

How Gaia Grounds Kepler, K2, TESS, and Direct Imaging

Jessie Christiansen Caltech/IPAC-NASA Exoplanet Science Institute Sagan Summer Workshop 2022

<u>christia@ipac.caltech.edu</u> 🔎 @aussiastronomer



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Overview

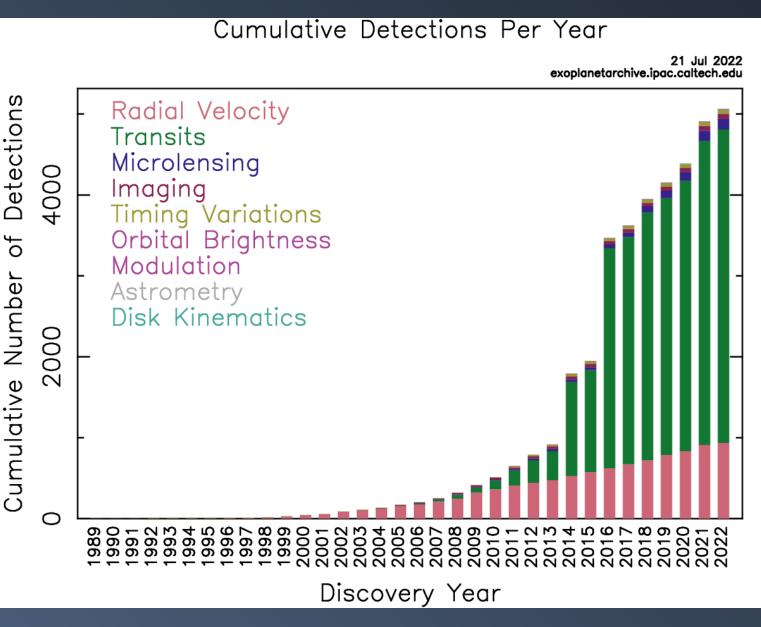


- 1. Planets! (Not from astrometry)
- 2. Improving our understanding of planets and planet populations with Gaia
- 3. Future open questions that Gaia will help tackle

Overview



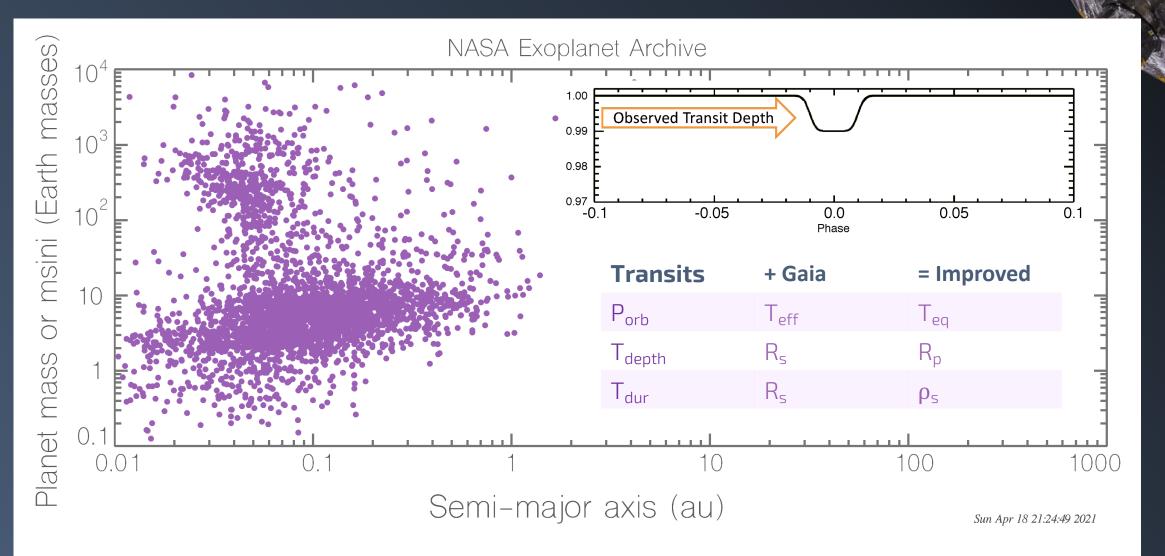
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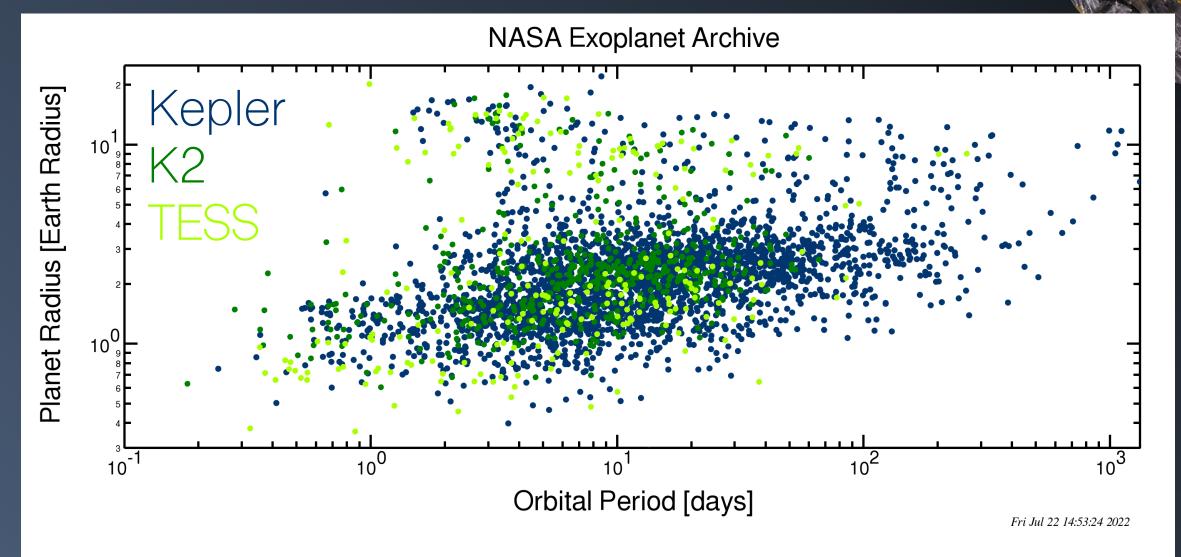


