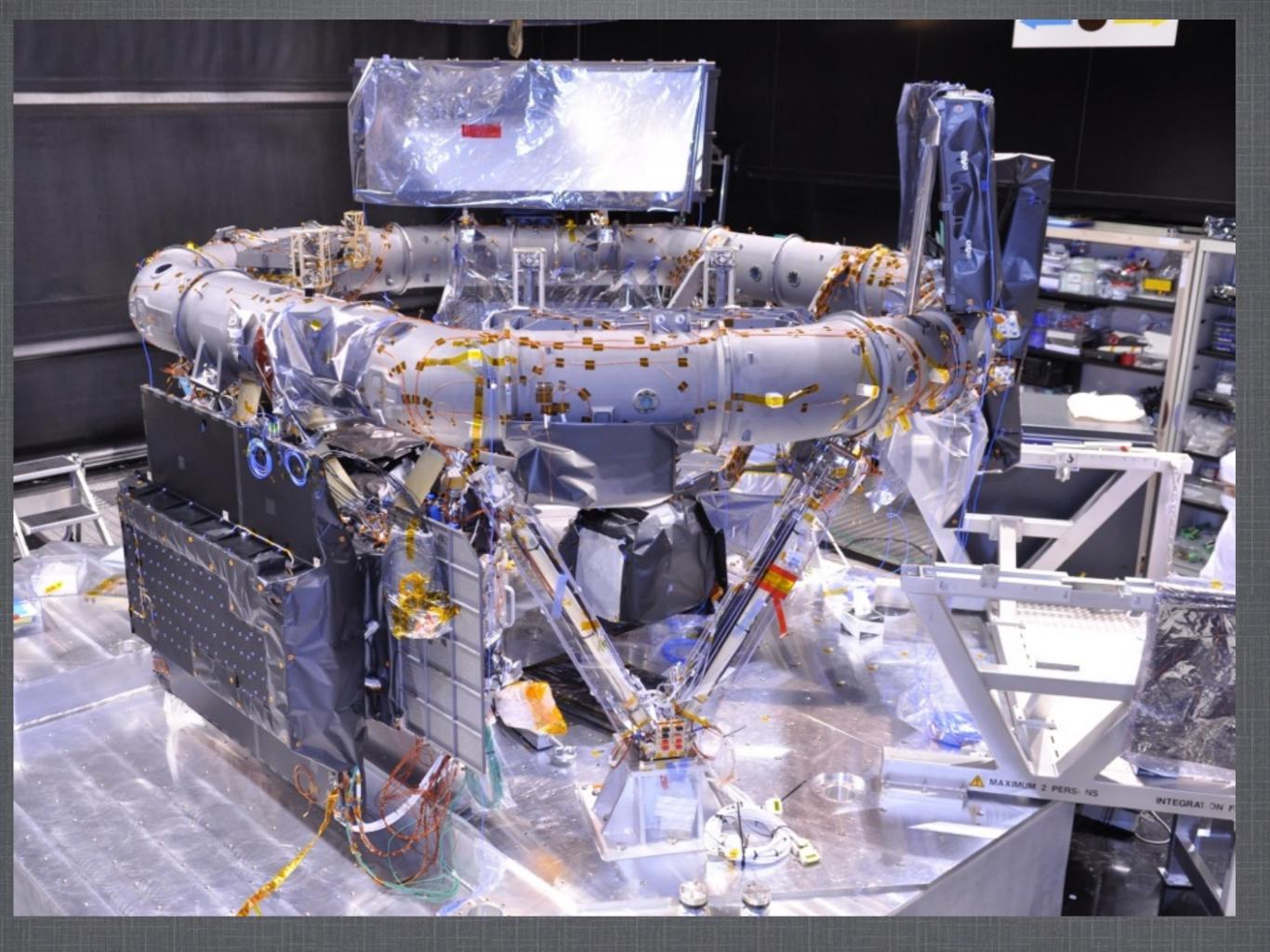
space-based parallax (e.g., Earth-Spitzer)

GAIA SPACE MISSION

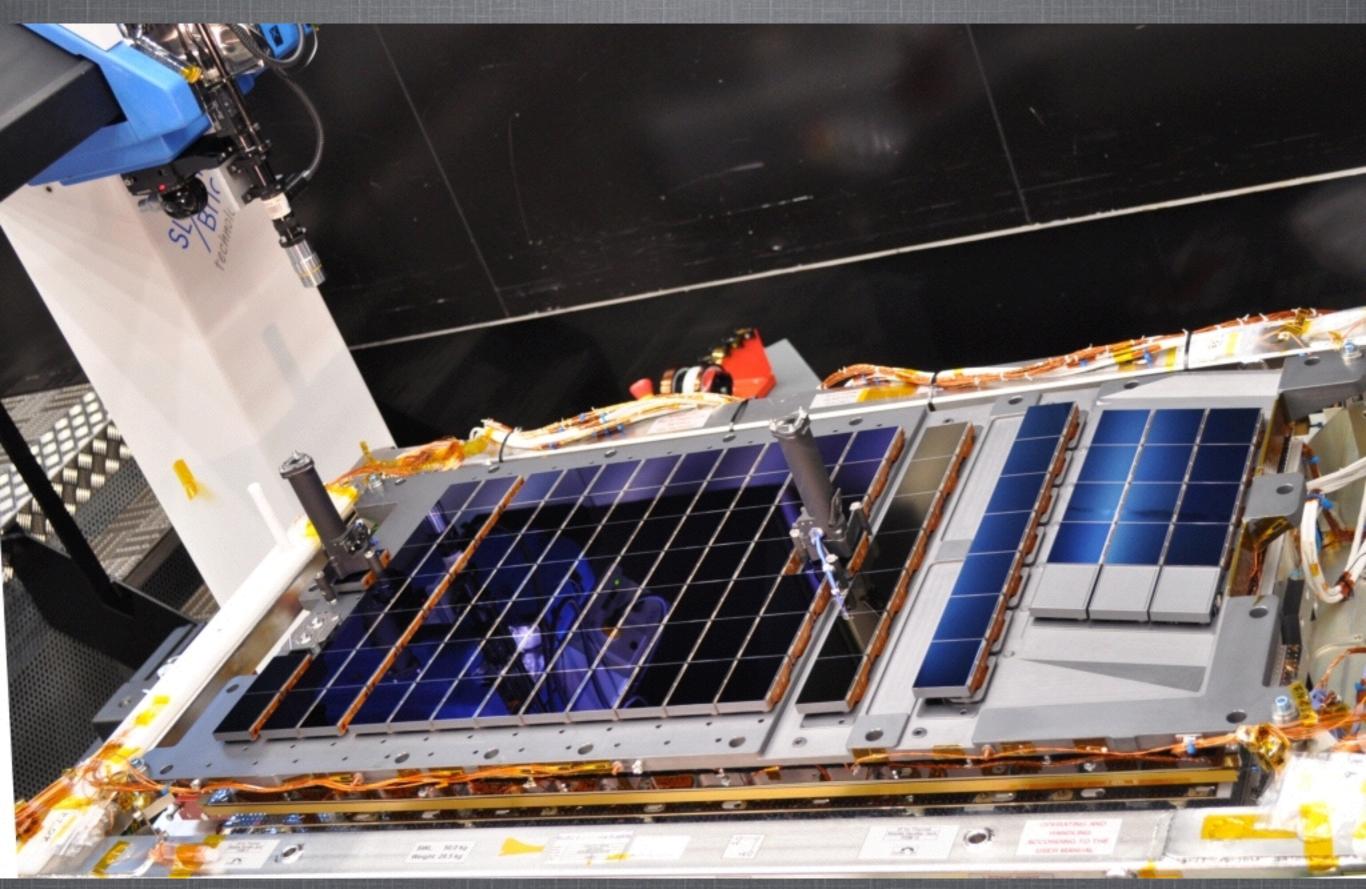
19 DEC 2013 9:12 UT



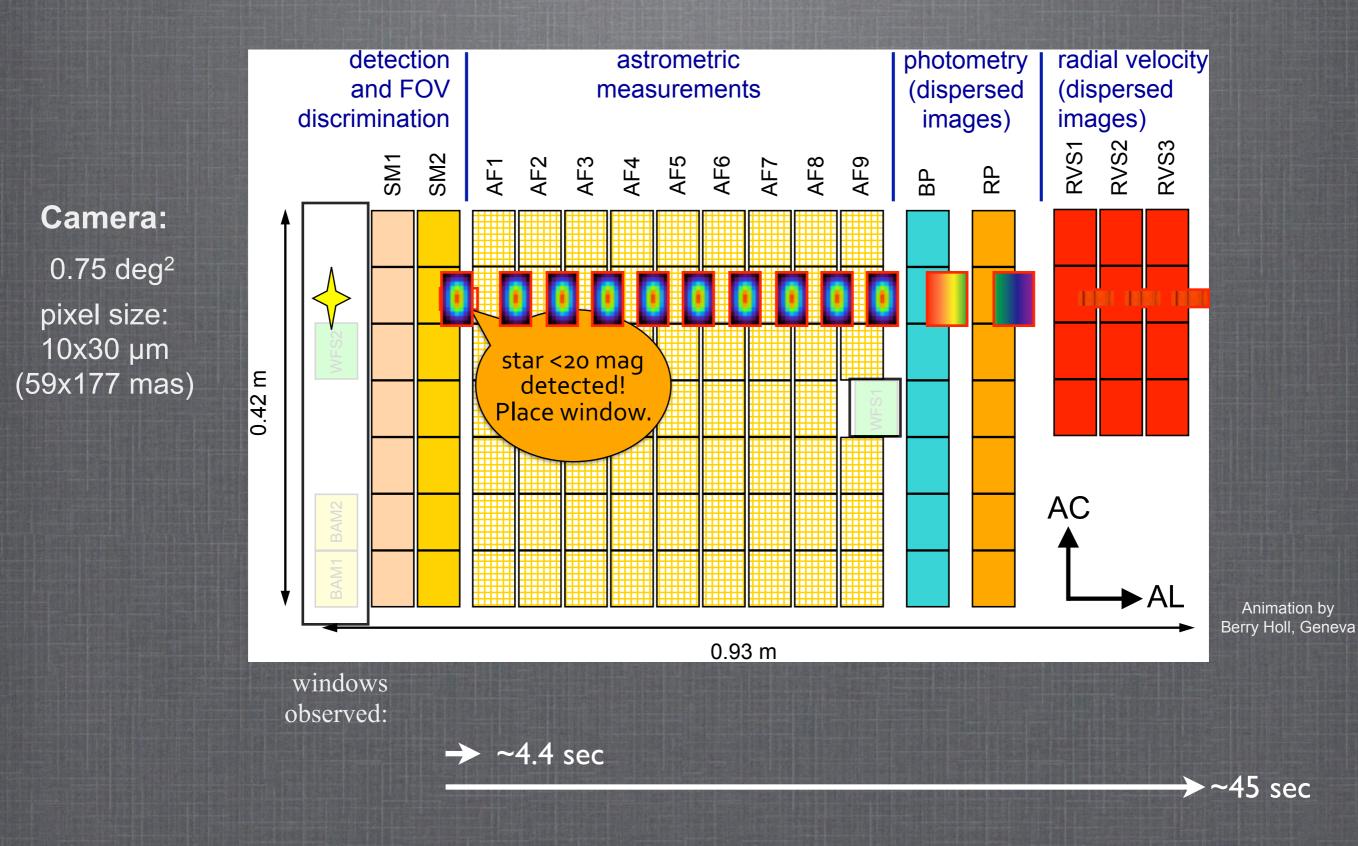




GAIA FOCAL PLANE

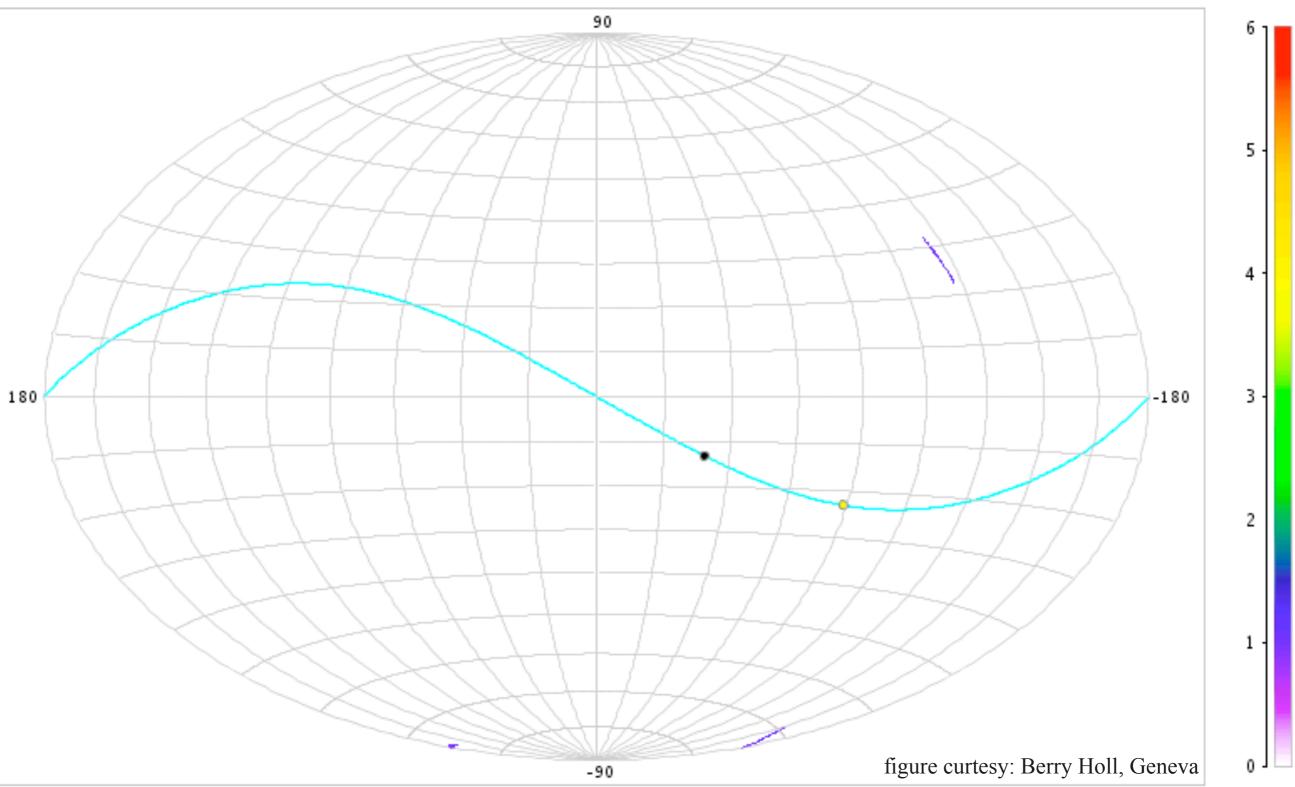


GAIA FOCAL PLANE



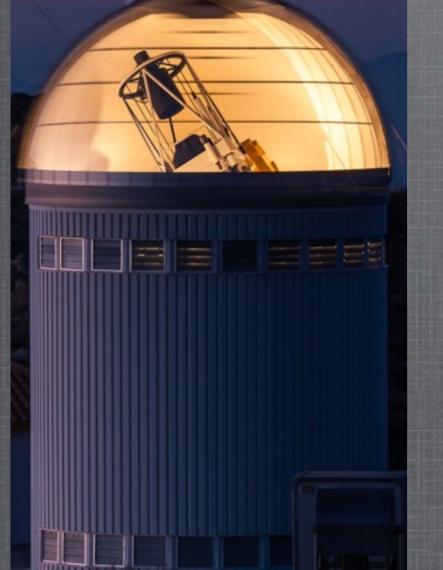
GAIA SKY SCANNING PATTERN

NSL field transits in ICRS after: 0 years 000 days 00 hr 10 min

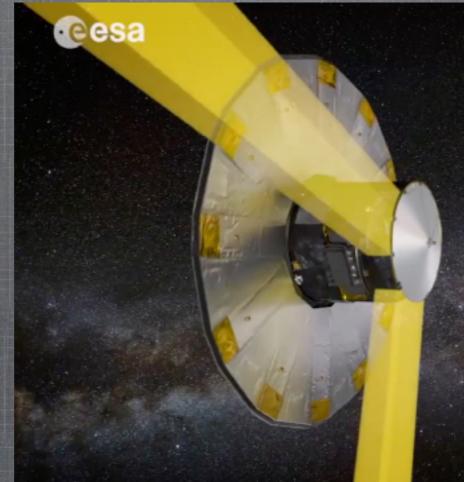


HOW TO FIND BLACK HOLES?