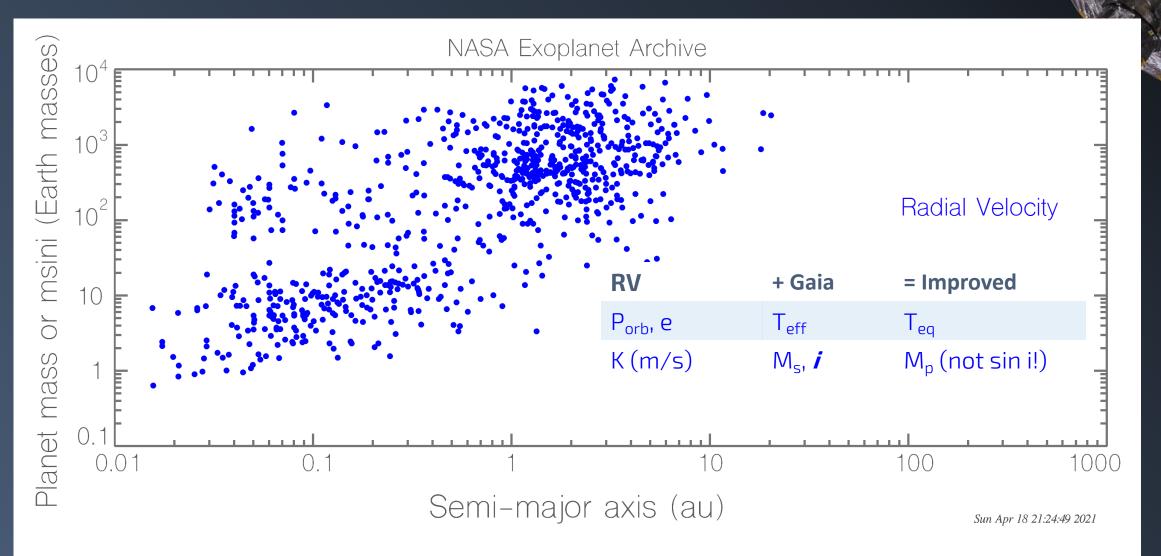


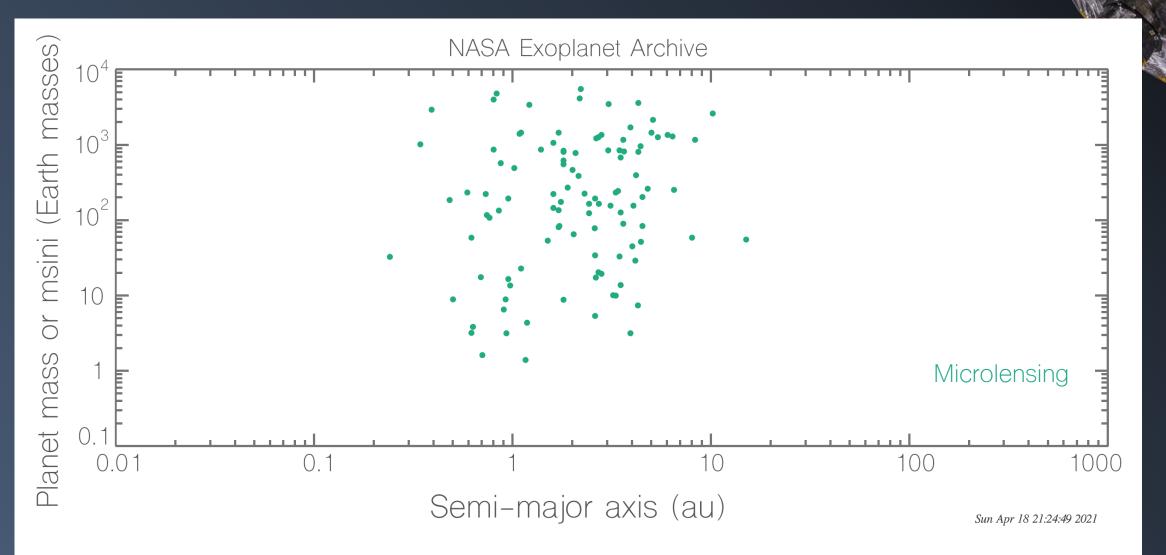
Transit: 3875 (inc 2 from Gaia! RV: 936 Microlensing: 130 Imaging: 61

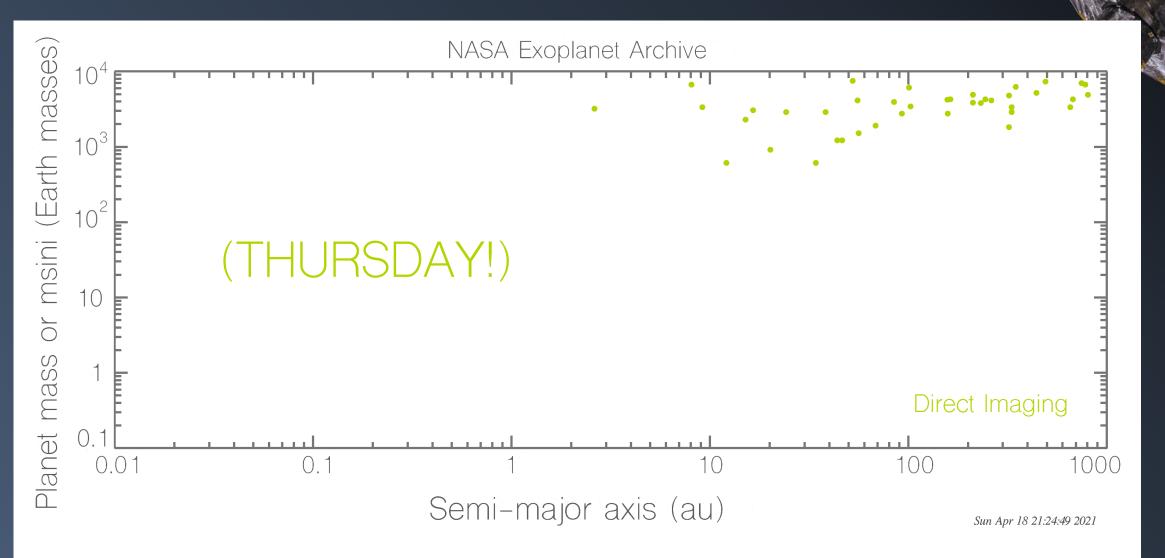
(Astrometry: 20,000?)

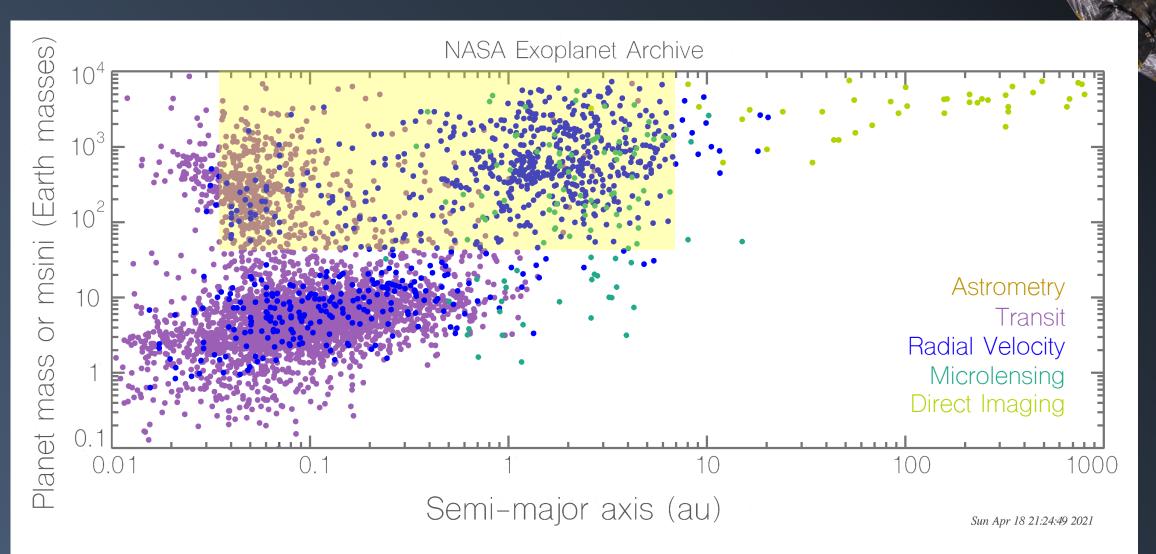












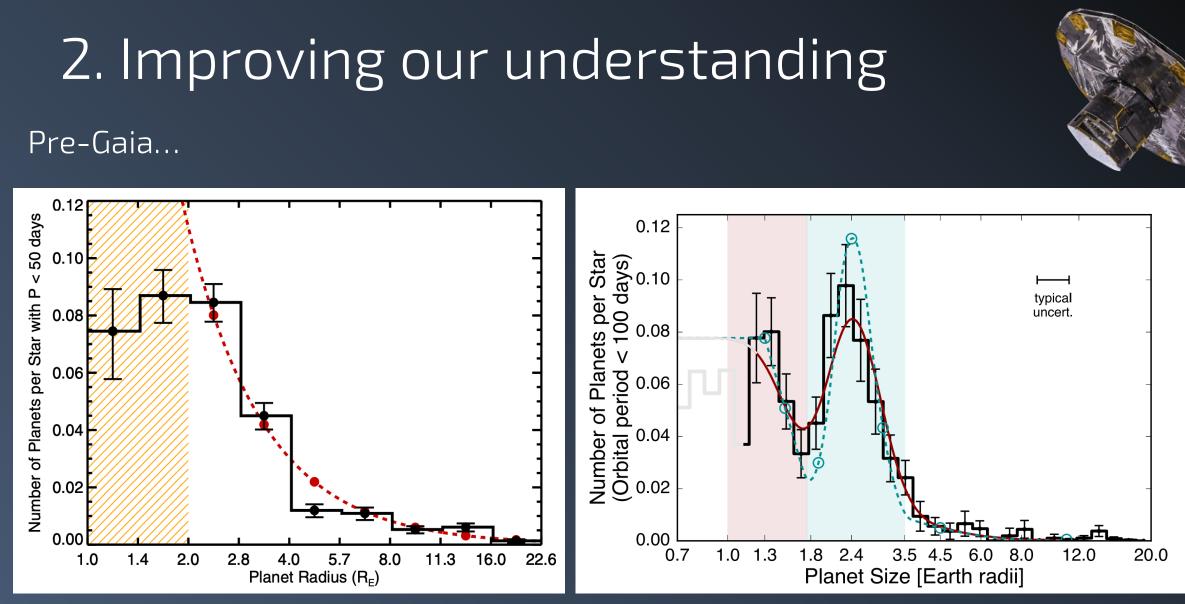
Overview



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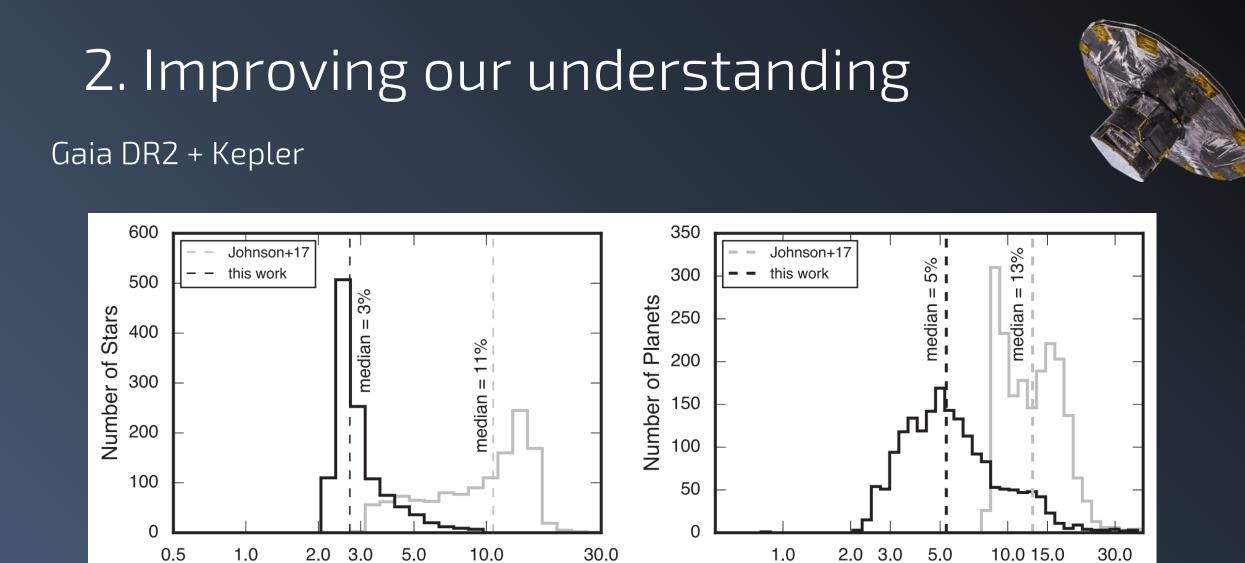
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Howard+2012