6GLF http://ogle.astrouw.edu.pl



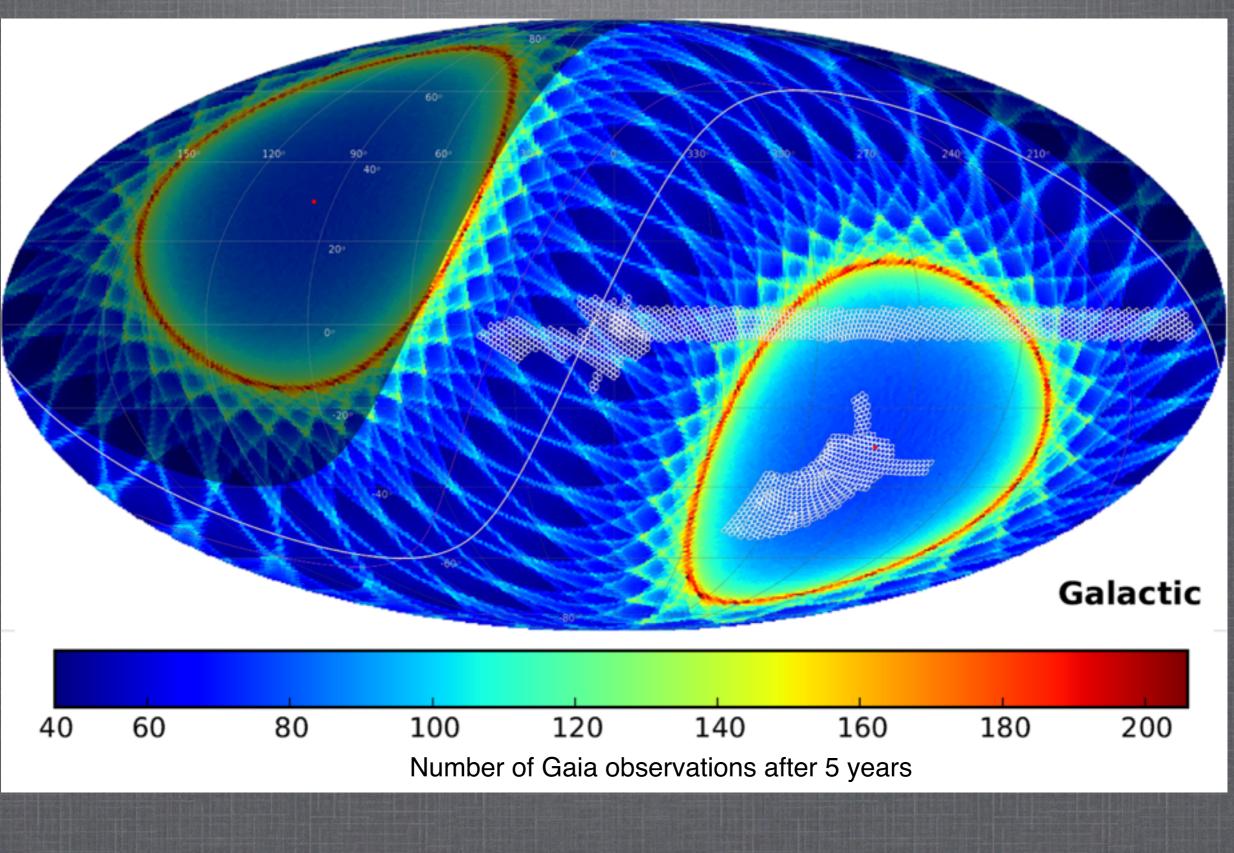
Gaia Science Alerts http://gsaweb.ast.cam.ac.uk/alerts





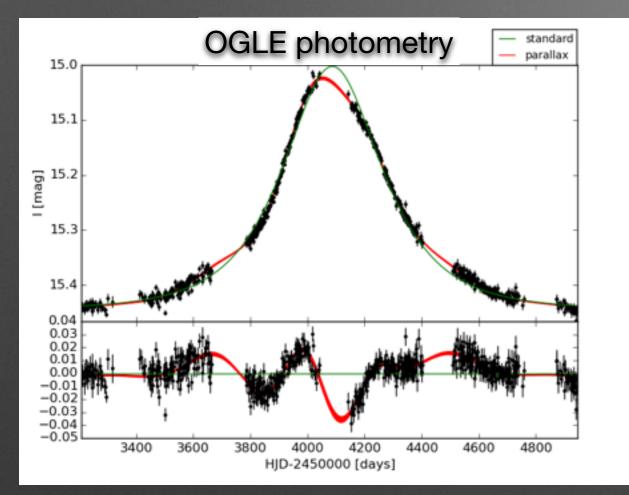
Polish 1.3m dedicated telescope in Las Campanas, Chile Surveying continuously since 1992. ESA space mission with 2x1.4m telescopes located in L2. In operation since 2014.

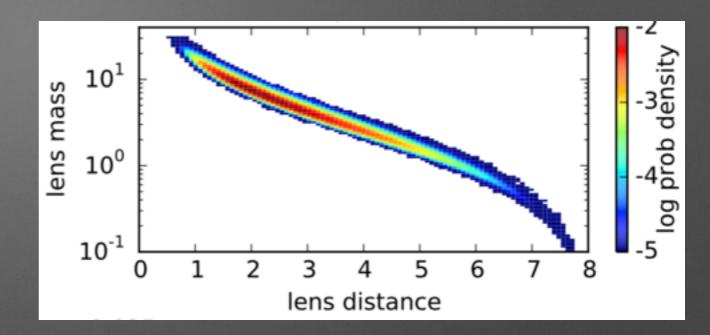
OGLE-GAIA SKY



STELLAR-MASS BLACK HOLES

OGLE3-ULENS-PAR-02 - candidate ~9MSun BH



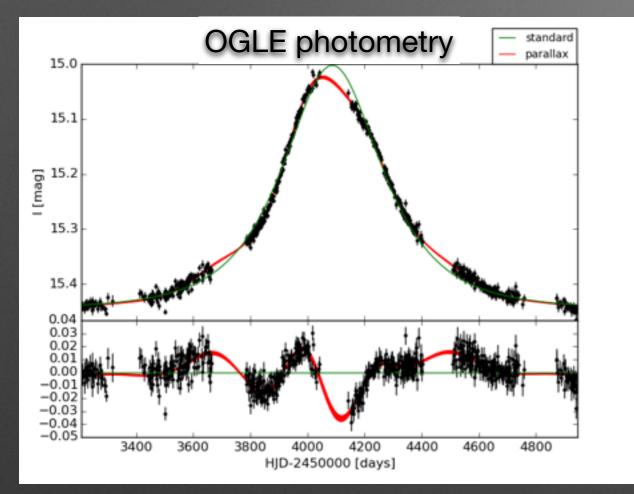


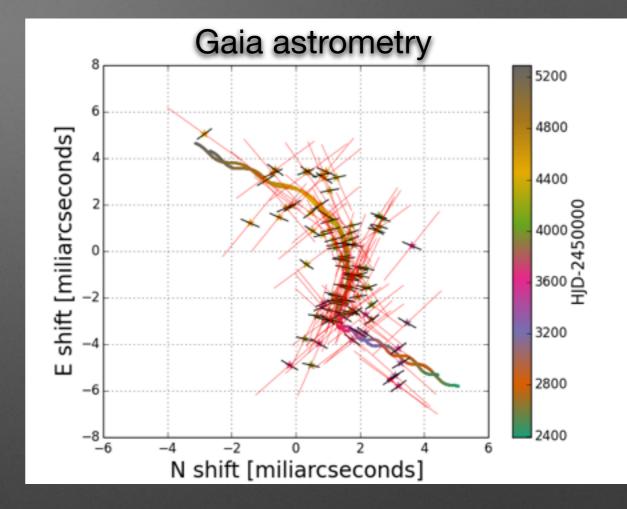
OGLE photometry from 2001-2008 and microlensing model probability distribution for the proper motion based on observed PM