Fulton+2017

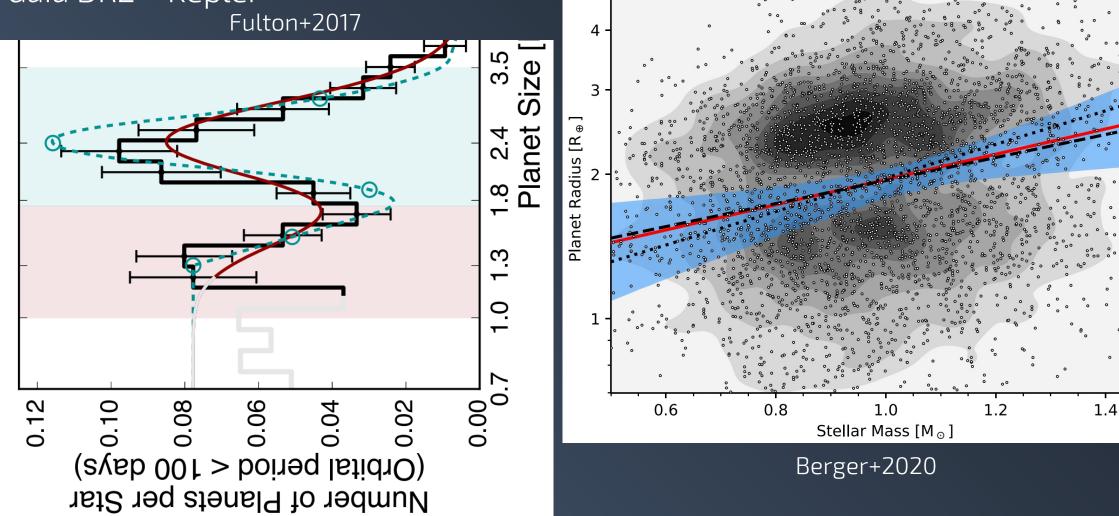


Fulton+2018

Fractional Planet Radius Uncertainty [%]

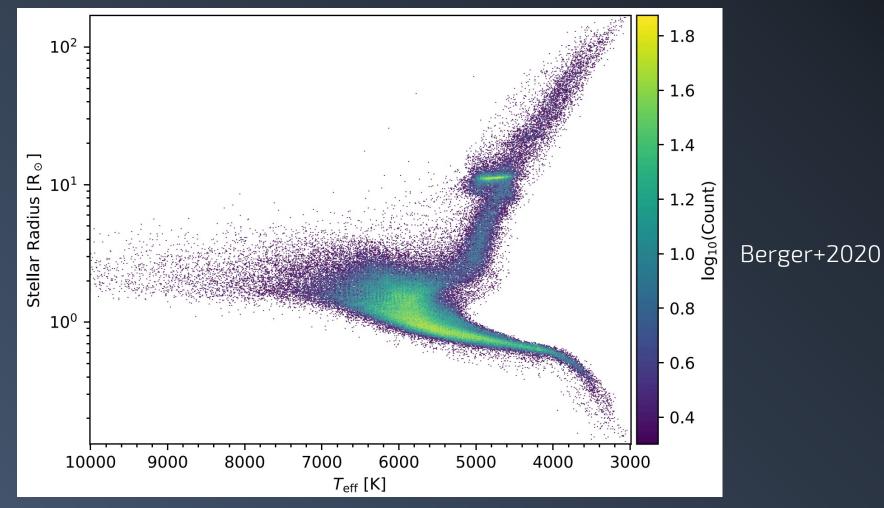
Fractional Stellar Radius Uncertainty [%]

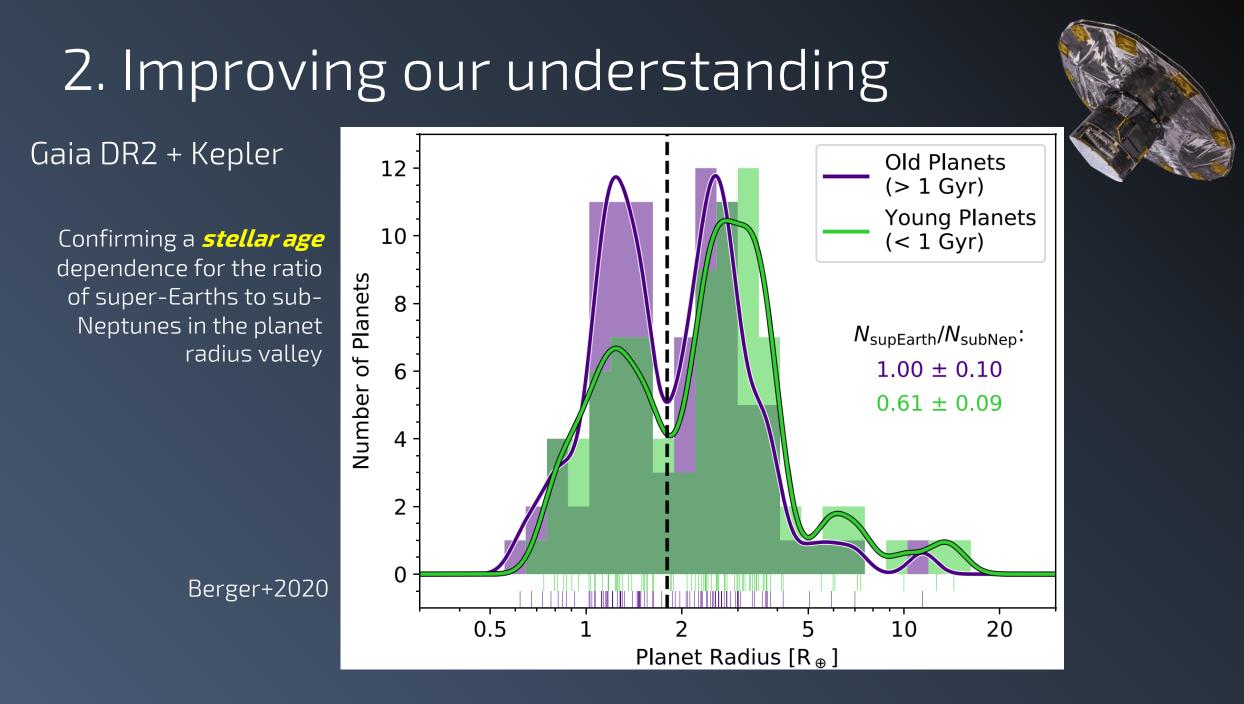
Gaia DR2 + Kepler

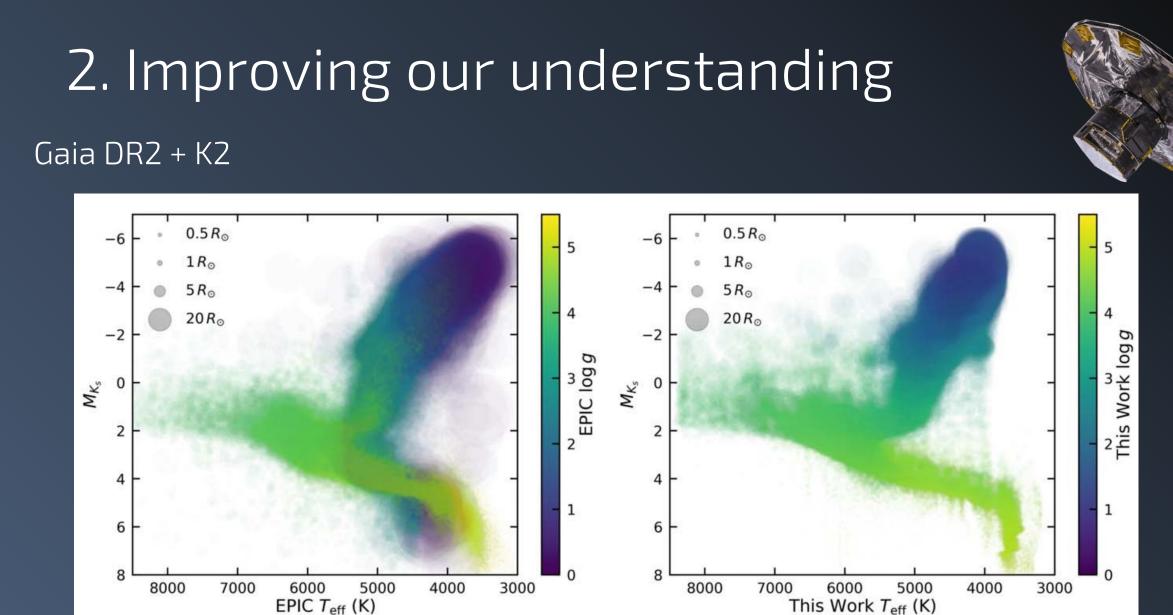




Gaia DR2 + Kepler





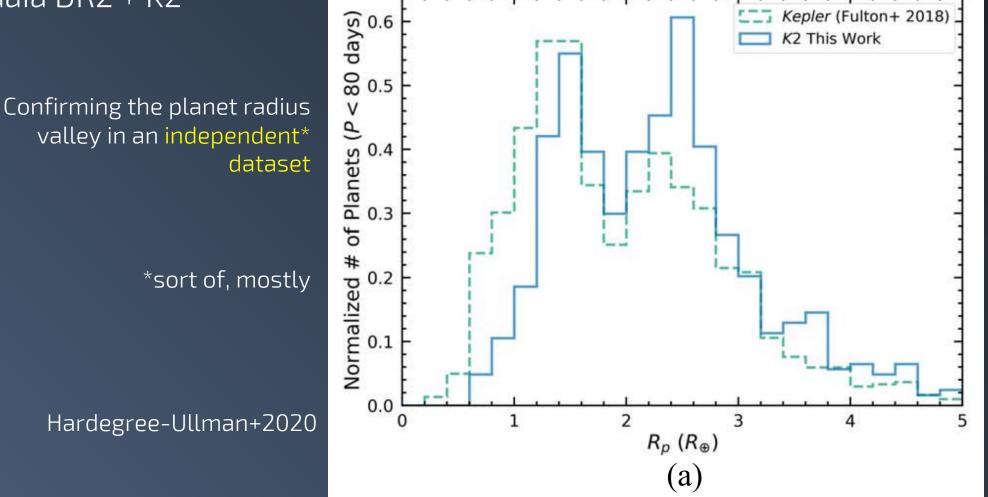


Hardegree-Ullman+2020

(b)

(a)

Gaia DR2 + K2





Gaia + TESS

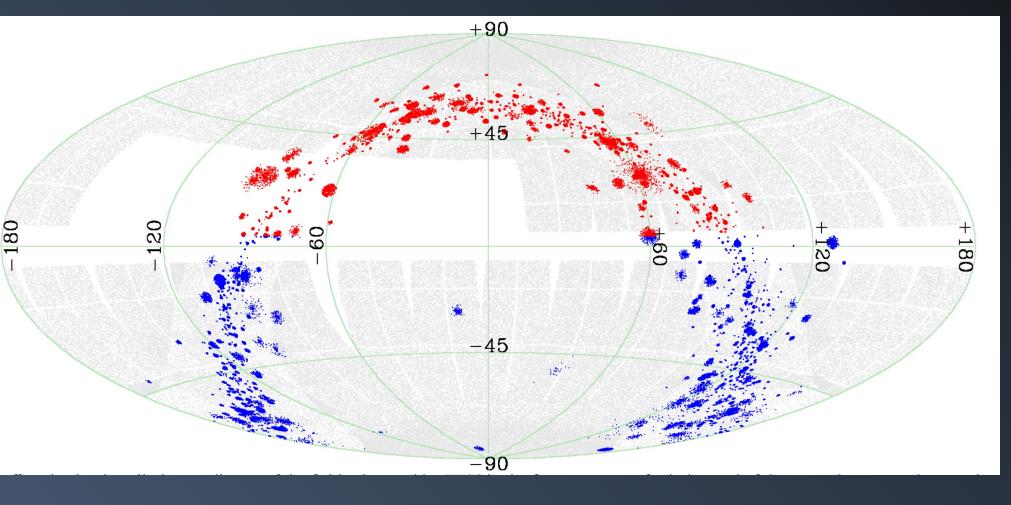
(TESS + Gaia = stellar astrophysics, solar system/asteroids, etc etc!)

Star clusters – astronomy's laboratory!

Age Chemistry Dynamical history

Constraining known clusters

Finding new clusters!



Nardiello+2020





Which name of a new star cluster found by Gaia is worse?