Mass, Distance estimate only

STELLAR-MASS BLACK HOLES

OGLE3-ULENS-PAR-02 - candidate ~9MSun BH





OGLE photometry from 2001-2008 and microlensing model

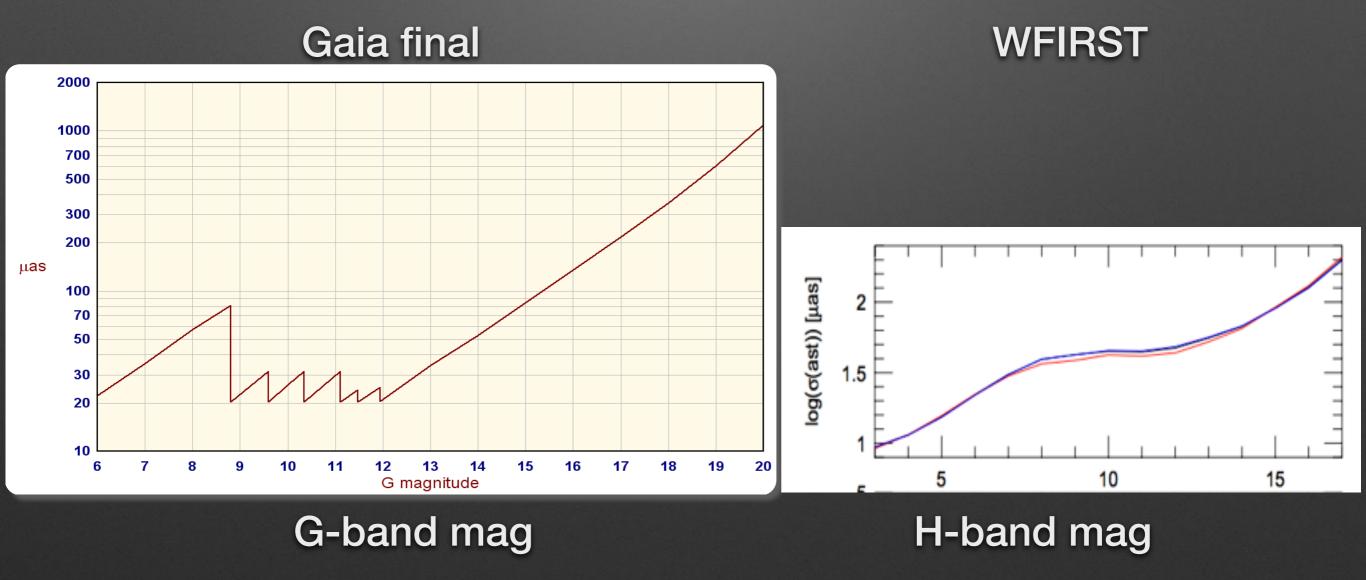
Mass, Distance

predicted Gaia astrometry for similar event

Rybicki in prep.

ASTROMETRIC PRECISION IN GAIA

50 mas : daily calibrations 1D precision! 0.1 mas : fortnightly calibrations final precision (in 2022):



Gould+2017

ASTROMETRIC PRECISION IN GAIA

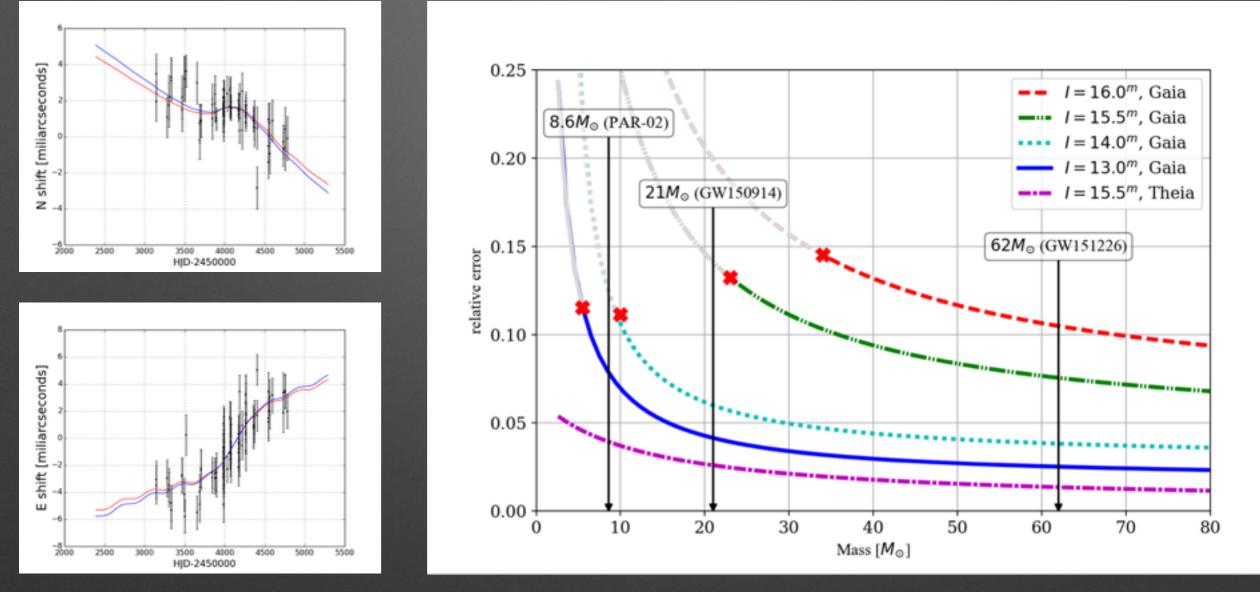
50 mas : daily calibrations 1D precision! 0.1 mas : fortnightly calibrations final precision (in 2022):



Gaia final

Gould+2017

STELLAR-MASS BLACK HOLES OGLE3-ULENS-PAR-02 - candidate ~9MSun BH



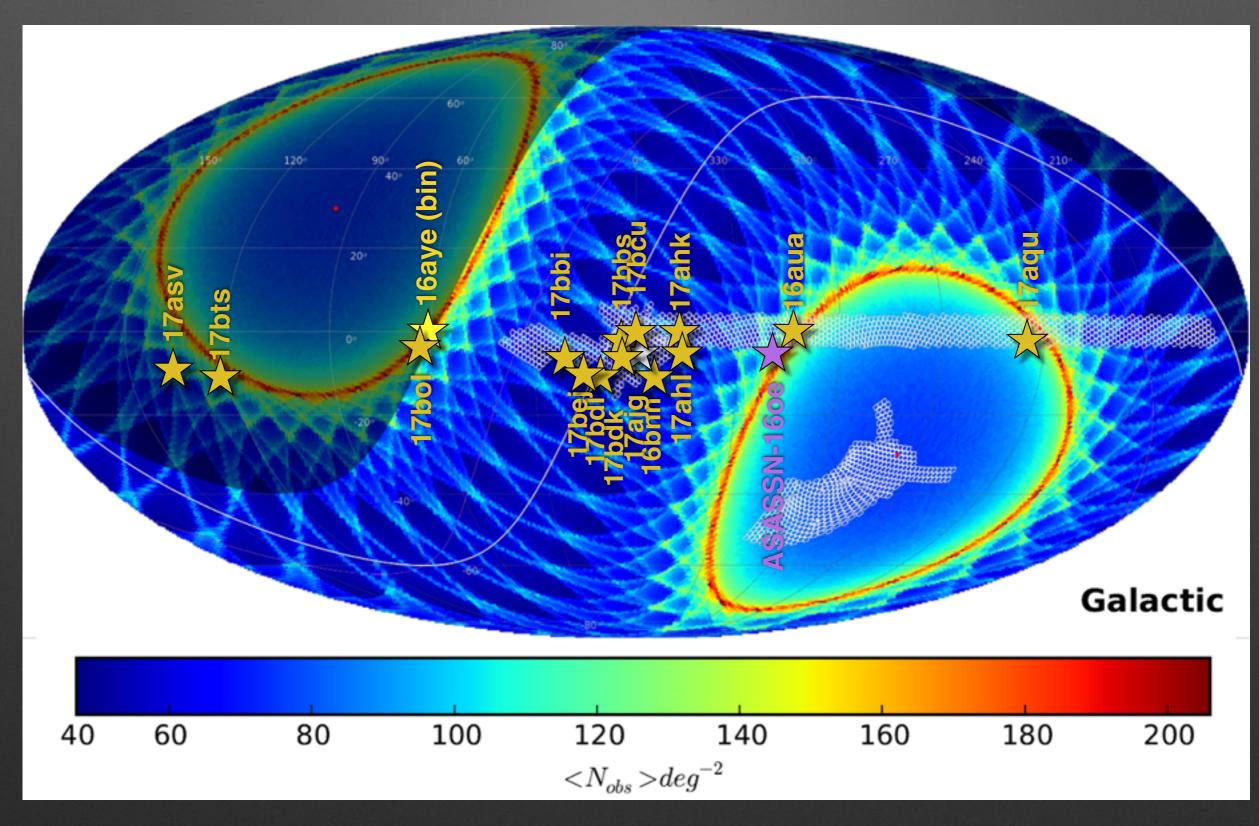
Combination of ground-based photometry and Gaia astrometry for long events will yield masses of black holes accurate to ~10% percent (brightness dependent)

see Kris Rybicki's POP

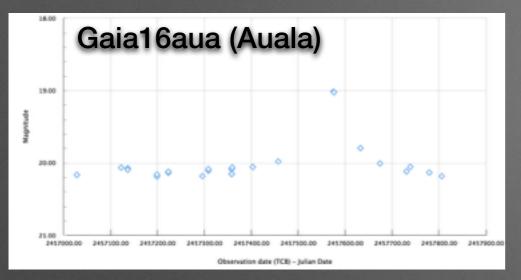
short break :)