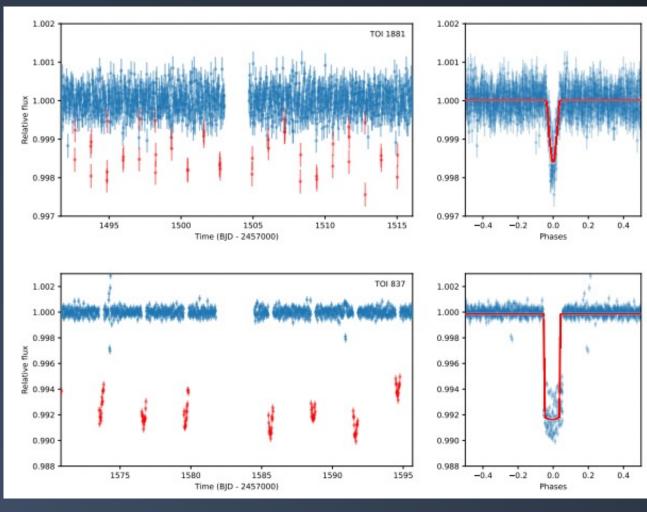
Gaia-Enceladus Sausage
Group X

Gaia DR2 + TESS

Identifying planet candidates in open clusters

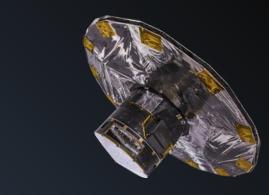
Gaia: positions and velocities

TESS: rotation periods





Sun+2022

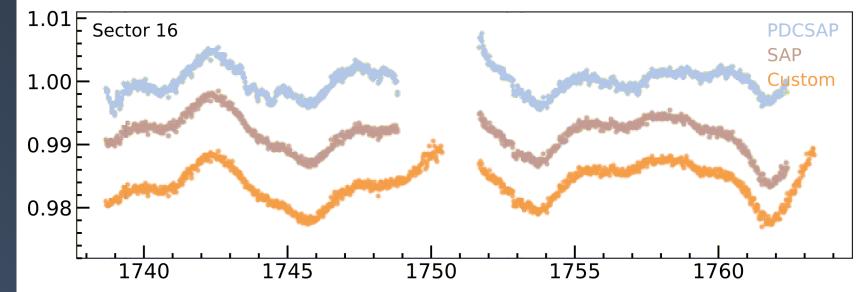


2. Improving our understanding Gaia DR2 + TESS

Identifying planet candidates in open clusters

Gaia: positions and velocities

TESS: rotation periods



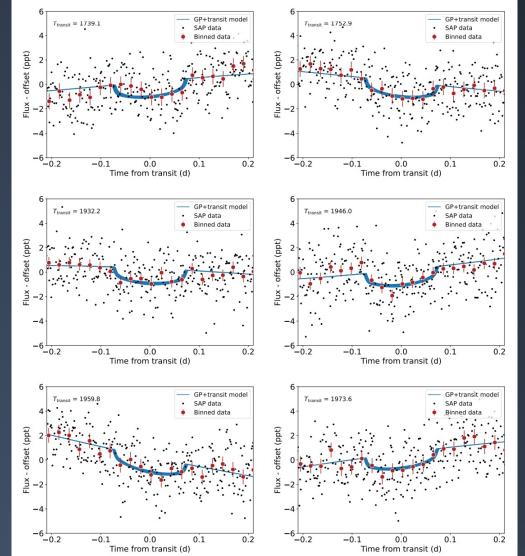
Newton+2021

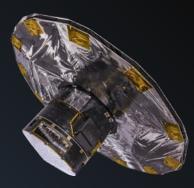
Gaia DR2 + TESS

Identifying planet candidates in open clusters

Gaia: positions and velocities

TESS: rotation periods





Newton+2021

Gaia clusters

Planet Name	Cluster	Age	Planet radius	Orbital period	Provenance
		(Myr)	R_\oplus	d	
AU Mic b	Beta Pictoris	22	4.20	8.463	Plavchan et al. (2020)
AU Mic c	Beta Pictoris	22	2.79	18.859	Gilbert et al. (2022)
DS Tuc A b	Tucana-Horologium	40	5.70	8.138	Newton et al. (2019)
EPIC 211822797 b	Praesepe	700	2.20	21.170	Mann et al. (2017)
HD 110082 b	MELANGE-1	250	3.20	10.183	Tofflemire et al. (2021)
HD 63433 b	Ursa Major	414	2.15	7.108	Mann et al. (2020)
HD 63433 c	Ursa Major	414	2.67	20.545	Mann et al. (2020)
HIP 67522 b	Sco-Cen OB	17	10.07	6.960	Rizzuto et al. (2020)
K2-100 b	Praesepe	700	3.88	1.674	Barragán et al. (2019)
K2-101 b	Praesepe	700	2.00	14.677	Mann et al. (2017)
K2-102 b	Praesepe	700	1.30	9.916	Mann et al. (2017)
K2-104 b	Praesepe	700	1.90	1.974	Mann et al. (2017)
K2-136 b	Hyades	700	0.99	7.975	Mann et al. (2018)
K2-136 c	Hyades	700	2.91	17.307	Mann et al. (2018)
K2-136 d	Hyades	700	1.45	25.575	Mann et al. (2018)
K2-25 b	Hyades	700	3.44	3.485	Stefansson et al. (2020)
K2-264 b	Praesepe	700	2.27	5.840	Rizzuto et al. (2018)
K2-264 c	Praesepe	700	2.77	19.663	Rizzuto et al. (2018)
K2-33 b	Upper Sco	9	5.04	5.425	Mann et al. (2016b)
K2-95 b	Praesepe	700	3.47	10.134	Obermeier et al. (2016)
Kepler-1627 b	Delta Lyra	38	3.82	7.203	Bouma et al. (2022)
TOI-1227 b	Epsilon Cha	11	9.57	27.364	Mann et al. (2022)
TOI-451 b	Pisces-Eridanus	120	1.91	1.859	Newton et al. (2021)
TOI-451 c	Pisces-Eridanus	120	3.10	9.193	Newton et al. (2021)
TOI-451 d	Pisces-Eridanus	120	4.07	16.365	Newton et al. (2021)
TOI-837 b	IC 2602	35	8.63	8.325	Bouma et al. (2020)
V1298 Tau b	Taurus	23	10.27	24.140	David et al. (2019b)
V1298 Tau c	Taurus	23	5.59	8.250	David et al. (2019a)
V1298 Tau d	Taurus	23	6.41	12.403	David et al. (2019a)
V1298 Tau e	Taurus	23	8.74	60.000	David et al. (2019a)

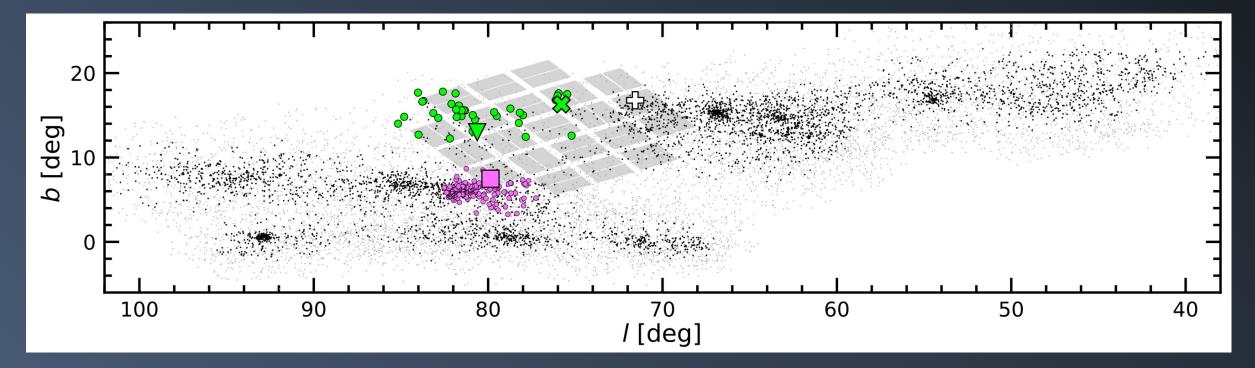


Newton+2021

Gaia clusters



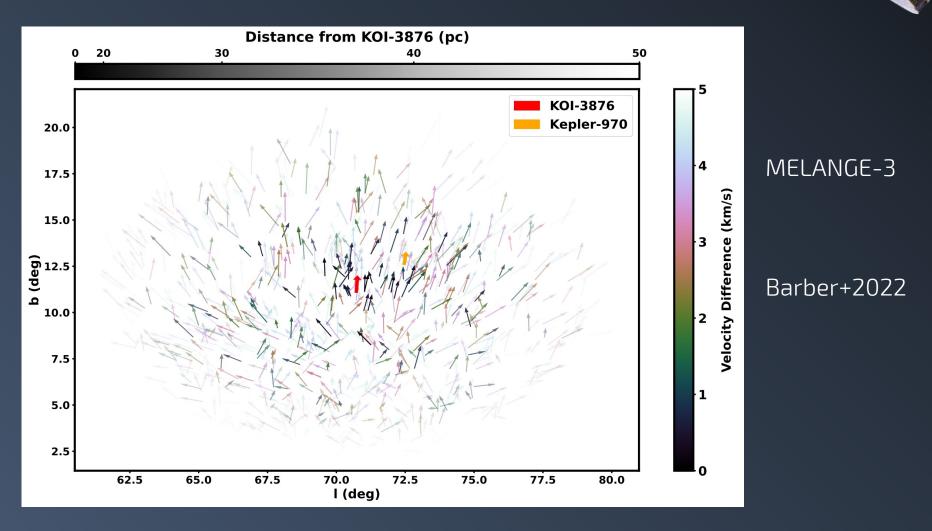
Bouma+2022 arXiv:02205.01112



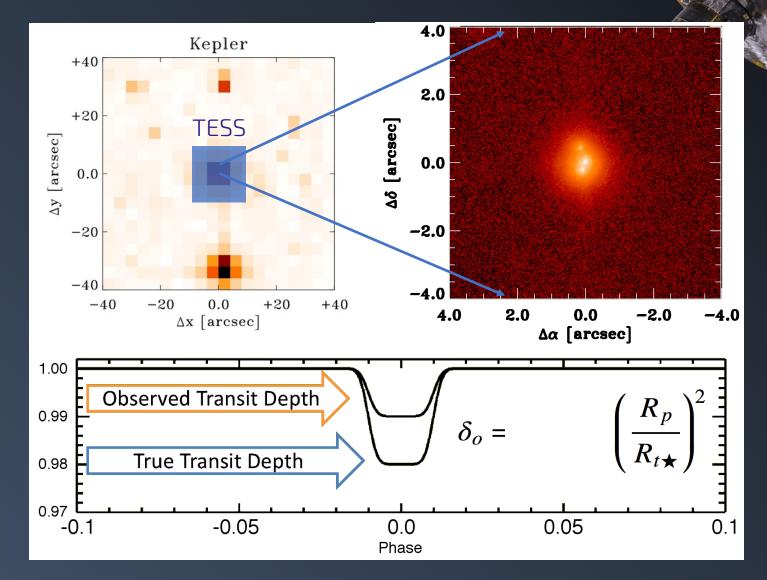
()Identifying EXTRAGALACTIC planets (e.g. Yoshida+2022)

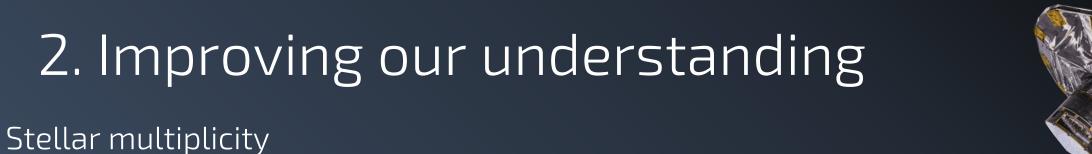


Gaia clusters



Stellar multiplicity





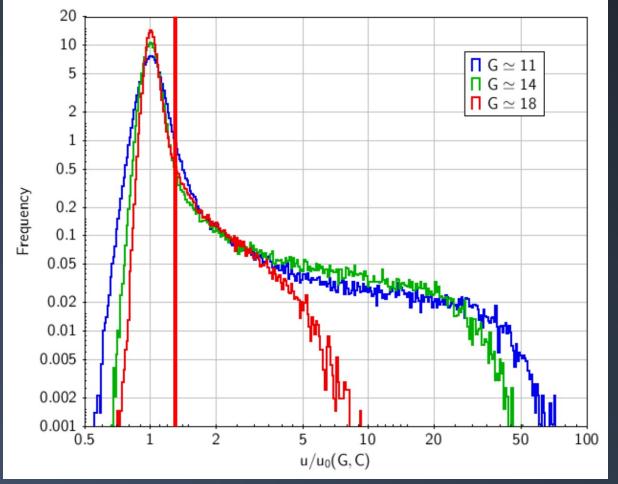
25001500K2/TESS Kepler 2000 1000 1500 1000 500 500 0 0.8 1.6 1.8 0.8 1.0 1.21.4 1.0 1.21.6 1.8 1.4 $\langle X_{R} \rangle = R_{p}(true)/R_{p}(obs.)$ $\langle X_{R} \rangle = R_{p}(true)/R_{p}(obs.)$