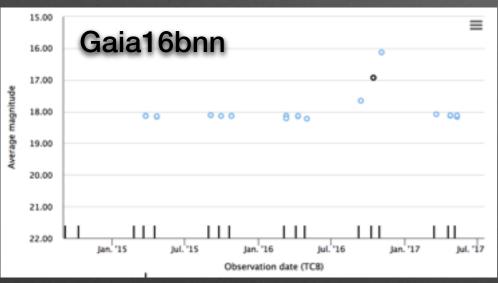


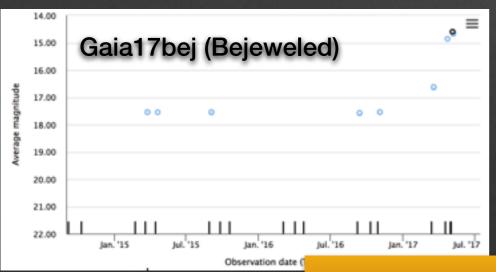
MICROLENSING EVENTS FROM GAIA July 2016 - July 2017

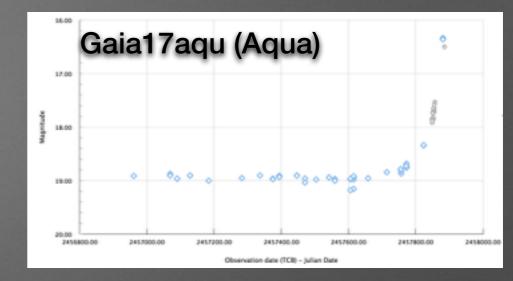


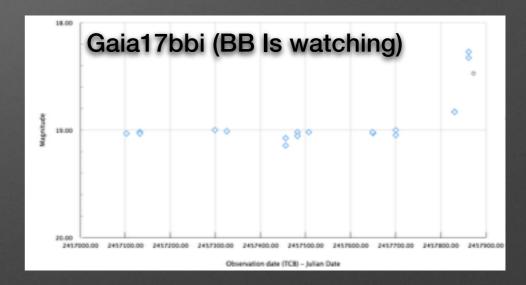
MICROLENSING EVENTS FROM GAIA

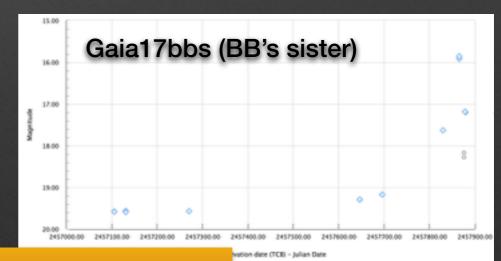








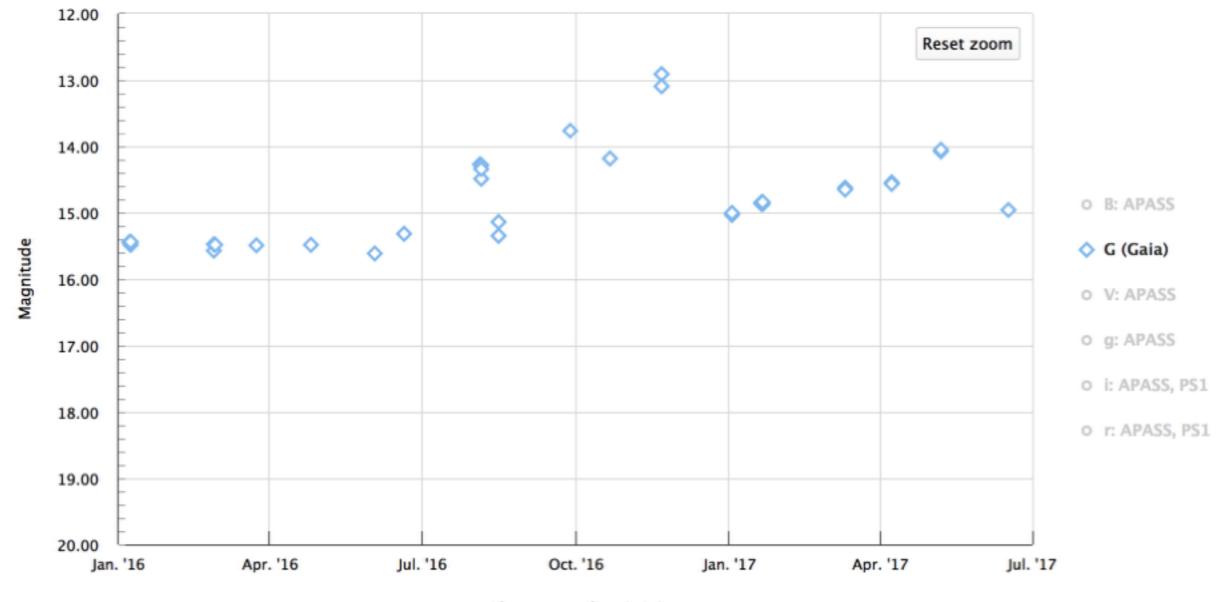




see Kasia Kruszynska's poster/POP

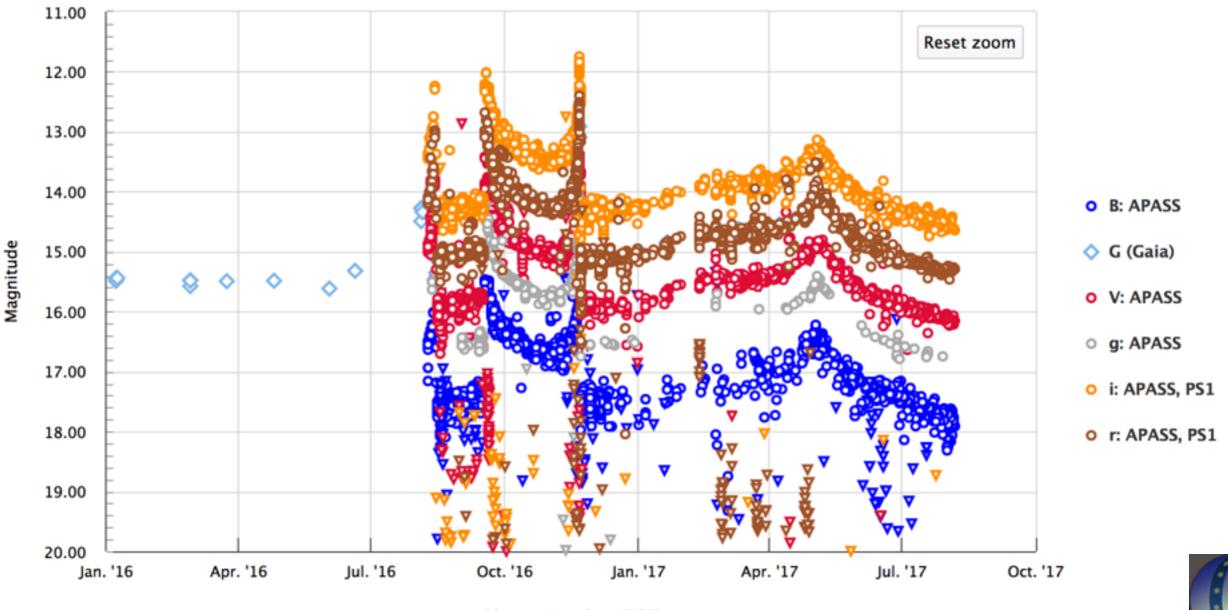
Gaia16AYErs Rock*





Observation date (TCB)