Ciardi+2015

Gaia identifies stellar binaries to within ~0.7"

= Astrometric noise Renormalized Unit Weight Error" (RUWE)"

"The median RUWE value for resolved close binaries (<0."75) is 5.56, compared to 1.04 for wider binaries (>0."75) and 1.03 for single targets." (Ziegler+2018)



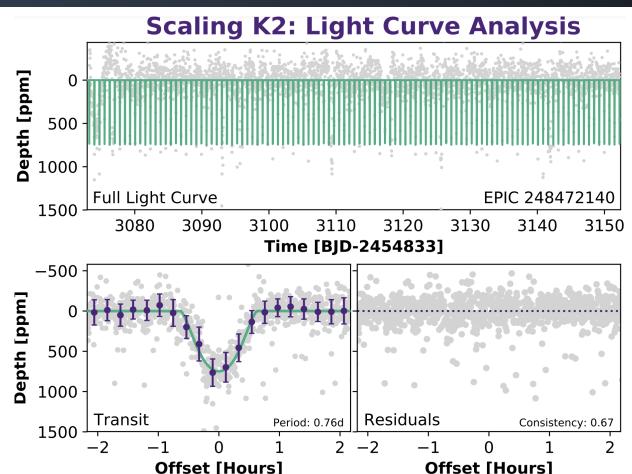


Lindegren+2018

Gaia helps vet planet candidates

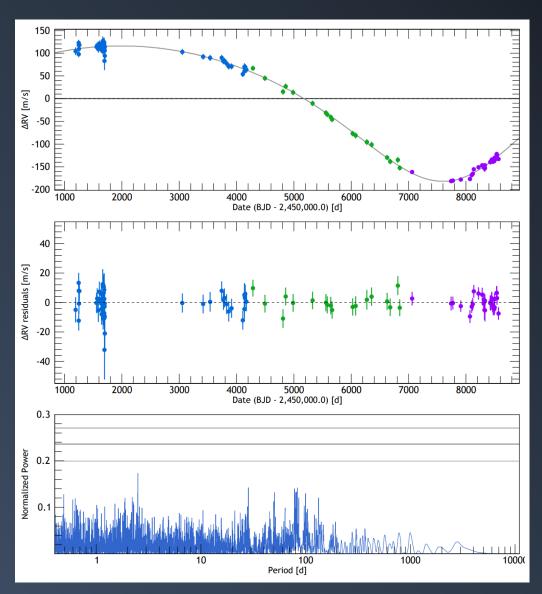
K2 candidate with a RUWE of 4.89

(credit: my spreadsheet of retired candidates and lost dreams)





Gaia helps refute planets



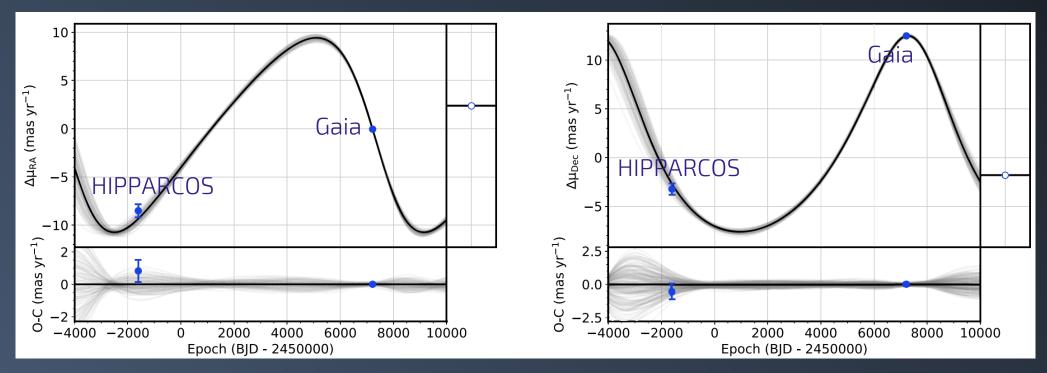
HD 92987 b

Mp sin *i* ~ 17 Jupiter masses (Rickman+2019)





HIPPARCOS-Gaia astrometry



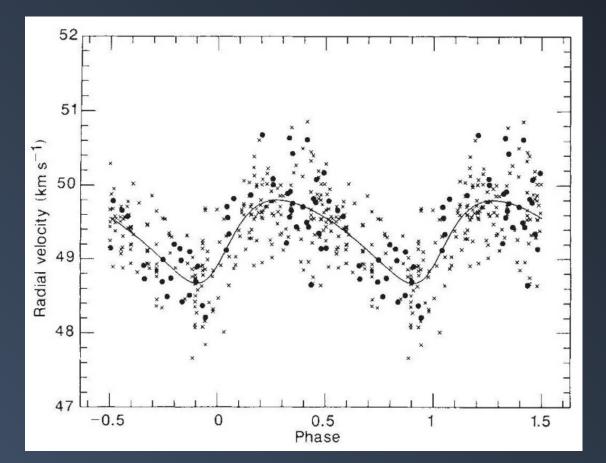
Venner+2021

i= 175 degrees (nearly pole-on)

Mp = 270 Jupiter masses



HIPPARCOS-Gaia astrometry



HD 114762 b The (real) (not) first planet

Mp sin *i* ~ 11 MJup (Latham+1989)

220 MJup (Gaia collaboration, 2022)

Many more improvements to RV planets pending...



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3. Future questions

DR3

• Full astrometric solution (inc. RUWE) for 1.46 billion objects

- Higher fidelity discrimination between single and binary/higher order multiplicity planetary systems
- Deeper exploration of planet demographics as a function of single vs. multiple stellar hosts
- Exploring demographics as a function of Galactic sub-components
- Uniform stellar parameters for 470 million objects, including spectral type
 - Ability to explore Kepler/K2/TESS demographics more finely, along more stellar parameter axes
- 12 elemental abundances for 5.5 (2.5?) million objects
 - Exploring demographics as a function of planetary building blocks on an unprecedented scale

3. Future questions



DR4

- Full astrometric, photometric, and RV solutions for 1.46 billion objects
 - Full orbits (!)
 - Acceleration of future direct imaging mission targets (anything that's accelerating too fast wouldn't be a good target!) (see Tim's talk tomorrow!)
- Exoplanets!
 - Exploring the demographics of solar-system-like Jupiter systems (see Alessandro's talk tomorrow!)