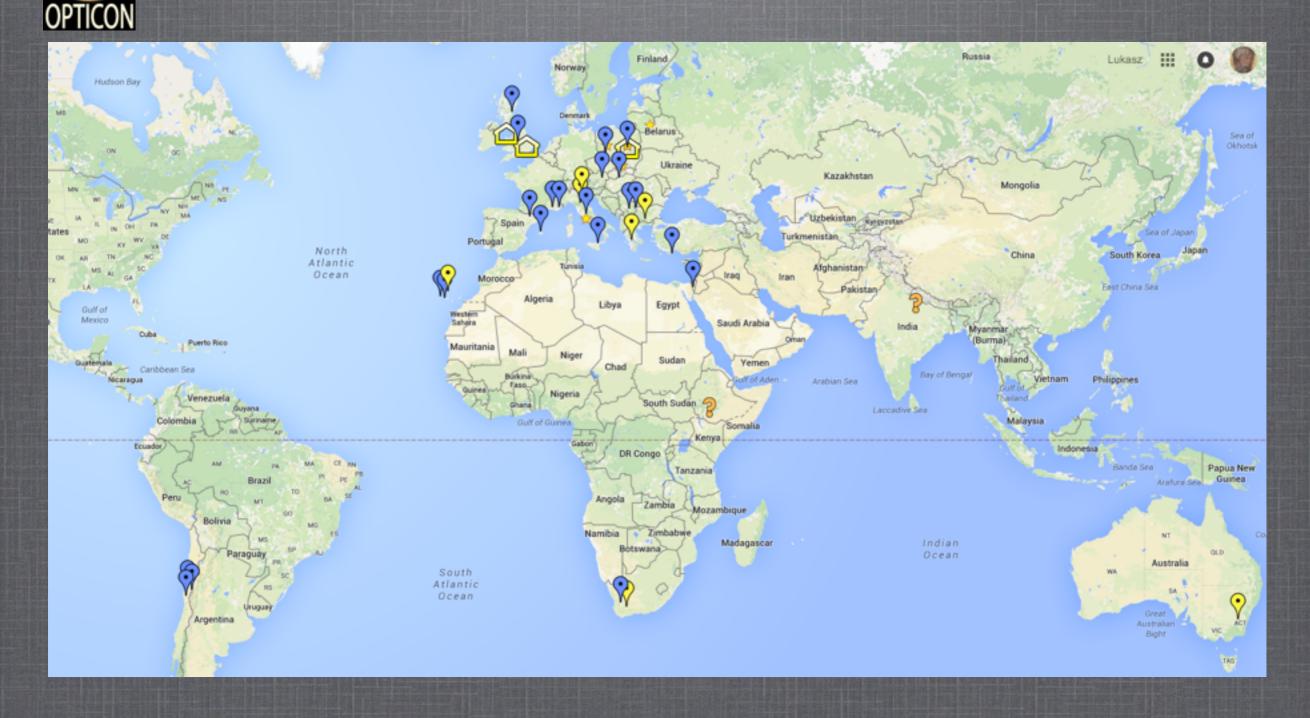
More than 20,000 data points collected in multiple bands by a network of volunteering observatories



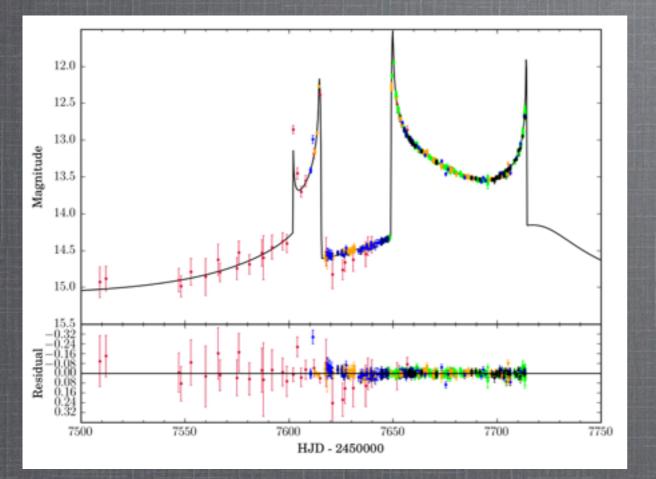
Observation date (TCB)

GAIA ALERTS FOLLOW-UP NETWORK

~20 active partners, ~30000 data points collected 2014-2016 2017-2020: continuation under OPTICON H2020

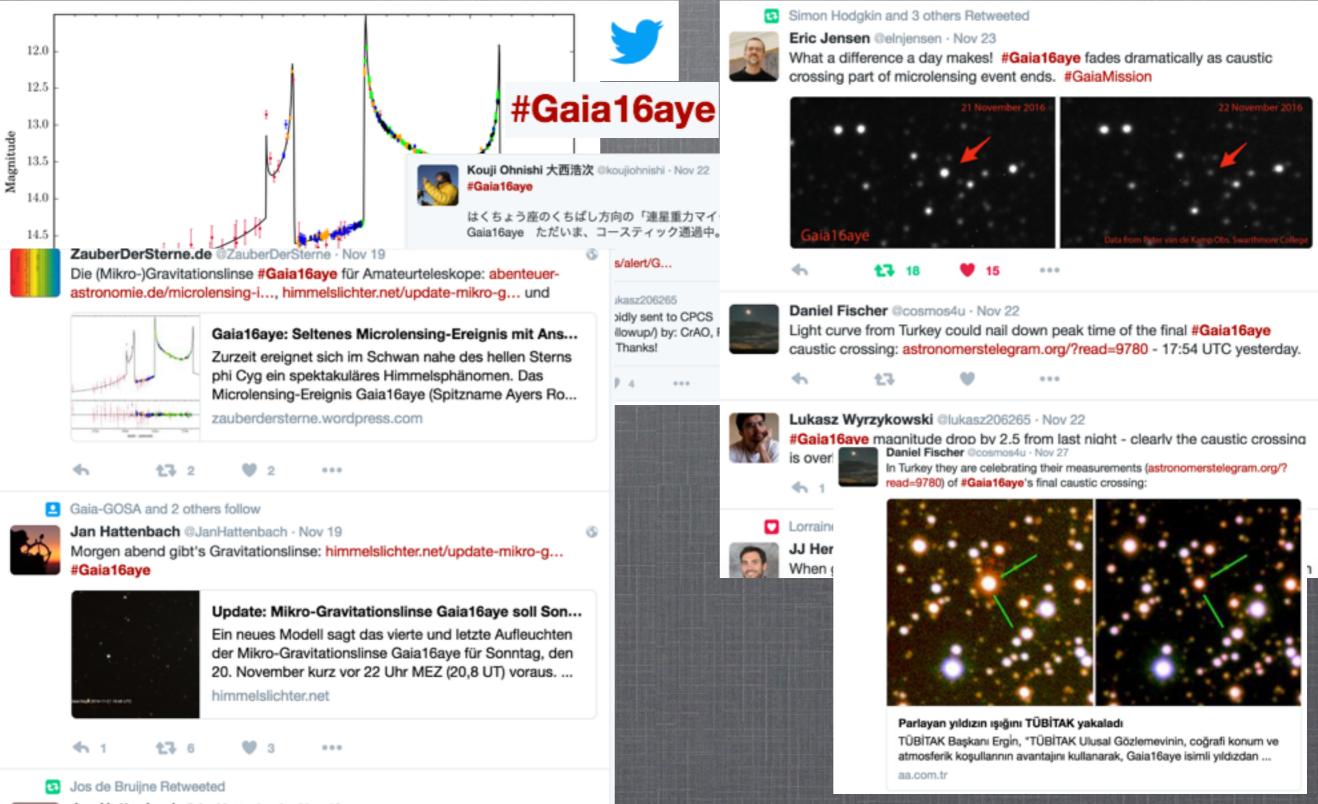


Catching the 4th caustic exit



Model prediction Caustic exit 21 November 2 am UT

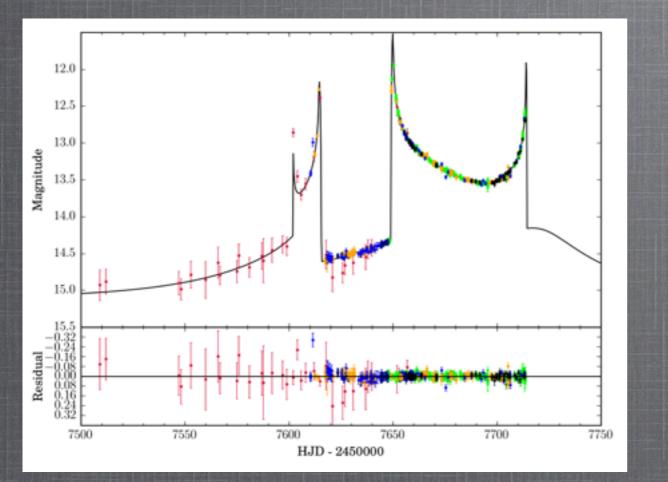
Catching the 4th caustic exit



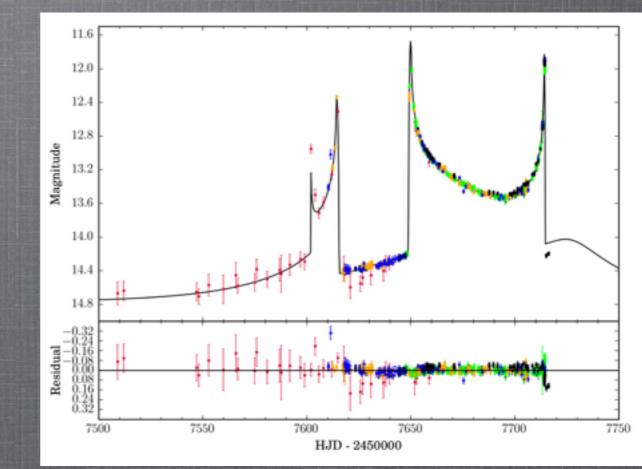
Ref.

Jan Hattenbach @JanHattenbach - Nov 18

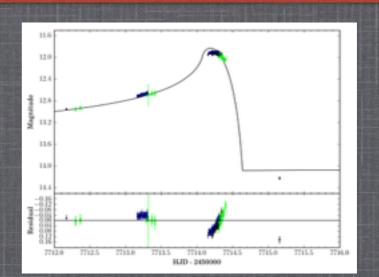
Catching the 4th caustic exit



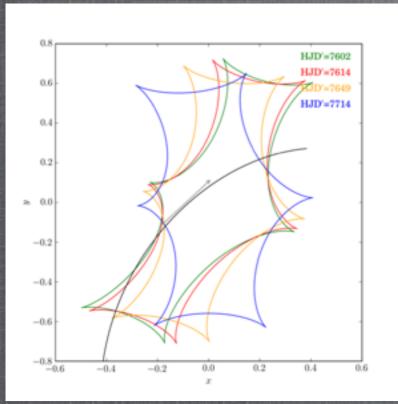


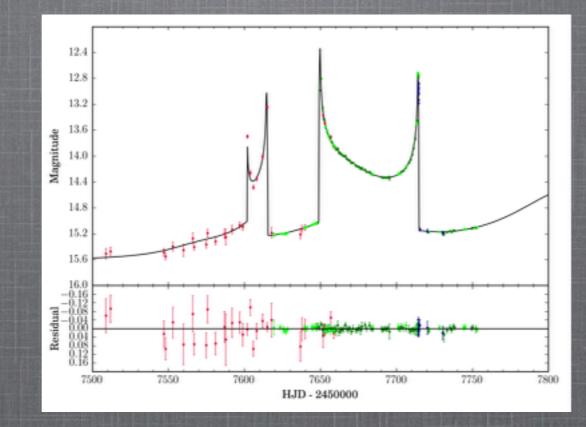


actual peak: 21 Nov ~16 UT

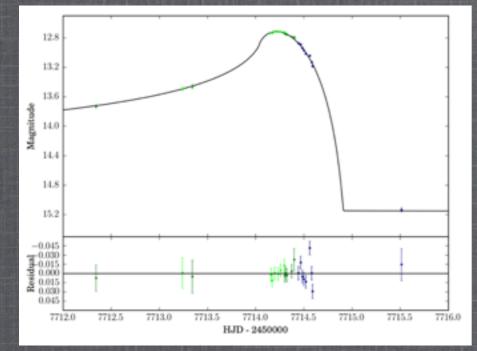


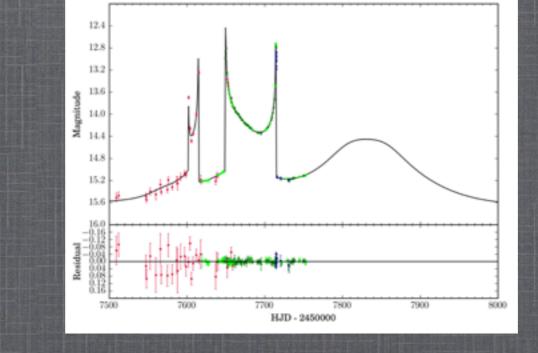
Preliminary full-Keplerian orbital solution of the binary lens





tE = 141d piE= 0.39 thetaE=3 mas $mu_rel=7mas/yr$ q=0.57 s=1.0fs=0.75





model by P.Mroz and J.Skowron

PRELIMINARY SOLUTION

~8kpc

2.3AU

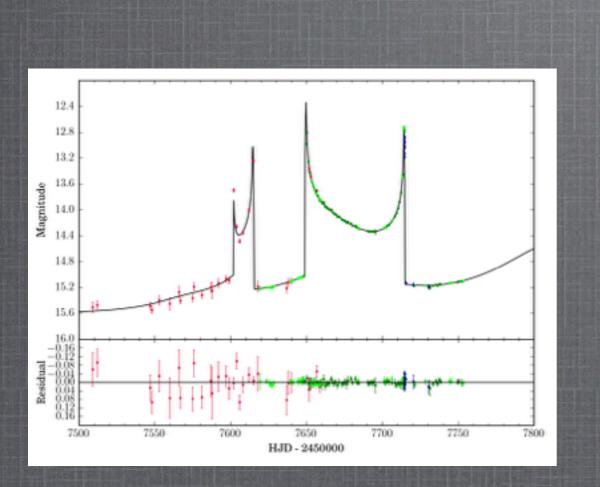


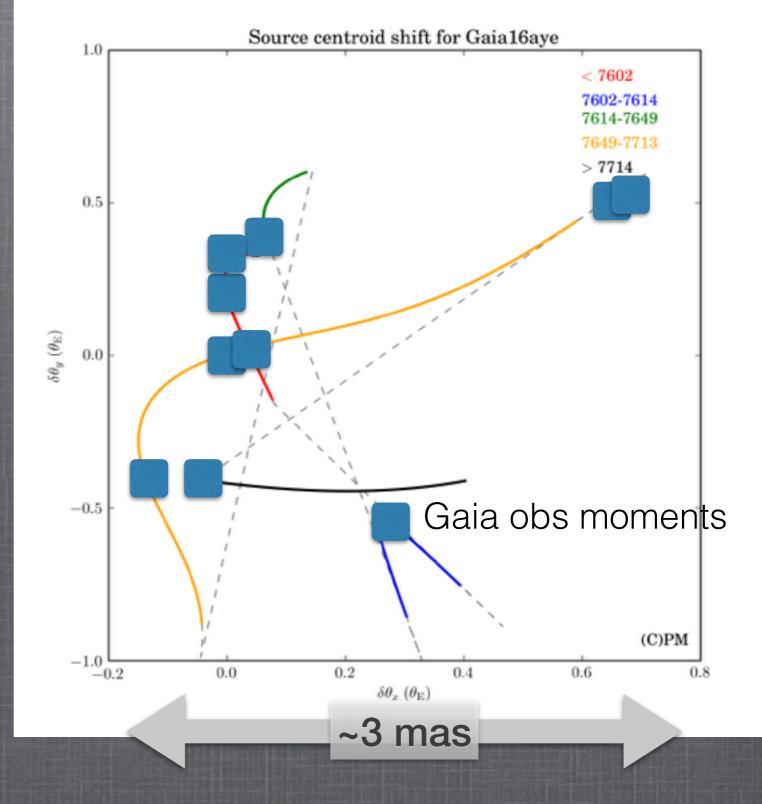


 $M_1 = 0.4 M_{Sun}$ $M_2 = 0.6 M_{Sun}$ P = 3.4 yrsincl = 60 deg ecc = 0.473 K3 giant R=~10 R_{Sun}

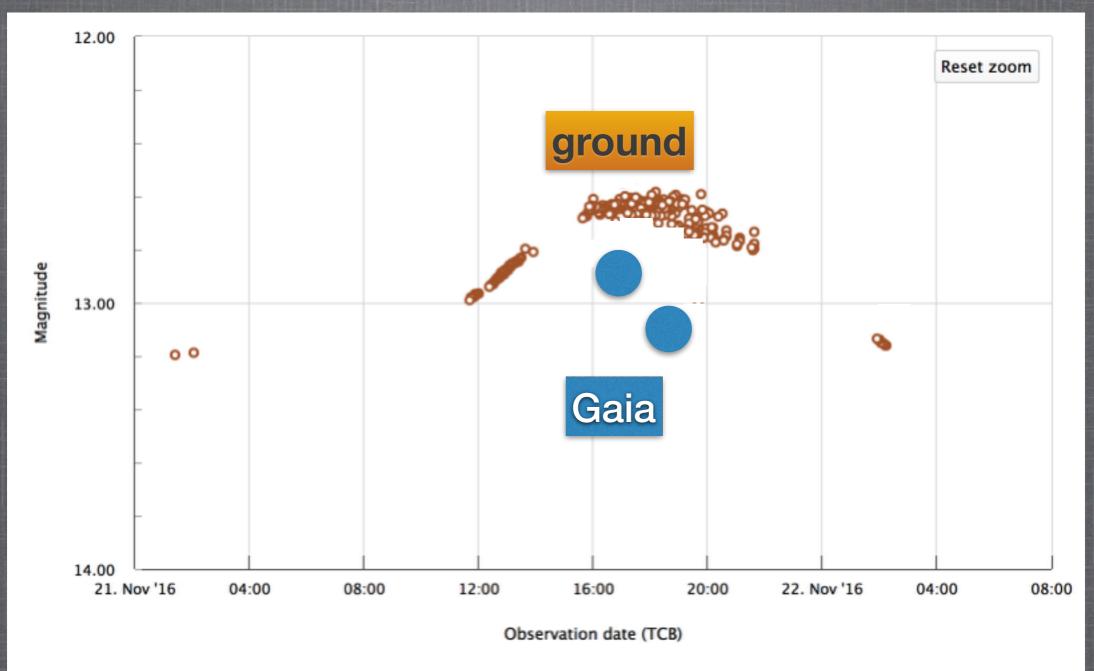
ASTROMETRY

First time ever chance to detect binary astrometric microlensing!





SPACE PARALLAX FROM GAIA?

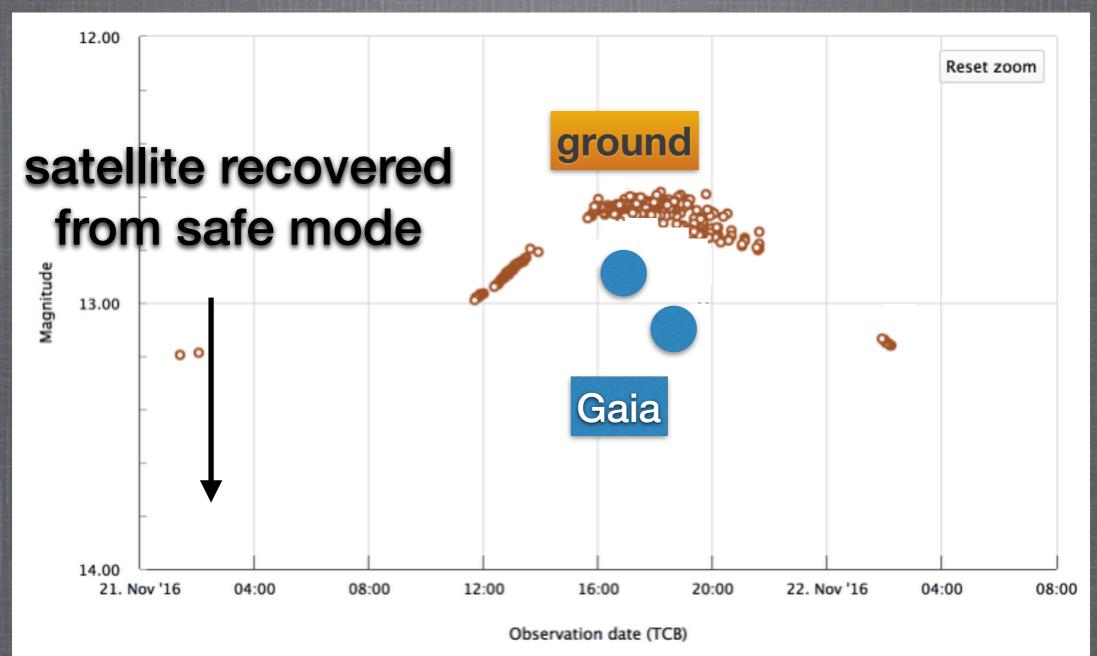


$$\boldsymbol{\pi}_{\mathrm{E}} = \frac{\mathrm{AU}}{D_{\perp}} \left(\frac{\Delta t_0}{t_{\mathrm{E}}}, \Delta u_0 \right),$$

Offset in both time and mag between Earth-based and Gaia-based observations

microlensing parallax

SPACE PARALLAX FROM GAIA?



$$\boldsymbol{\pi}_{\mathrm{E}} = \frac{\mathrm{AU}}{D_{\perp}} \left(\frac{\Delta t_0}{t_{\mathrm{E}}}, \Delta u_0 \right),$$

Offset in both time and mag between Earth-based and Gaia-based observations microlensing parallax



GAIA DATA RELEASES



DR1 14 September 2016:

- positions for 1 billion stars
- astrometric solution for 2 million stars (TGAS) (proper motion, parallax)

DR2 April 2018:

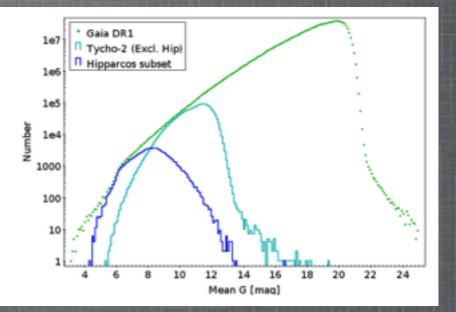
- astrometric solution for 1 biliion stars
- colors, temperatures, radial velocity

DR3 2020:

- improved astrometry
- low-resolution spectra for all objects
- radial velocity, variable stars
- non-single stars

DR4 (Final) 2022:

- Full astrometric, photometric and radial velocity catalogs
- All per-epoch astrometry and photometry
- Exo-planets

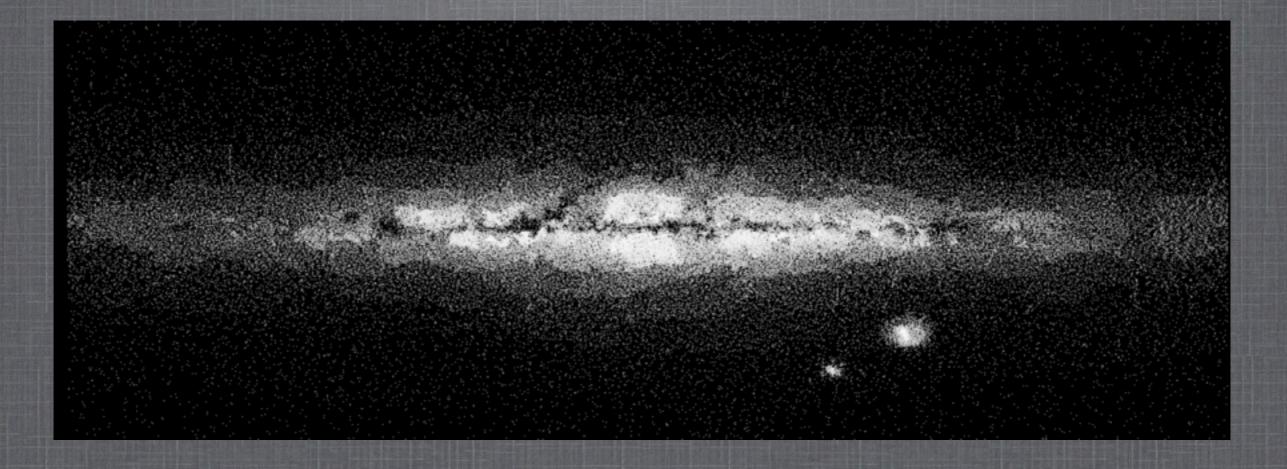




gaia archive



https://gea.esac.esa.int/archive/



SUMMARY

- Gaia will soon provide 3D map of the Milky Way
- all data public in 2022 (after your PhD!)
- Gaia sampling in the Bulge is too poor for microlensing
- but Gaia finds microlensing in the MW Disk
- astrometric microlensing from Gaia only for brighter events
- space parallax possible Earth Gaia (in L2)
- WFIRST should have astrometric microlensing capability — great for measuring theta_E for black hole lenses!